

### TRACING THE LIVES, DEATHS, & EXPLOSIONS OF MASSIVE STARS

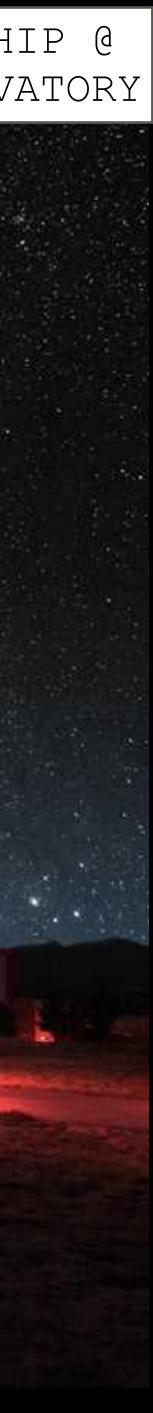
Sarafina Nance PH.D. CANDIDATE UC BERKELEY / LBNL



Royal Astronomical Society of Canada



#### SUMMER INTERNSHIP @ MCDONALD OBSERVATORY



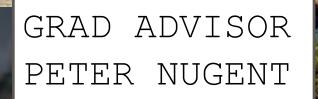










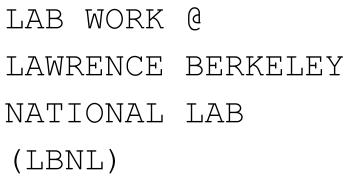


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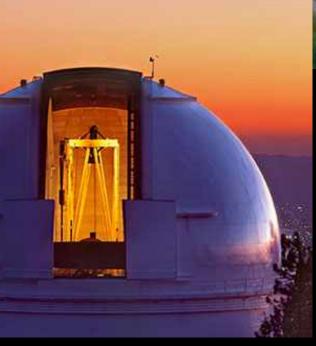




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#### **COMPUTATIONAL COSMOLOGY CENTER**









# RESEARCH INTERESTS: I like stars that explode (aka SUPERNOVAE). I try to understand: 1. WHICH stars explode? 2. HOW do they explode? 3. What do they tell us about the FATE OF THE UNIVERSE?



@starstrickenSF

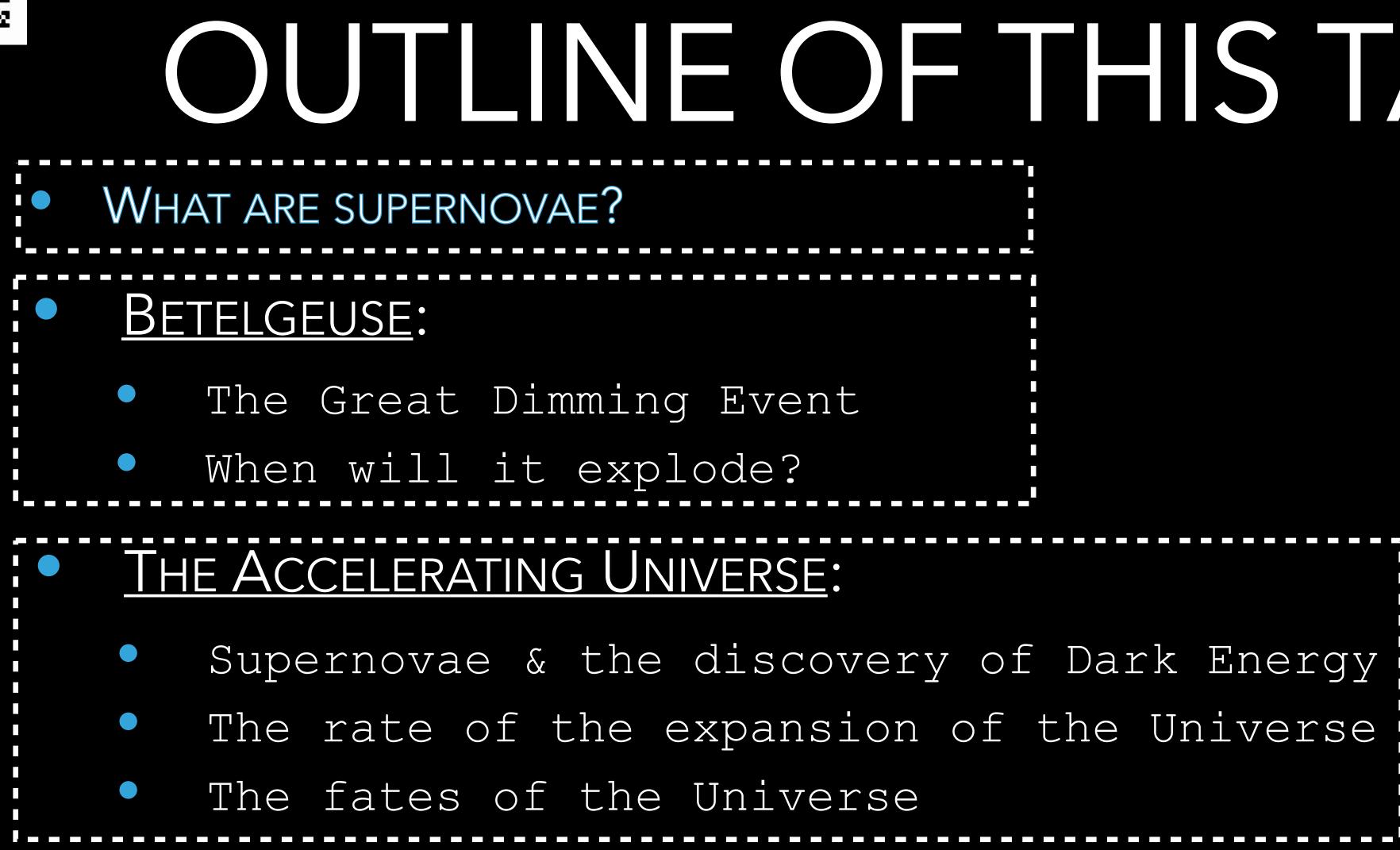


# • WHAT ARE SUPERNOVAE?



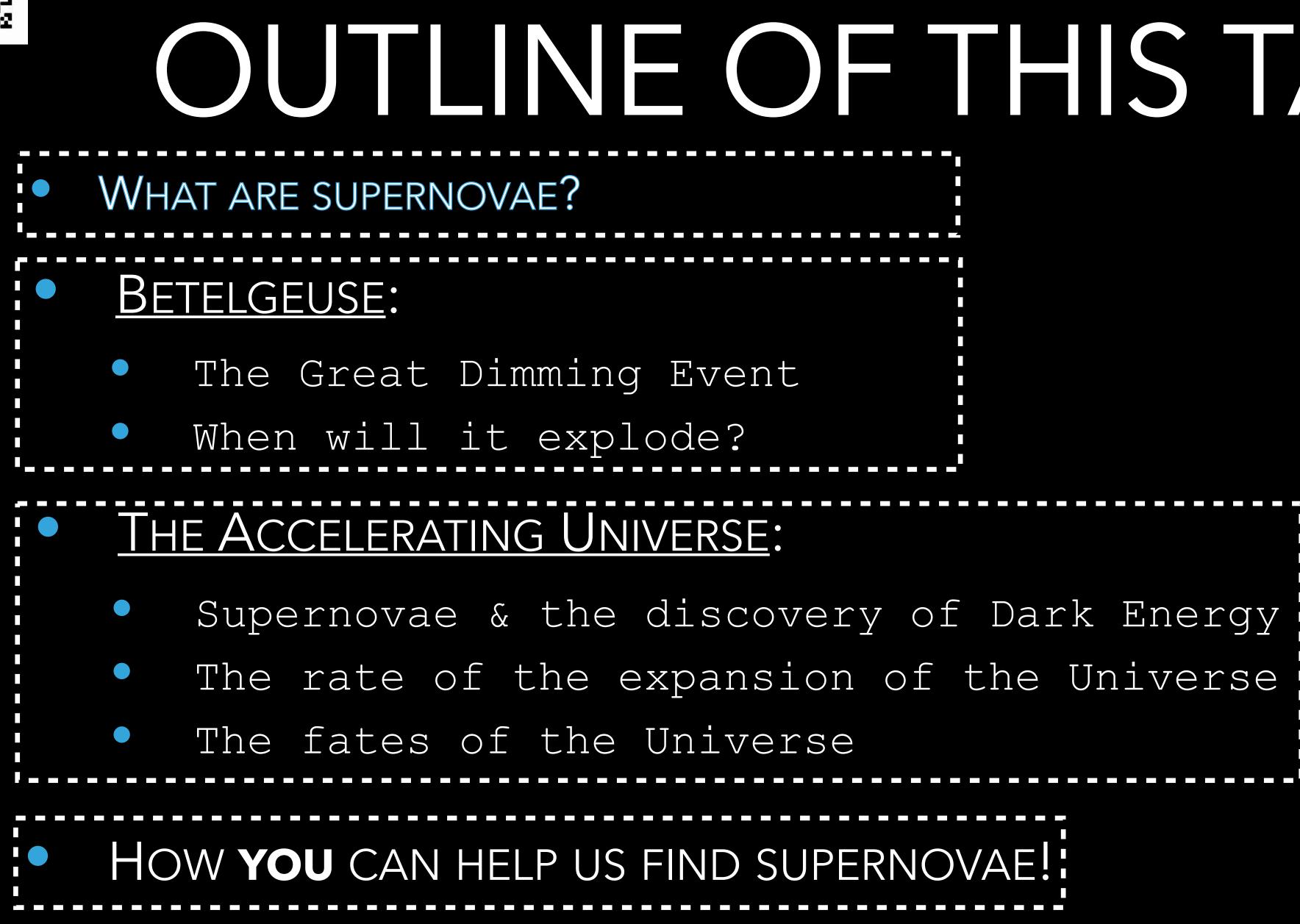
# OUTLINE OF THIS TALK WHAT ARE SUPERNOVAE? <u>BETELGEUSE</u>: The Great Dimming Event When will it explode?





# OUTLINE OF THIS TALK





# OUTLINE OF THIS TALK

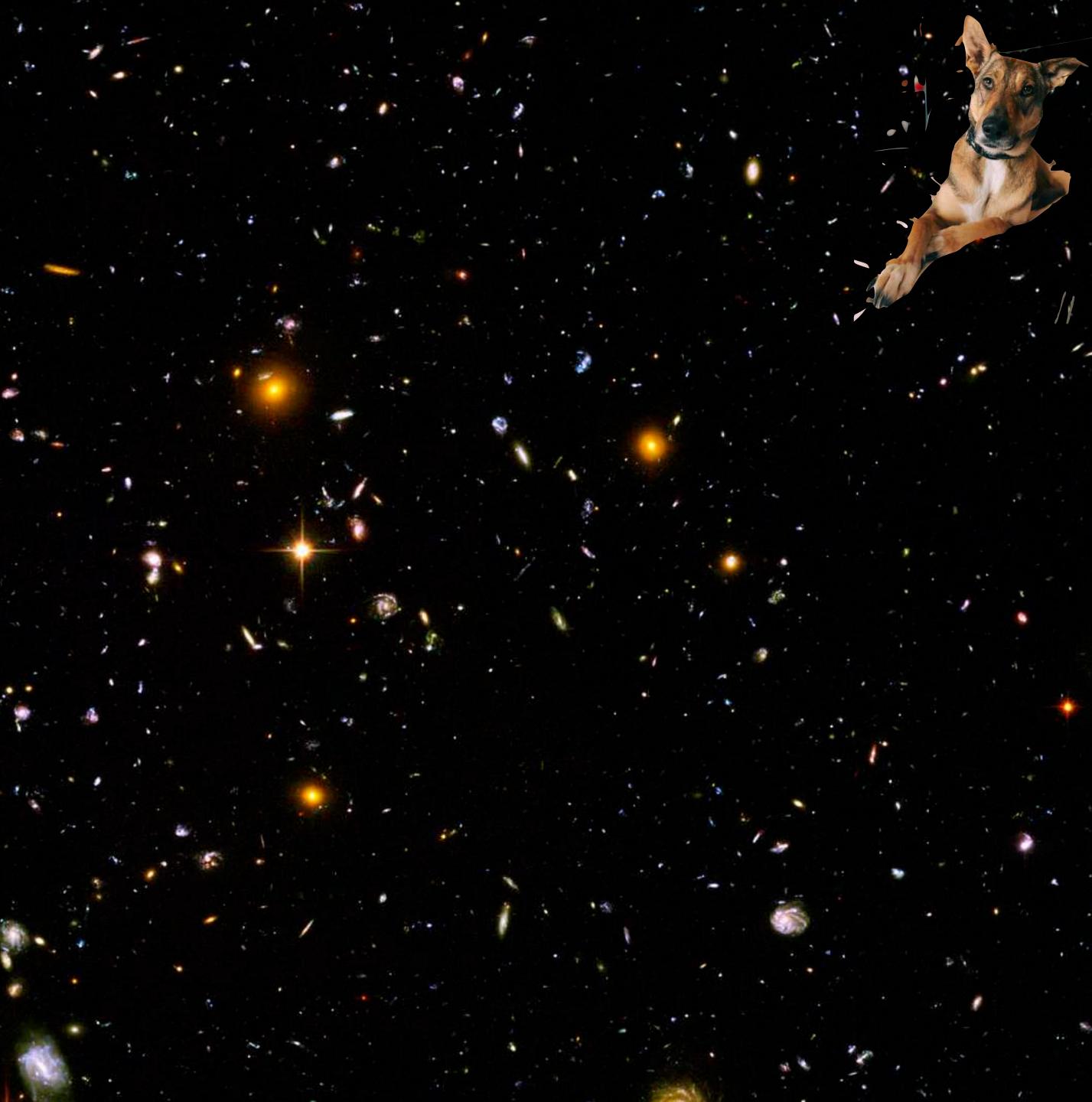


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NASA; ESA; G. Illingworth, D. Magee, and P. Oesch, University of California, Santa Cruz; R. Bouwens, Leiden University; and the HUDF09 Team

