1 Black Hole Spin Networks and Shadows in FTFT

1.1 Black Hole Spin Networks

In loop quantum gravity (LQG), black hole horizons are quantized as spin networks, with area:

$$A_H = 8\pi\gamma l_{\rm Pl}^2 \sum_i \sqrt{j_i(j_i+1)},\tag{1}$$

where $\gamma \approx 0.274$, $l_{\rm Pl} = \sqrt{\hbar G/c^3}$, and j_i are spin labels [1]. The Fonooni Temporal Field Theory (FTFT) modifies this via the scalar field ϕ_T (mass $\sim 6.82 \times 10^{-13} \,\mathrm{eV}$) through the action:

$$S = \int d^4x \sqrt{-g} \left[\frac{1 - 2\kappa\phi_T}{16\pi G} R + \frac{1}{2} (\partial_\mu \phi_T) (\partial^\mu \phi_T) - \frac{1}{2} m_{\phi_T}^2 \phi_T^2 \right].$$
 (2)

The horizon area becomes:

$$A_H^{\rm FTFT} \approx (1 + 2\kappa\phi_T) 8\pi\gamma l_{\rm Pl}^2 \sum_i \sqrt{j_i(j_i+1)},\tag{3}$$

with $\kappa \sim 10^{-10} M_{\rm Pl}^{-1}$, $\phi_T \approx \phi_0 \approx 10^{-10} M_{\rm Pl}$. The ϕ_T dynamics induce spin transitions:

$$H_{\phi_T}\psi_{j_i} \approx \kappa \phi_T \left[\sqrt{j_i(j_i+1)} - \sqrt{(j_i \pm 1)(j_i \pm 1+1)} \right] \psi_{j_i \pm 1}.$$
 (4)

This increases entropy by $(1+2\kappa\phi_0)$, potentially affecting GW echo damping, testable with LIGO O4 [2].

1.2 Black Hole Shadows

The black hole shadow is defined by the photon sphere's critical curve, observable by the Event Horizon Telescope (EHT) [3]. In the FTFT metric:

$$ds^{2} = -\left(1 - \frac{2GM}{r}\right)(1 + 2\kappa\phi_{T})dt^{2} + \left(1 - \frac{2GM}{r}\right)^{-1}(1 + 2\kappa\phi_{T})dr^{2} + r^{2}(1 + 2\kappa\phi_{T})d\Omega^{2}, \quad (5)$$

the photon sphere is at:

$$r_{\rm ph}^{\rm FTFT} \approx \frac{3GM}{c^2} \left(1 - \frac{4\kappa\phi_0}{3} \right), \quad \phi_0 \approx 10^{-10} M_{\rm Pl}. \tag{6}$$

The shadow diameter is:

$$\theta_{\rm shadow}^{\rm FTFT} \approx \frac{6\sqrt{3}GM}{c^2 D} \left(1 - \frac{4\kappa\phi_0}{3}\right).$$
(7)

For M87^{*} ($M \approx 6.5 \times 10^9 M_{\odot}$, $D \approx 16.8$ Mpc), the correction is $\sim 10^{-18} \mu$ as, below EHT's resolution. For Kerr black holes, ϕ_T 's oscillations ($\omega_T \approx 165$ Hz) may induce asymmetric shadow distortions, detectable with EHT upgrades by 2030 [3].

References

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- [2] LIGO Scientific Collaboration and Virgo Collaboration, "GWTC-3: Compact Binary Coalescences Observed by LIGO and Virgo During the Second Part of the Third Observing Run," Phys. Rev. X 11, 021053 (2021).
- [3] Event Horizon Telescope Collaboration, "First M87 Event Horizon Telescope Results. I. The Shadow of the Supermassive Black Hole," Astrophys. J. Lett. 875, L1 (2019).