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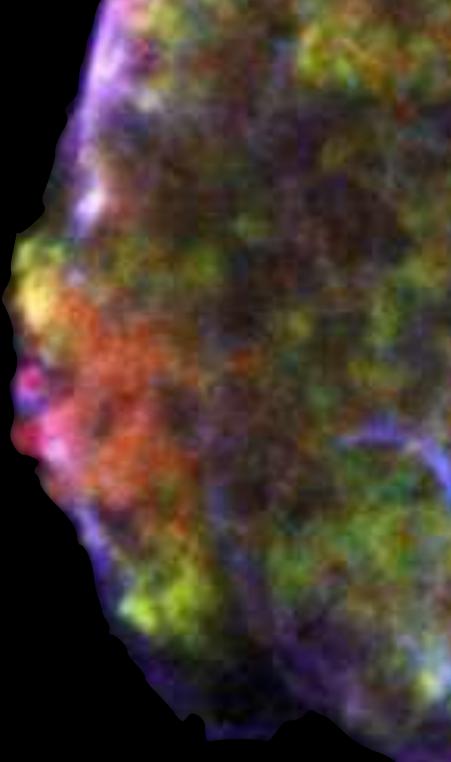
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# OK, SO WHAT EVEN IS A SUPERNOVA? (PL: SUPERNOVAE)



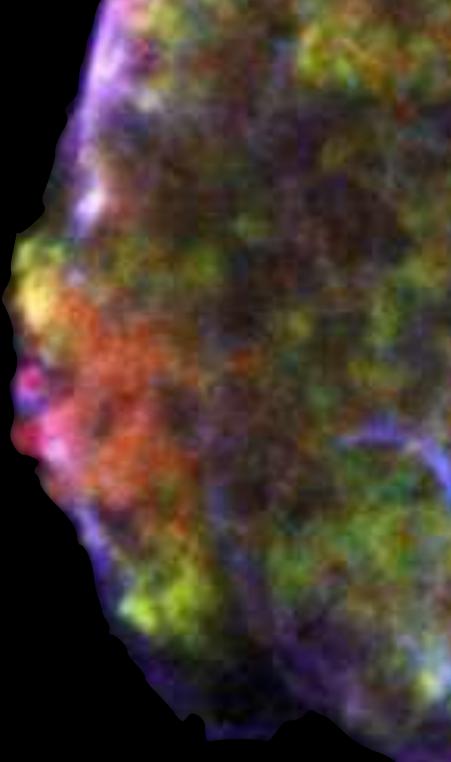




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# OK, SO WHAT EVEN IS A SUPERNOVA? (PL: SUPERNOVAE)

#### **BIGGEST EXPLOSIONS** IN THE UNIVERSE





#### SPEEDS 10-20% SPEED OF LIGHT\* $(ENERGY = 10^{28})$ TONS OF TNT)

# OK, SO WHAT EVEN IS A SUPERNOVA? (PL: SUPERNOVAE)

#### **BIGGEST EXPLOSIONS** IN THE UNIVERSE





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# OK, SO WHAT EVEN IS A SUPERNOVA? (PL: SUPERNOVAE)

#### **BIGGEST EXPLOSIONS** IN THE UNIVERSE

#### OUTSHINE GALAXIES MADE OF HUNDREDS OF BILLIONS OF STARS



#### SPEEDS 10-20% SPEED OF LIGHT\* $(ENERGY = 10^{28})$ TONS OF TNT)



# OK, SO WHAT EVEN IS A SUPERNOVA? (PL: SUPERNOVAE)

#### **BIGGEST EXPLOSIONS** THE UNIVERSE

IN 1 MONTH, EMITS AS MUCH ENERGY AS THE SUN WILL **EVER** EMIT OVER ITS **10 BILLION YEAR LIFETIME** 

#### OUTSHINE GALAXIES MADE OF HUNDREDS OF BILLIONS OF STARS



# Why should we care about supernovae? Why are they cool?



Supernovae <u>trigger star formation</u> & <u>disperse elements</u> throughout the interstellar medium. THUS, THEIR DEATHS TRIGGER THE BIRTHS OF NEW STARS.

FIRE Simulation By: Kareem El-Badry (my cousin!)

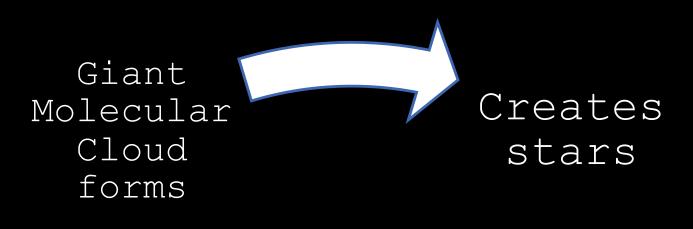




FIRE Simulation By: Kareem El-Badry (my cousin!)



### COSMOLOGICAL z=0.34 SIMULATION OF GALAXY EVOLUTION





Gas cools

Supernovae explode

Gas gets blown out of galaxy

FIRE Simulation By: Kareem El-Badry (my cousin!) Timescale ~3 Gigayears

Temperature color map. Purple dots are cooler

#### 1 kpc





# Supernovae are the birthplace of heavy elements in the Universe.







## Carl Sagan (my hero)



# A HOW COMMON ARE SUPERNOVAE?

1 supernova per second in the Universe

1-3 supernovae seen from Earth per day

1-2 supernovae per galaxy (like the Milky Way) per 100 years

Phil Hopkins, Caltech 2015 FIRE Simulations

#### FIRE Simulation: Flying through a Milky-Way-like Galaxy in the X-Ray



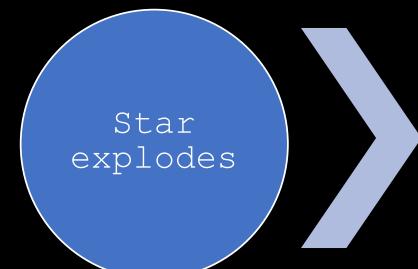


# What do these supernovae actually look like from Earth?



# HISTORICAL SUPERNOVAE

#### Formation of Supernova Remnants



Expelled gas travels outwards & runs into surrounding material

Gas heats up 🔵 surrounding medium & glows

Chanda X-Ray Satellite Telescope

Cast shadows at night was visible during the day for months

Seen by ancient observers in:

Switzerland, Egypt, Iraq, China, Japan, and North America.

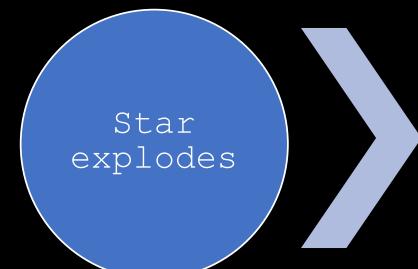
# SN 1006





# HISTORICAL SUPERNOVAE

#### Formation of Supernova Remnants



Expelled gas travels outwards & runs into surrounding material

Gas heats up 🔵 surrounding medium & glows

Hubble Space Telescope

Bright enough to be seen in daylight for

23 days!

Seen by ancient observers in:

China, Japan, Arab countries, and Native American folks

> Visible in the night sky for 653 days

# Crab Nebula / SN 1054



Star = Supernova

Rendition of SN 1054 by Ancestral Puebloans (*Hisatsinom* in Hopi) in Peñasco Blanco - a Chacoan Ancestral Puebloan great house(Chaco Canyon)in New Mexico, USA.

Petroglyph from 1054.

Hand shows relative size of sky objects

Crescent moon shows phase / position of moon relative to the supernova on July 4, 1054

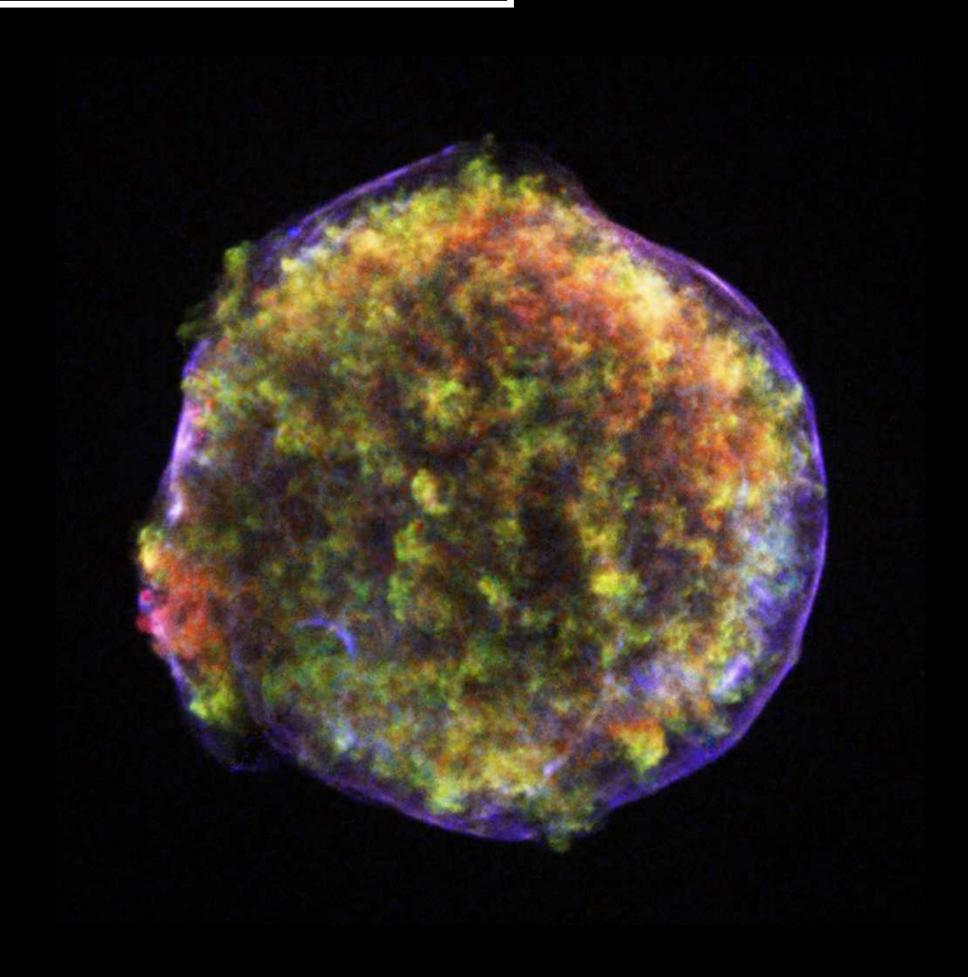




# B Cas/SN 1572Tycho's SN Remnant



Tycho Brahe, discovered SN by eye 1572







# B Cas/SN 1572nt Also seen in...

AL AARAAF,

TAMERLANE,

MINOR POINS.

BY EDGAR A. POD.

BALTIMORN: HATCH & DUNNING.

1829.



Prince of Denmarke. By William Shakefpeare. Newly imprinted and enlarged to almost as much againe as it was, according to the true and perfect



Tycho Brahe, discovered SN by eye 1572

ULYSSES

BY

JAMES JOY

Printed by I.R. for N. L. and are to be fold at his thoppe vnder Saint Dunftons Church in

Chanda X-Ray Satellite Telescope

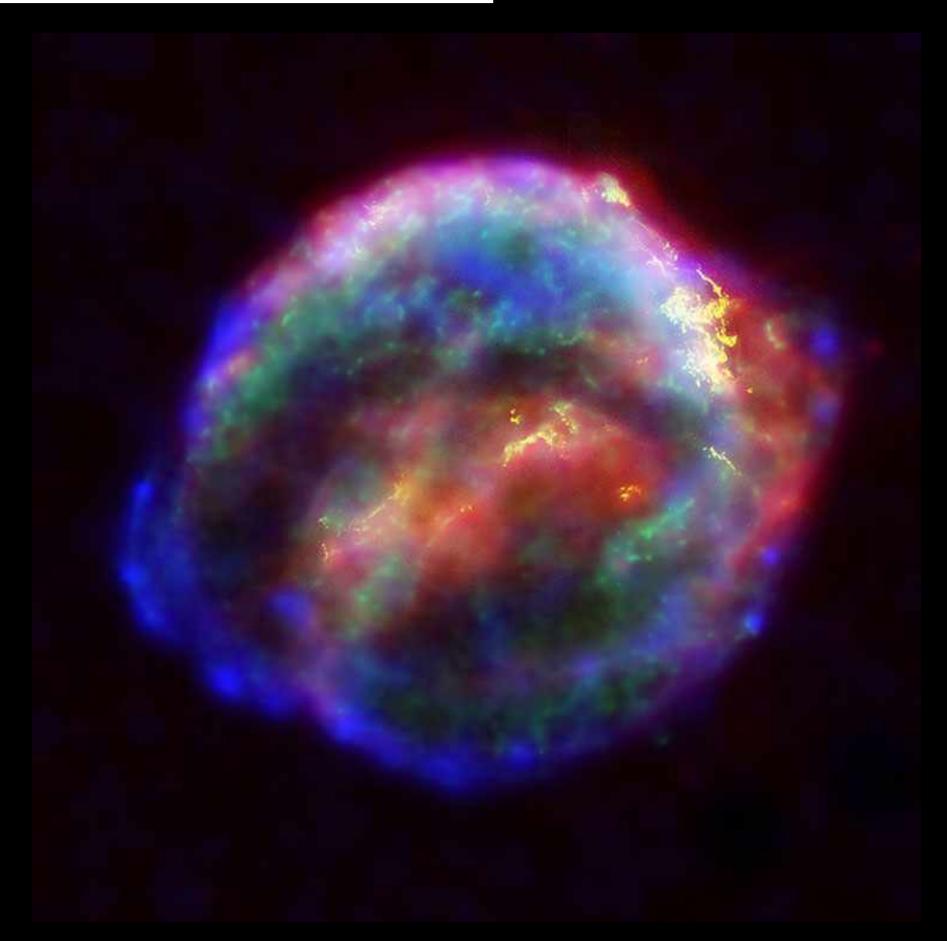






Johannes Kepler (Brahe's GRADUATE STUDENT), discovered SN by eye 1604





CXO/HST/Spitzer Space Telescope



2nd supernovae observed in a generation in the Milky Way. But there haven't been any observed in our galaxy since, over 400 years ago... We are due for <u>another one!</u>

Johannes Kepler (Brahe's GRADUATE STUDENT), discovered SN by eye 1604



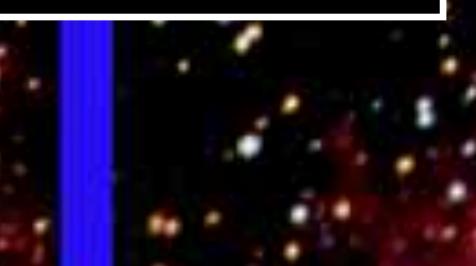
CXO/HST/Spitzer Space Telescope



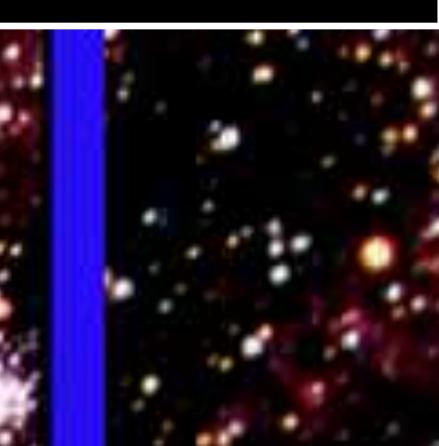
The Large Magellanic Cloud, just 170,000 lightyears away (10<sup>18</sup> miles).

This was the first naked-eye supernova in the era of modern astronomy.

# SN 1987A



On Feb 24, 1987, a star exploded in our cosmic backyard.







# SN 1987A

#### Formation of Supernova Remnants



Here, the ejecta slams into the material released by the progenitor system ~20,000 years before the star exploded.

Hubble Space Telescope, NASA, ESA; Video compilation: Mark McDonald

#### 1994 - 2004

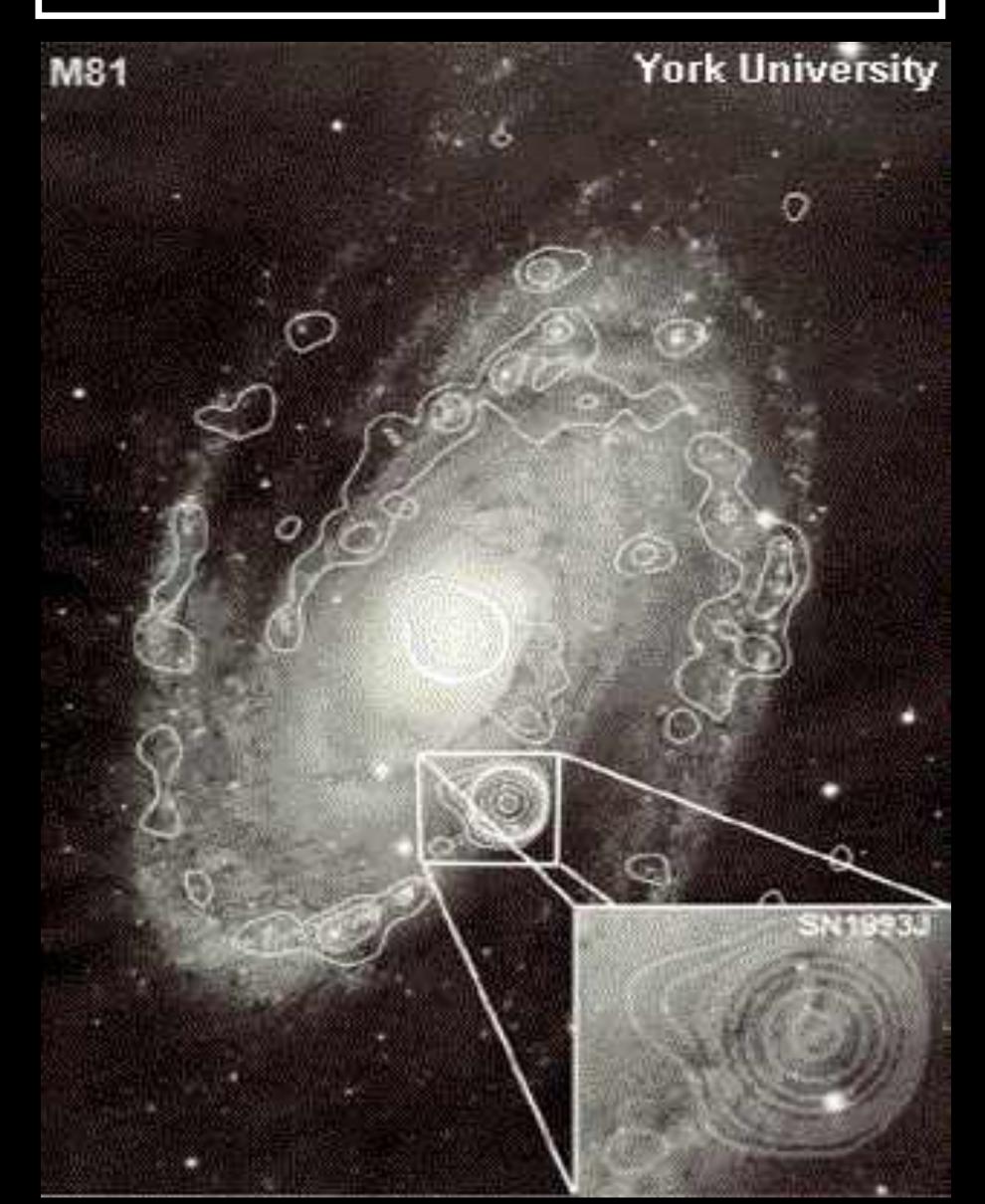


#### 09/1994



# **OTHER** SUPERNOVAE VISBLE FROM

23 consecutive VLBI Images



#### SN 1993J Galaxy: M81

Date: May 1993 - Feb 2000 Distance: 11 Megalightyears Velocity of outflows: 10,000 mi/s -> 6,000 mi/s

Color scale: Brightness of radio emission (blue = faintest, red = brightest)







#### SN 1994D Galaxy: NGC 4526

# OTHER SUPERNOVAE VSBLE FROM EARTH

Supernova as bright as the core of the galaxy (~100 billion stars) Date: May 1993 — Jeb 2000 Distance: 20 Megalightyears (~10<sup>20</sup> miles)





#### SN 1998S Galaxy: NGC 3877

# OTHER SUPERNOVAE VISBLE FROM

Pedro Ré, Marine Biologist at University of Lisbon / Portugal Time: 3.5 months Distance: 42 Megalightyears





# How are supernovae formed?



# Type Ia Supernova





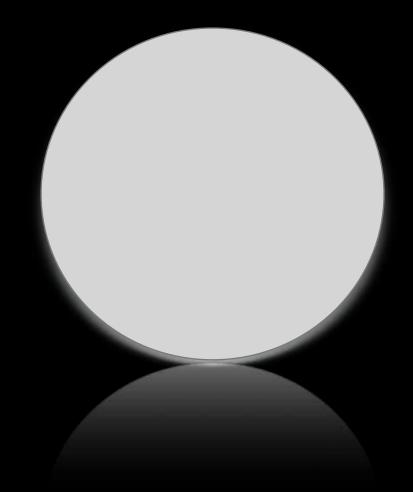
# SUN

# EARTH

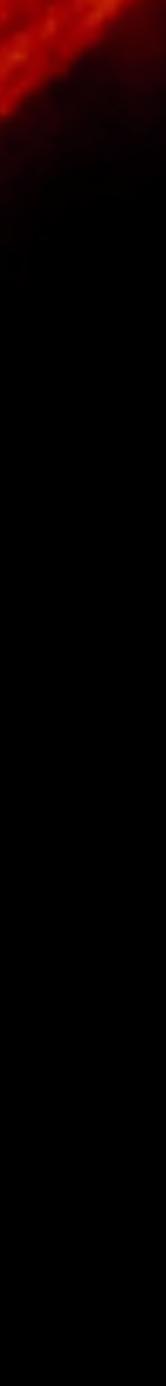
# WHITE DWARF

NASA / SDO







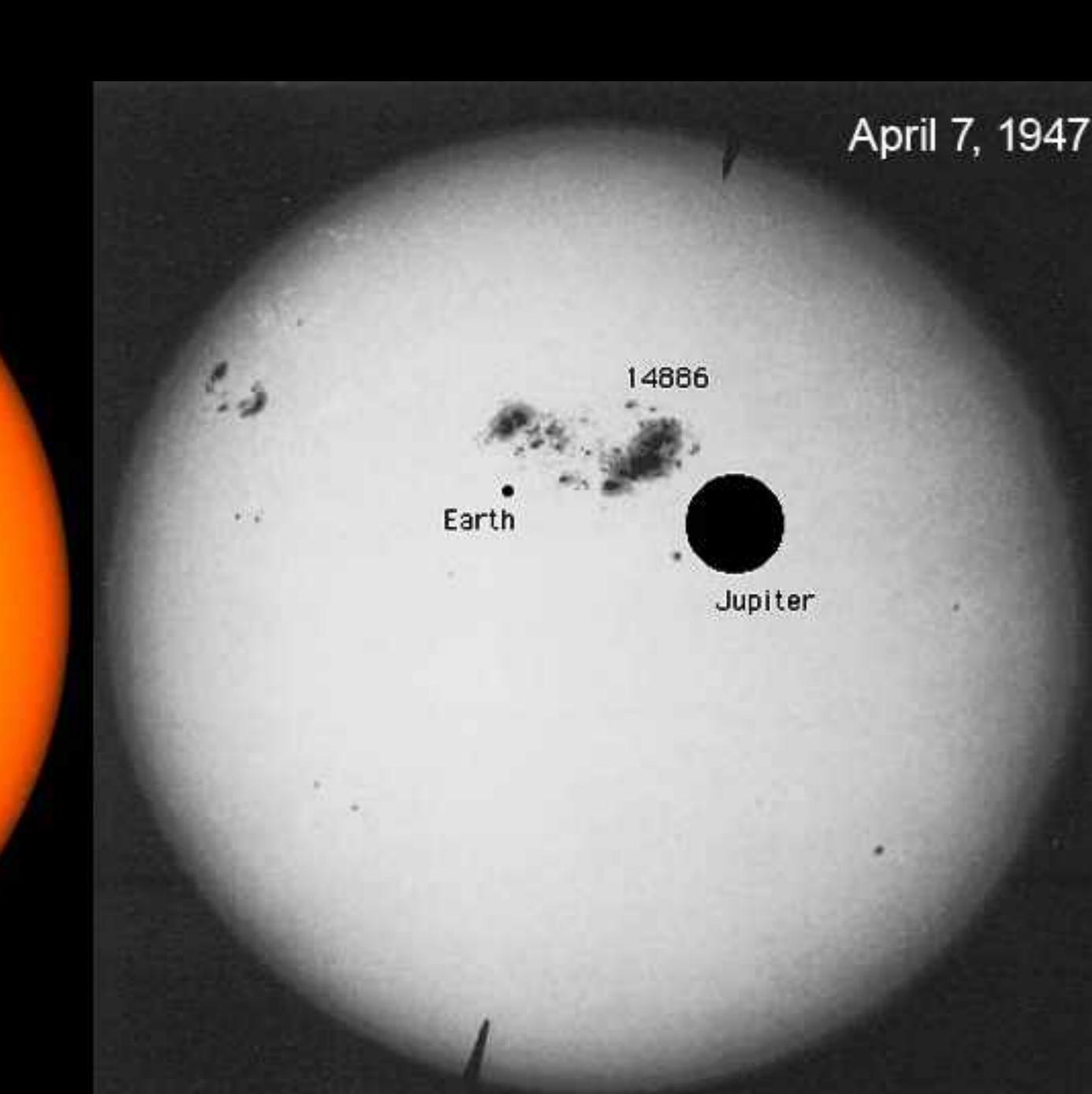




### May 11, 2012



NASA / SDO & the Carnegie Institution





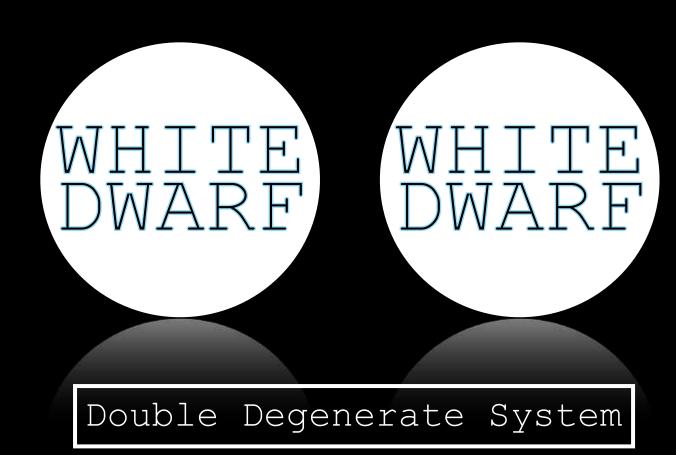
Occurs in a binary system, where the mass transfer triggers a thermonuclear supernova.

### MATN SEQUENCE STAR OR ANT STAR



Single Degenerate System

## Type Ia Supernova





# Core Collapse Supernovae

NASA, ESA, and F. Summers and G. Bacon STScI) Simulation credit: S. Orlando

(INAF-Osservatorio Astronomico di Palermo)

SN 1987A Galaxy: Large Magellanic Cloud

# 1987

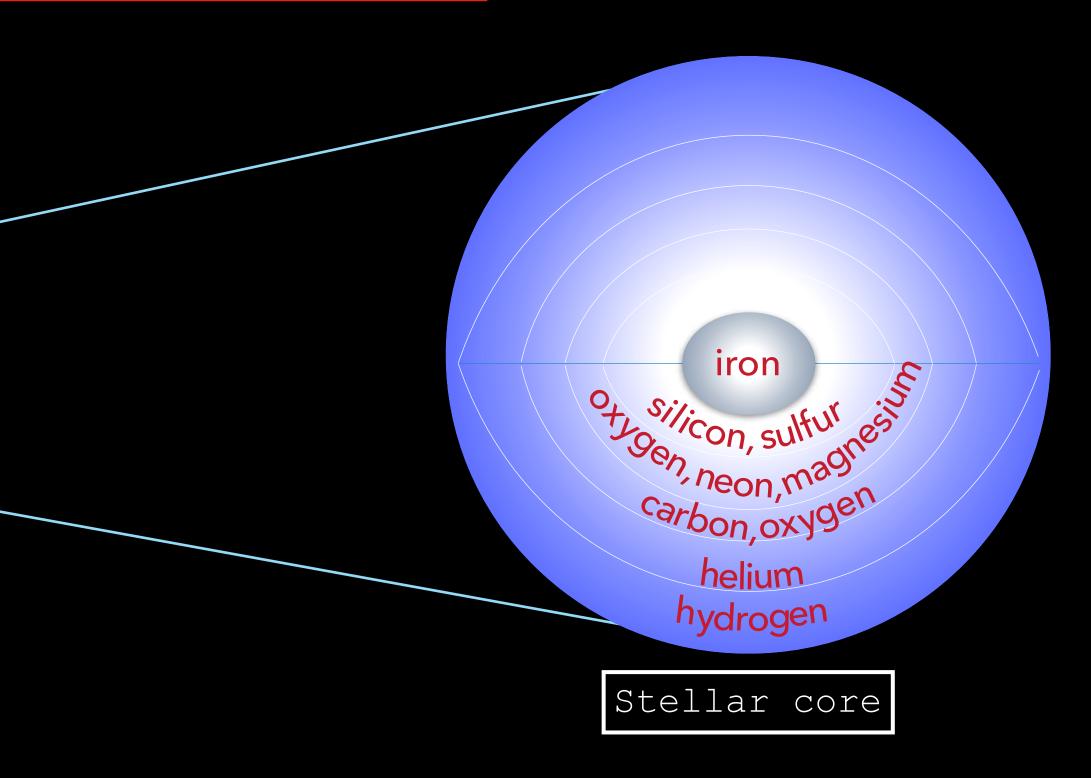


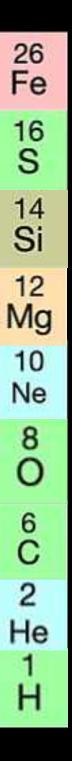


# Core Collapse Supernovae

Occurs at the end of a <u>massive star</u>'s life. The star runs out of fuel, collapses, and subsequently explodes

### CORE OF A RED SUPERGIANT







### $1M_{\odot}=1~\text{SOLAR}~\text{MASS}=\text{THE}~\text{MASS}~\text{OF}~\text{OUR}~\text{SUN}$

# DWARF WHITE $M < 8M_{\odot}$ $M < 25 M_{\odot}$ NEUTRON **STAR** STARS

# SUPERNOVA M > 8Mo

 $M > 25 M_{\odot}$ 

## BLACK HOLE

SINGULARITY







### $1M_{\odot}=1~\text{SOLAR}~\text{MASS}=\text{THE}~\text{MASS}~\text{OF}~\text{OUR}~\text{SUN}$

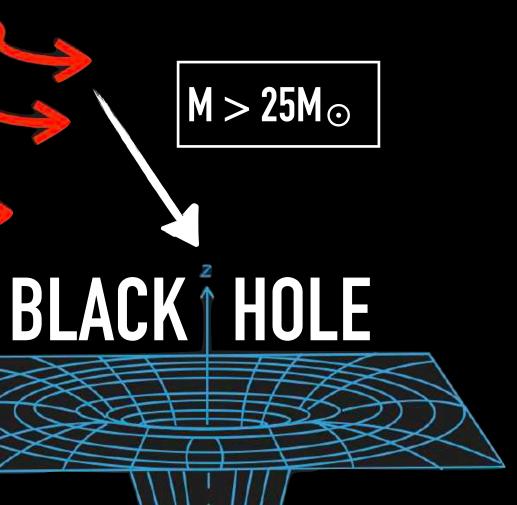
# WHITE $M < 8M_{\odot}$ DWARF

Supernovae are powerful stellar explosions that push the boundaries of our physics knowledge.

# **STARS** © SARAFINA NANCE

# SUPERNOVA

SINGULARITY









MAGNETAR

### **DWARF STARS**

## NEUTRON STAR

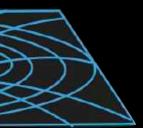
© SARAFINA NANCE

# STELLAR GRAVEYARD

PULSAR

### They form some of the most exotic phenomena in our Universe.







# AND WHEN 2 OF THESE SPEEDY OBJECTS COLLIDE... They create gravitational waves.





### COMBINATIONS:

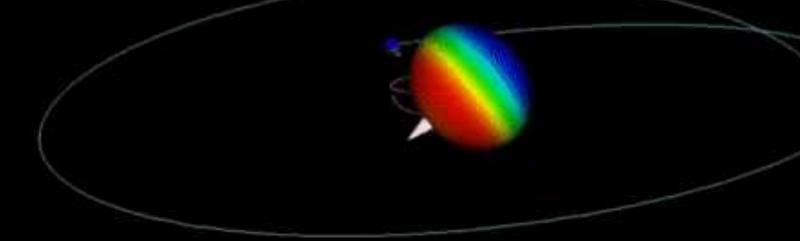
2 BLACK HOLES (BBH)

### NEUTRON STARS (BNS)

NEUTRON STAR +

1 BLACK HOLE

(NSBH)

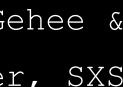




Binary Black Holes (BBH) Spins: 0.91 & 0.3 Mass Ratio: 6:1

by: Robert McGehee & Alex Streicher, SXS







Binary Black Holes (BBH) Masses:  $\sim 30 M_{\odot}$ Time: 1.3 billion years ago

By: Simulating eXtreme Spacetimes Project





# How do we find supernovae?



# SEARCH FOR SUPERNOVAE

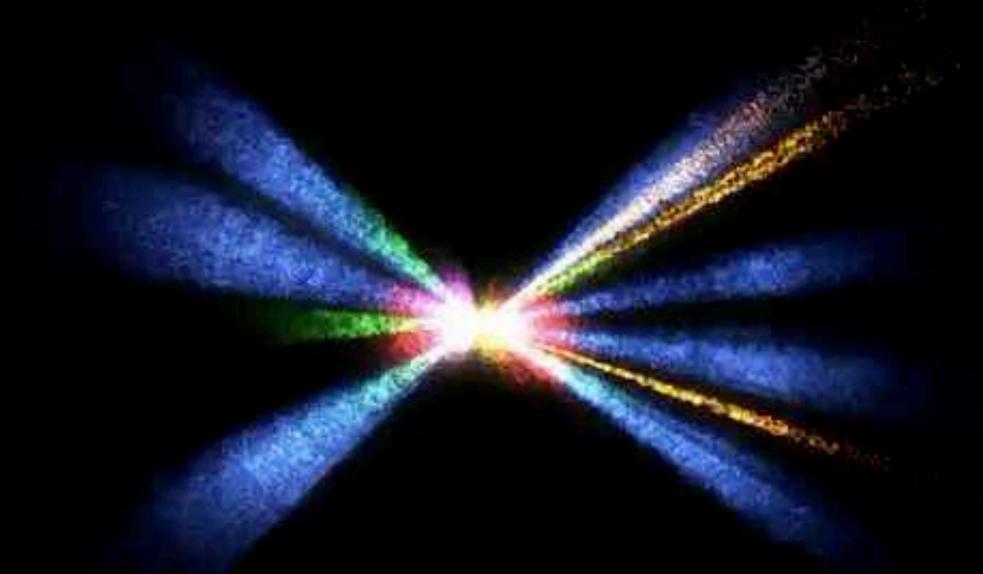
1 supernova per galaxy per 100 years

1 supernova per 100 galaxies per year

100 supernovae per 10,000 galaxies per year (if you monitor frequently enough)

Animation by: Samuel Hinton

### 3-D Visualization of DES+OzDES, WiggleZ, GAMA, 2dF, 6dF, & SDSS Surveys



CENTER: US DOTS: GALAXIES OBSERVED Orange dots: DES + supernovae measured by OzDES. Blue dots: WiggleZ Dark Energy Survey. Green dots: Galaxy And Mass Assembly Survey (GAMA). Red dots: 2dF Galaxy Redshift Survey. Yellow dots: 6dF Galaxy Survey. Purple dots: Sloan Digital Sky Survey (SDSS).



# SEARCH FOR SUPERNOVAE

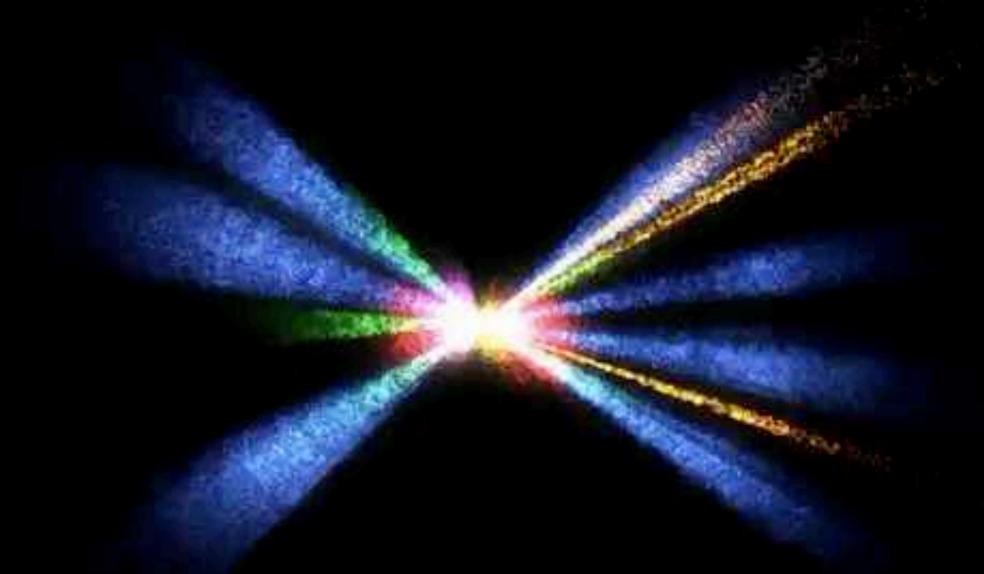
1 supernova per galaxy per 100 years

# Thus, we need to look at a whole bunch of galaxies every single night!!!

100 supernovae per 10,000 galaxies per year (if you monitor frequently enough)

Animation by: Samuel Hinton

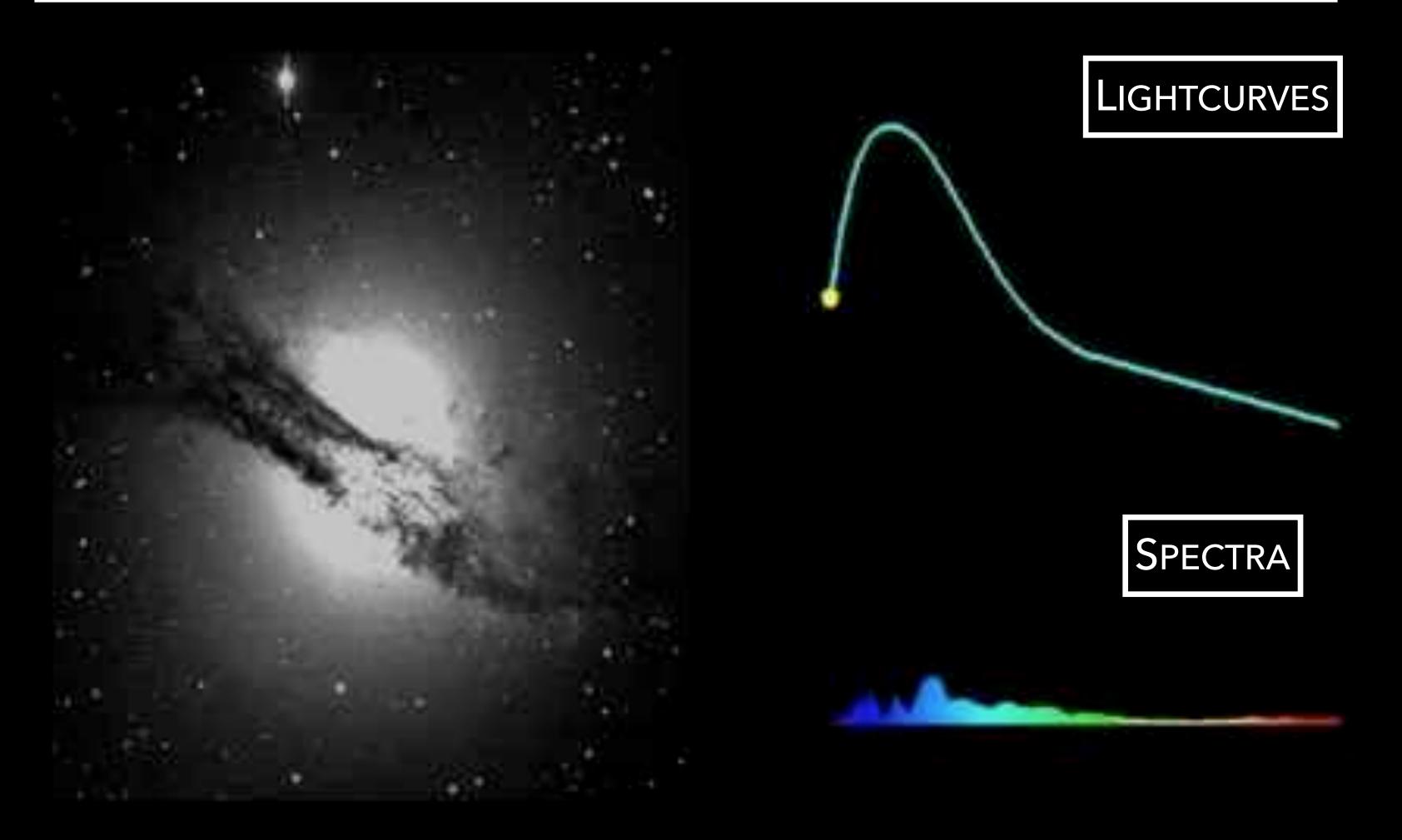
### 3-D Visualization of DES+OzDES, WiggleZ, GAMA, 2dF, 6dF, & SDSS Surveys



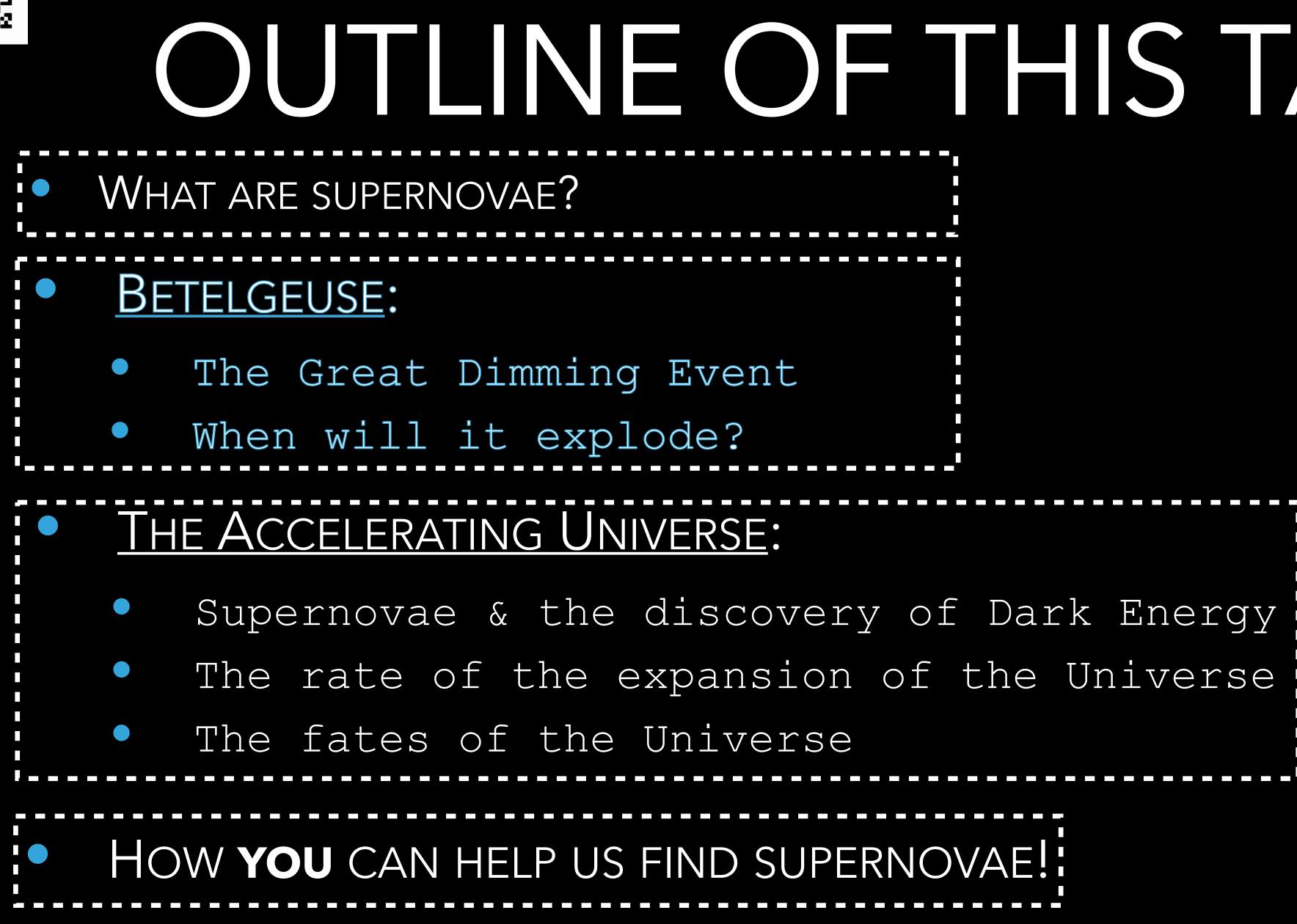
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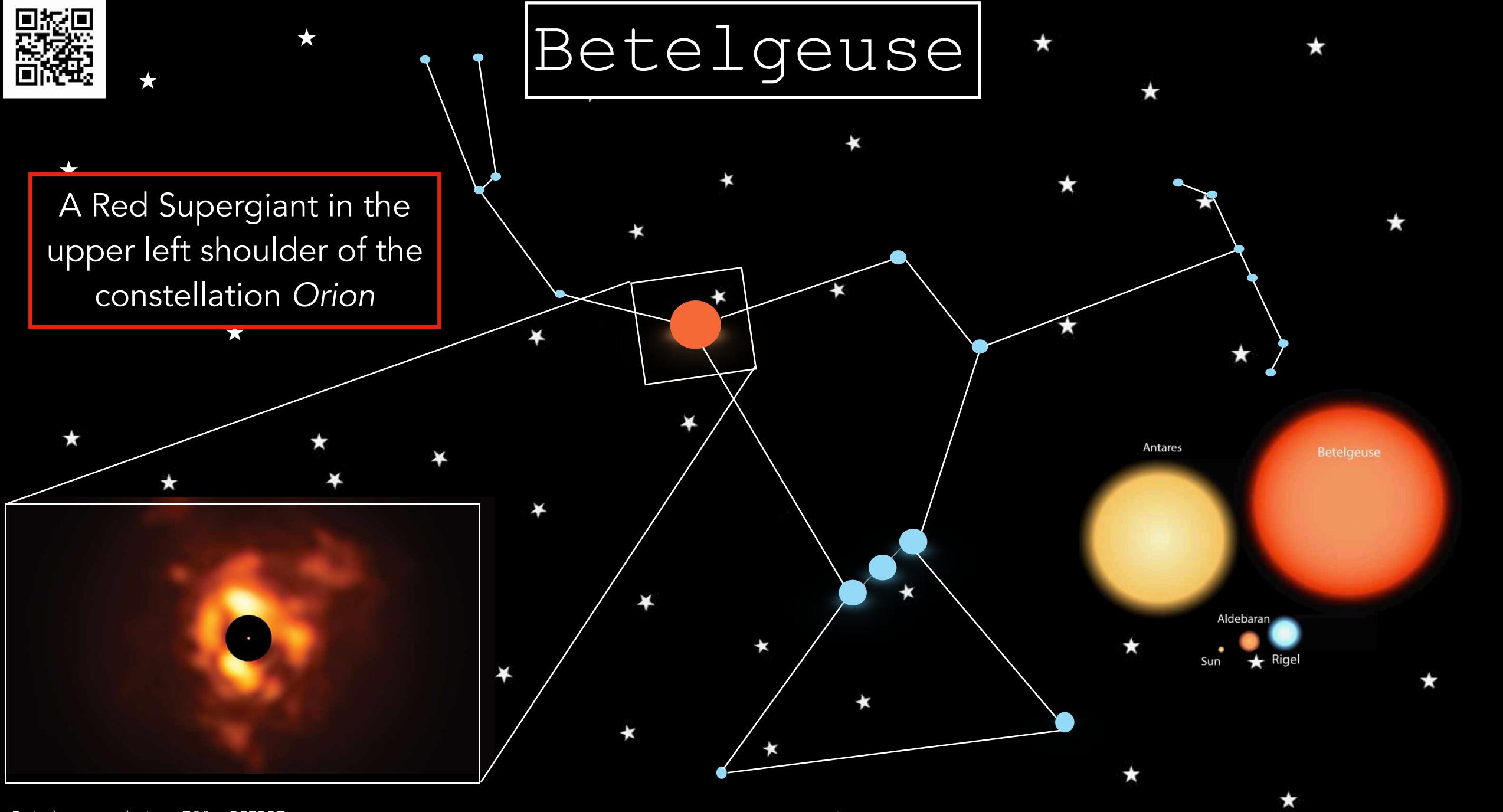
# What kind of data do we get?





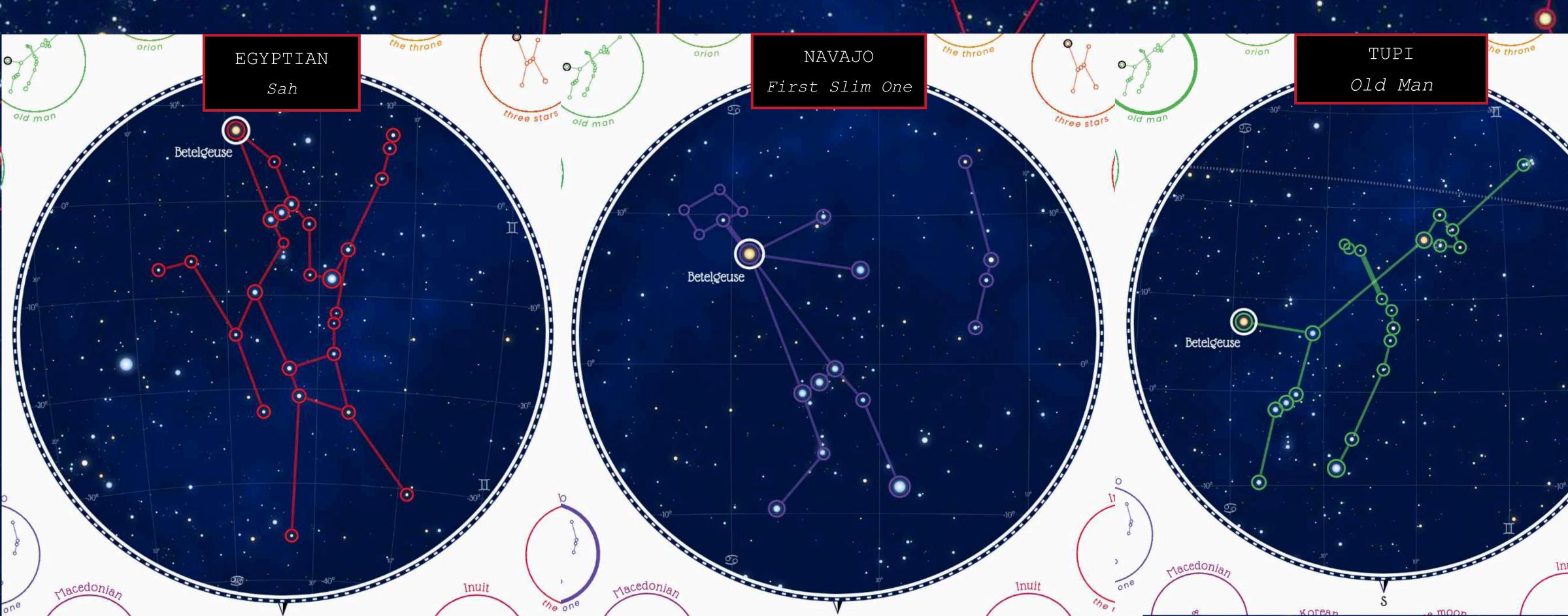


# OUTLINE OF THIS TALK



Betelgeuse photo: ESO, PIERRE KERVELLA, M. MONTARGÈS *ET AL* 



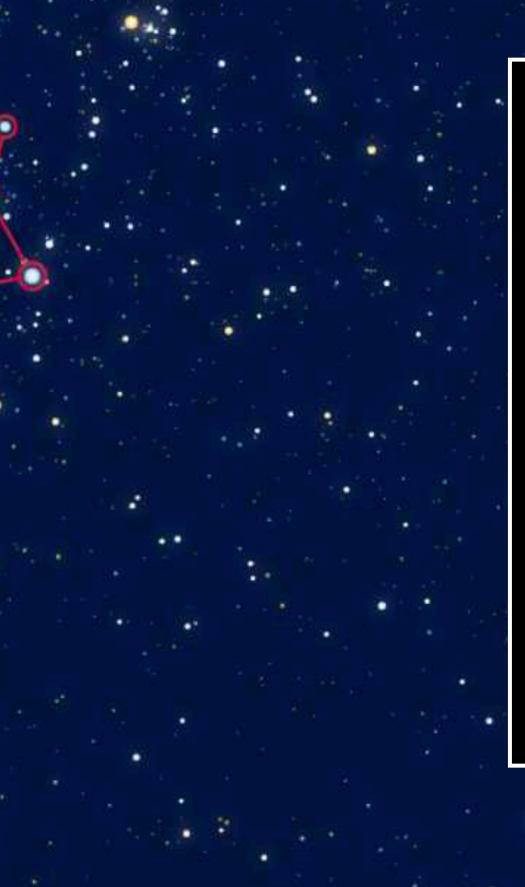


# Betelgeuse: IN CULTURES AROUND THE WORLD

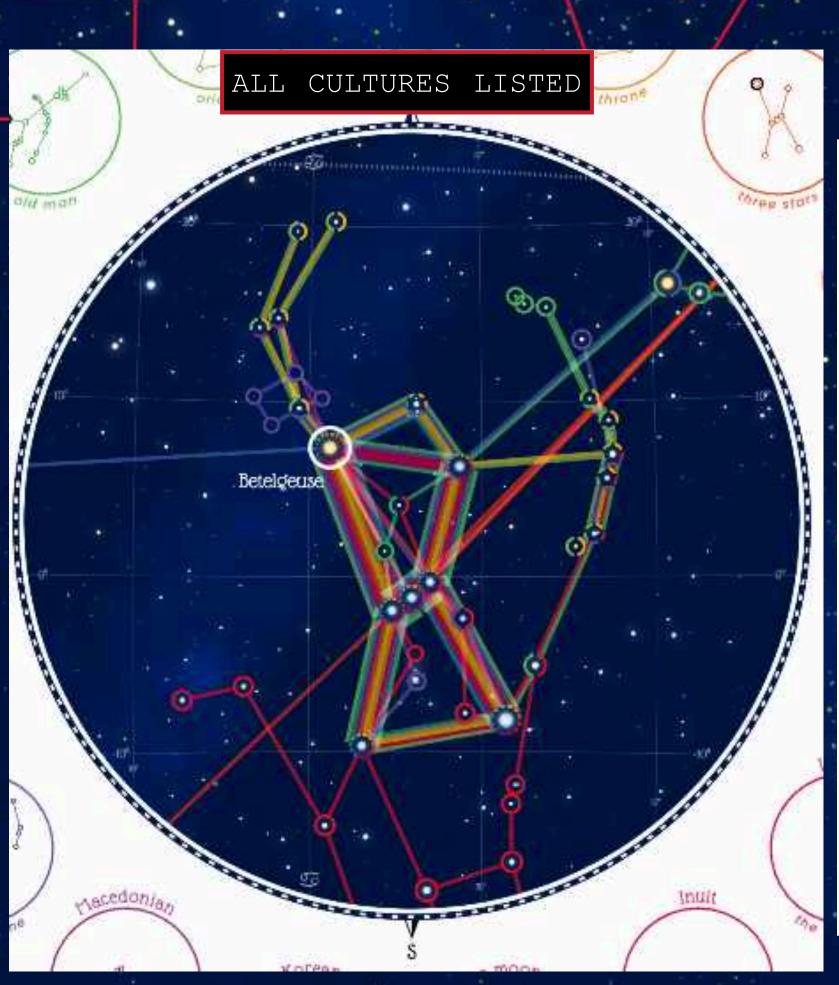
Data viz from astronomer and data visualization lover: Nadieh Bremer











# Betelgeuse: IN CULTURES AROUND THE WORLD

Data viz from astronomer and data visualization lover: Nadieh Bremer

Japanese Moon Stations Korean Macedonian Ojjbwe Sardinian Tukano Tupi Western





# Is Betelgeuse going to explode soon?



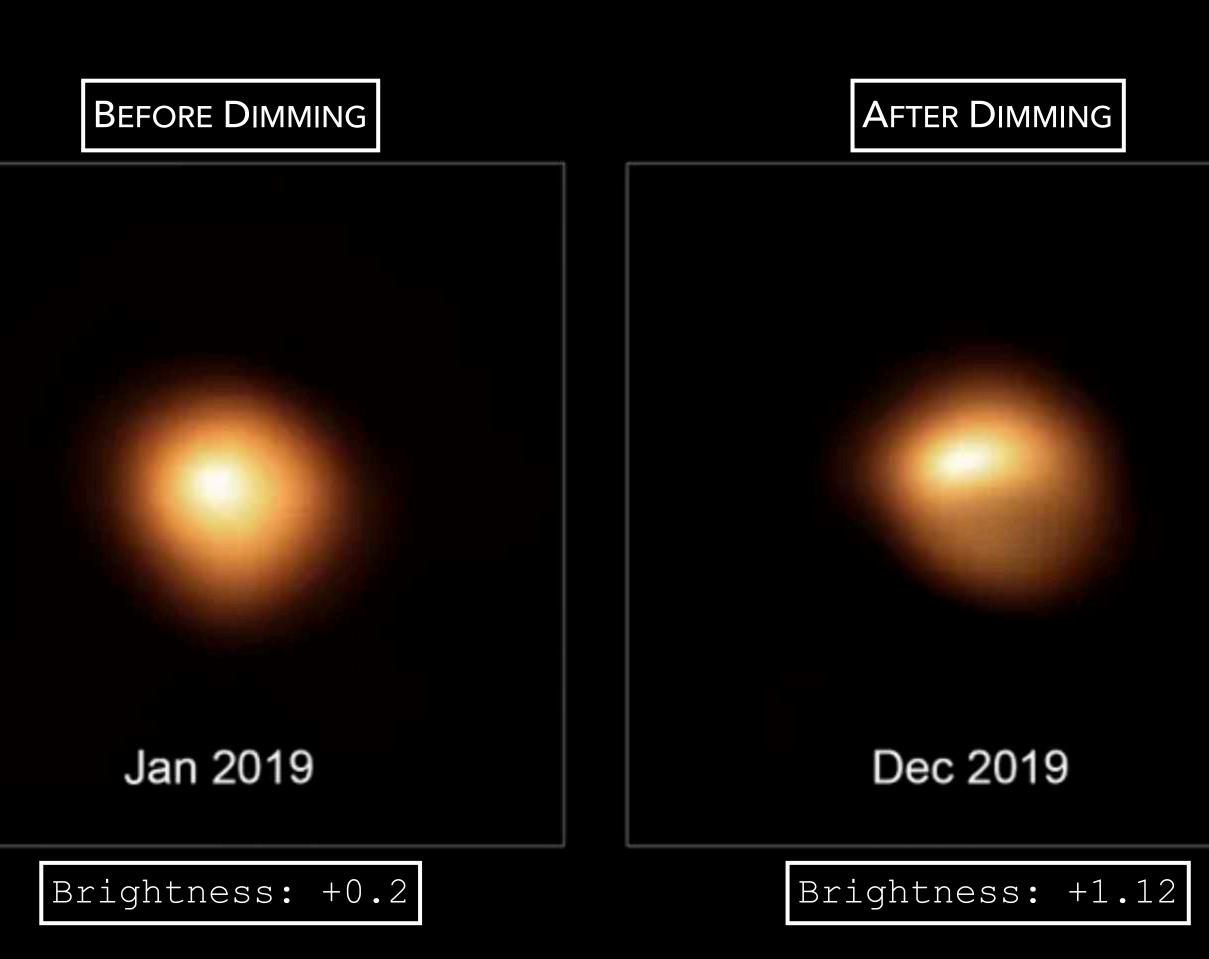
# THE GREAT DIMMING EVENT

From 10/2019 - 2/2020, something really weird happened.

Betelgeuse dimmed. A lot.

DOES THAT MEAN IT IS GOING TO EXPLODE SOON????

SPHERE on ESO's VLT ESO / M. Montarges et al.









By Carlo Inigo Monzon Updated March 10, 2020 15:00 +08

SPHERE on ESO's VLT ESO / M. Montarges et al. © CC BY 4.0 / ESO/L, Calçada / A plume on Betelgeuse artist's imp

### see what happens next," Will Betelgeuse go supernova? Science. uncover mystery of star's dimming brightness

Scientists were able to uncover the reason why the brightness of the Betelgeuse star has



FOR





By Carlo Inigo Monzon Updated March 10, 2020 15:00 +08

SPHERE on ESO's VLT ESO / M. Montarges et al.

### Will Betelgeuse go supernova? Science. uncover mystery of star's dimming brightness

Scientists were able to uncover the reason why the brightness of the Betelgeuse star has



USE GOING TO

L, Calçada / A plume on Betelgeus





By Carlo Inigo Monzon Updated March 10, 2020 15:00 +08

SPHERE on ESO's VLT ESO / M. Montarges et al. Is Betelgeuse Dying? Star Continues to Cet Dimme

## Short answer:

### Will Betelgeuse go supernova? Science. uncover mystery of star's dimming brightness

the present dimming cycle seems over, j

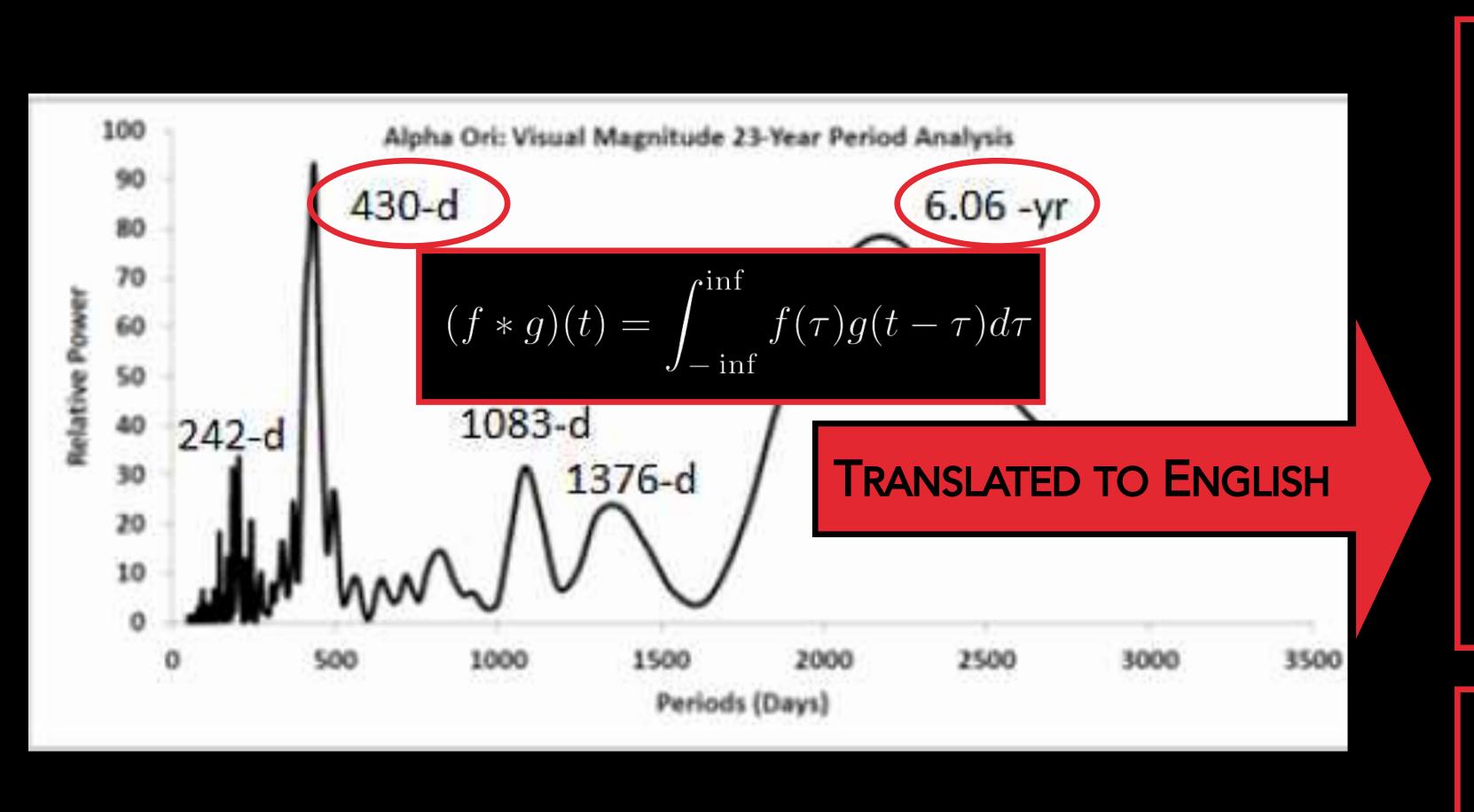
Scientists were able to uncover the reason why the brightness of the Betelgeuse star has



USE GOING TO

L, Calçada / A plume on Betelgeus





## What does the data say?

### Betelgeuse has 2 distinct periods.

When combined, they create a periodic "beating" effect, which predicted the star would reach a minimum brightness between 2/14/20 & 2/28/20.

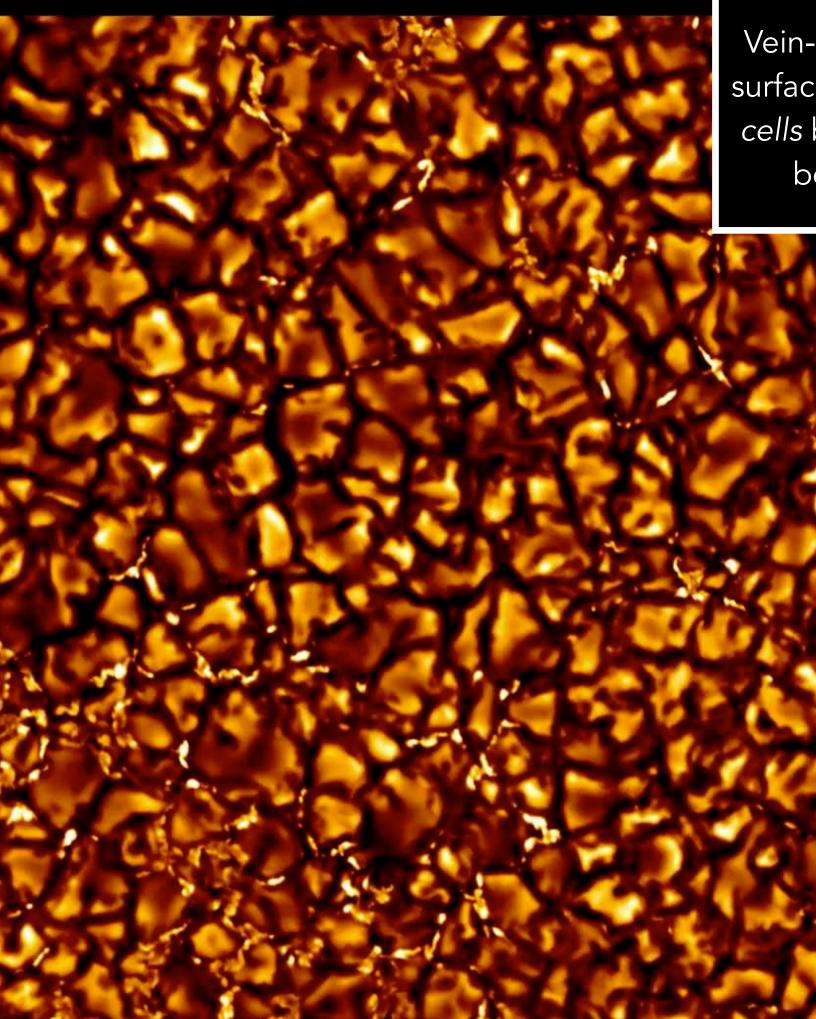
Guess what? It did just that.

After that, it re-brightened... right on schedule!!!









Observer: Vasco Henriques Data reductions : Vasco Henriques / Luc Rouppe van der Voort Date: 25 May 2017 Wavelength: 3950 Å, near Ca II H and K

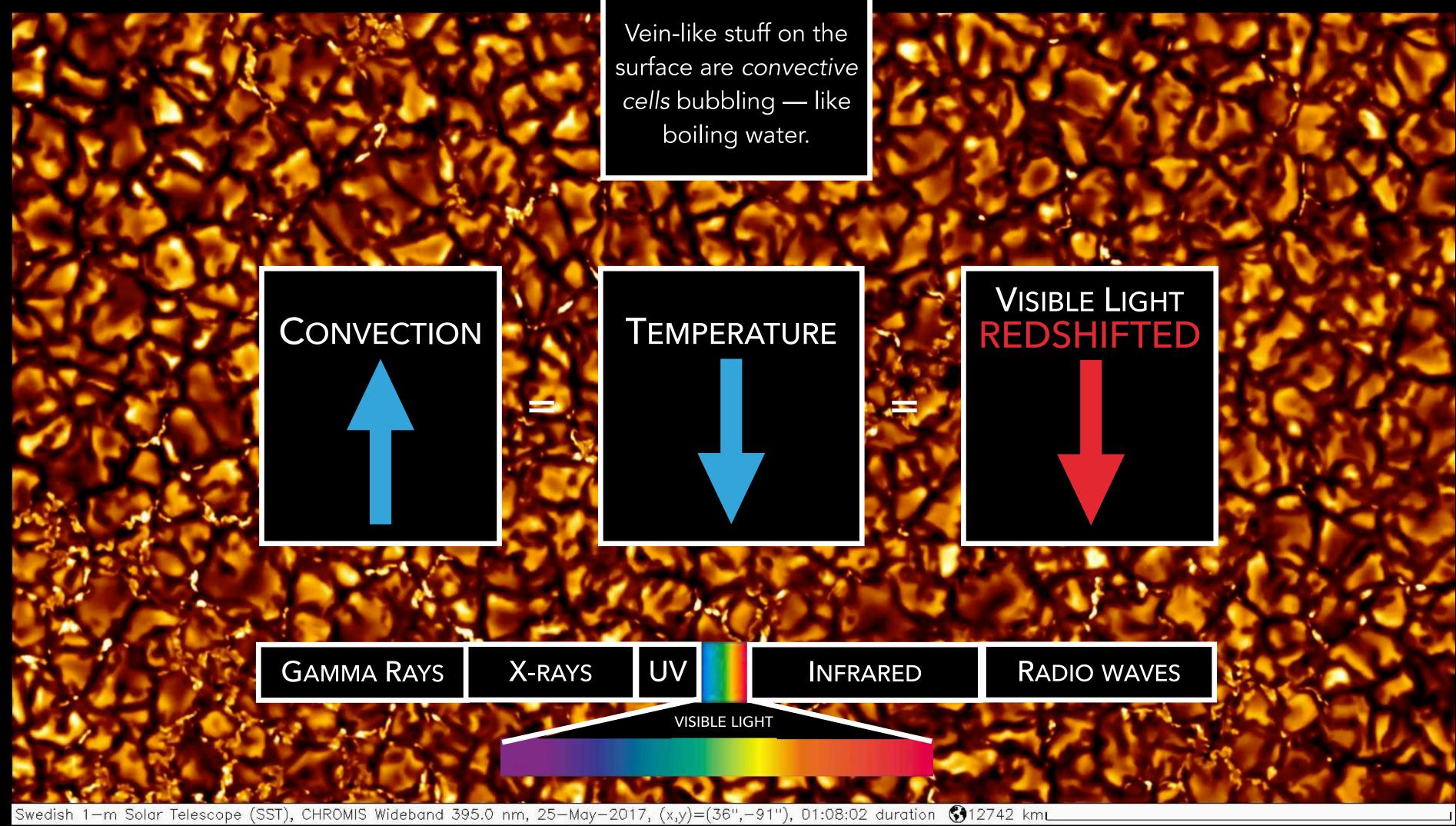
## Theory #1

Vein-like stuff on the surface are convective *cells* bubbling — like boiling water.









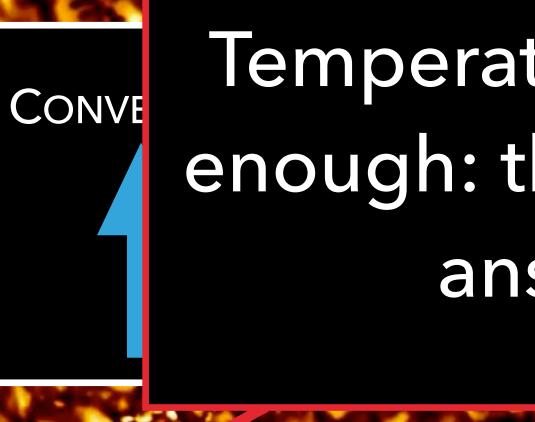
Observer: Vasco Henriques Data reductions : Vasco Henriques / Luc Rouppe van der Voort Date: 25 May 2017 Wavelength: 3950 Å, near Ca II H and K







Vein-like stuff on the surface are *convective cells* bubbling — like boiling water.



🖬 edish 1−m Solar Telescope (SST), CHROMIS Wideband 395.0 nm, 25−May−2017, (x,y)=(36'',−91''), 01:08:02 duration 🚯12742 km∟

X-RAYS

UV

VISIBLE LIGHT

JAMMA RAYS

Observer: Vasco Henriques Data reductions : Vasco Henriques / Luc Rouppe van der Voort Date: 25 May 2017 Wavelength: 3950 Å, near Ca II H and K Temperature not low enough: this is not the answer!

INFRARED

ble Light SHIFTED

RADIO WAVES





Betelgeuse burped out a shell of material that obscured the star, making it appear dimmer.

## Notice the star EXPANDING & SHRINKING, thus growing BRIGHTER & DIMMER







Betelgeuse burped out a shell of material that obscured the star, making it appear dimmer.

We see evidence of "gray" dust around Betelgeuse – this seems to be the right answer!

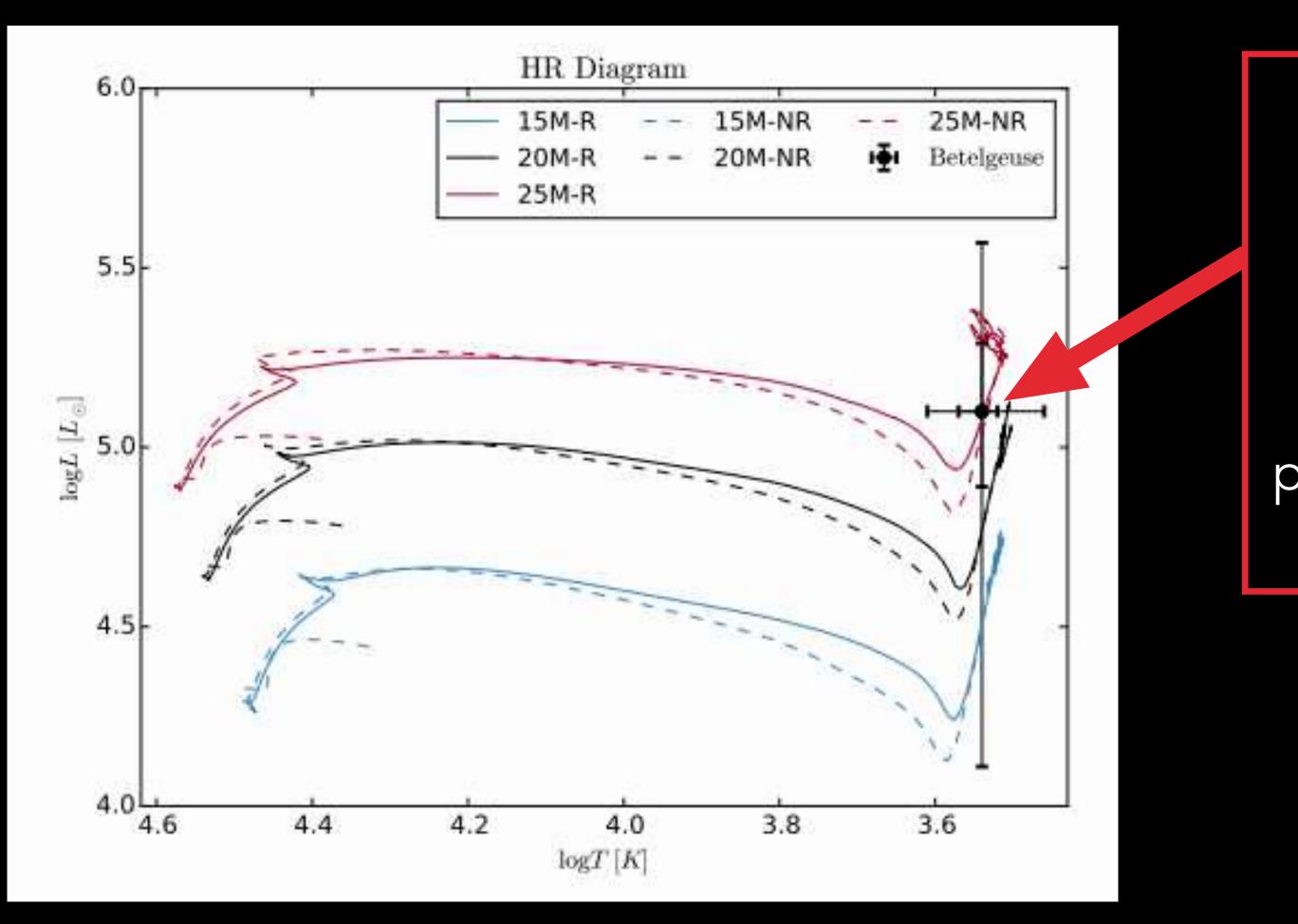
Notice the star EXPANDING & SHRINKING, thus growing BRIGHTER & DIMMER





If Betelgeuse is NOT going to explode soon, then when WILL it explode?



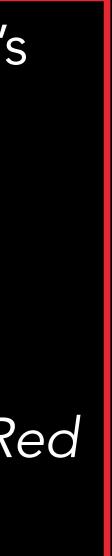


Wheeler, Nance, et al. 2016

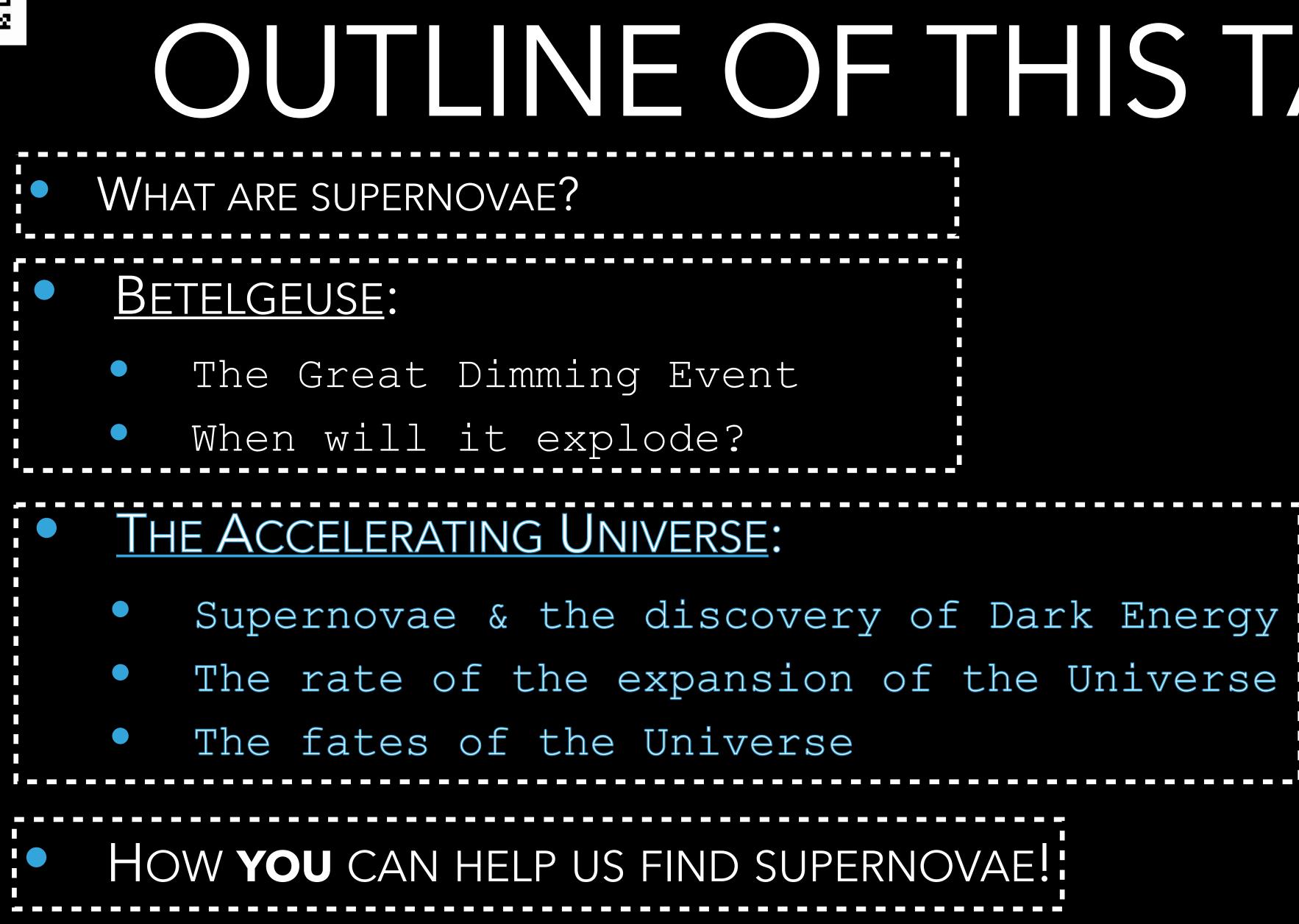
Compare observations with simulations

> Simulations comparing Betelgeuse's observed characteristics with its simulated characteristics place Betelgeuse at the base of the Red Giant Branch

AKA not exploding for another ~100,000 years (or so)







# OUTLINE OF THIS TALK



# Searching the Cosmos with the HUBBLE SPACE TELESCOPE



### Most Distant Supernova Ever

A type of supernova called a Superluminous Supernova

Mat Smith & DES Collaboration

## DES16C2nm

### Thought to have exploded

### **10.5 BILLION YEARS AGO**

(That's when the Universe was just 3.3 billion years old, 1/4 of its current age)

# 15th September 2015





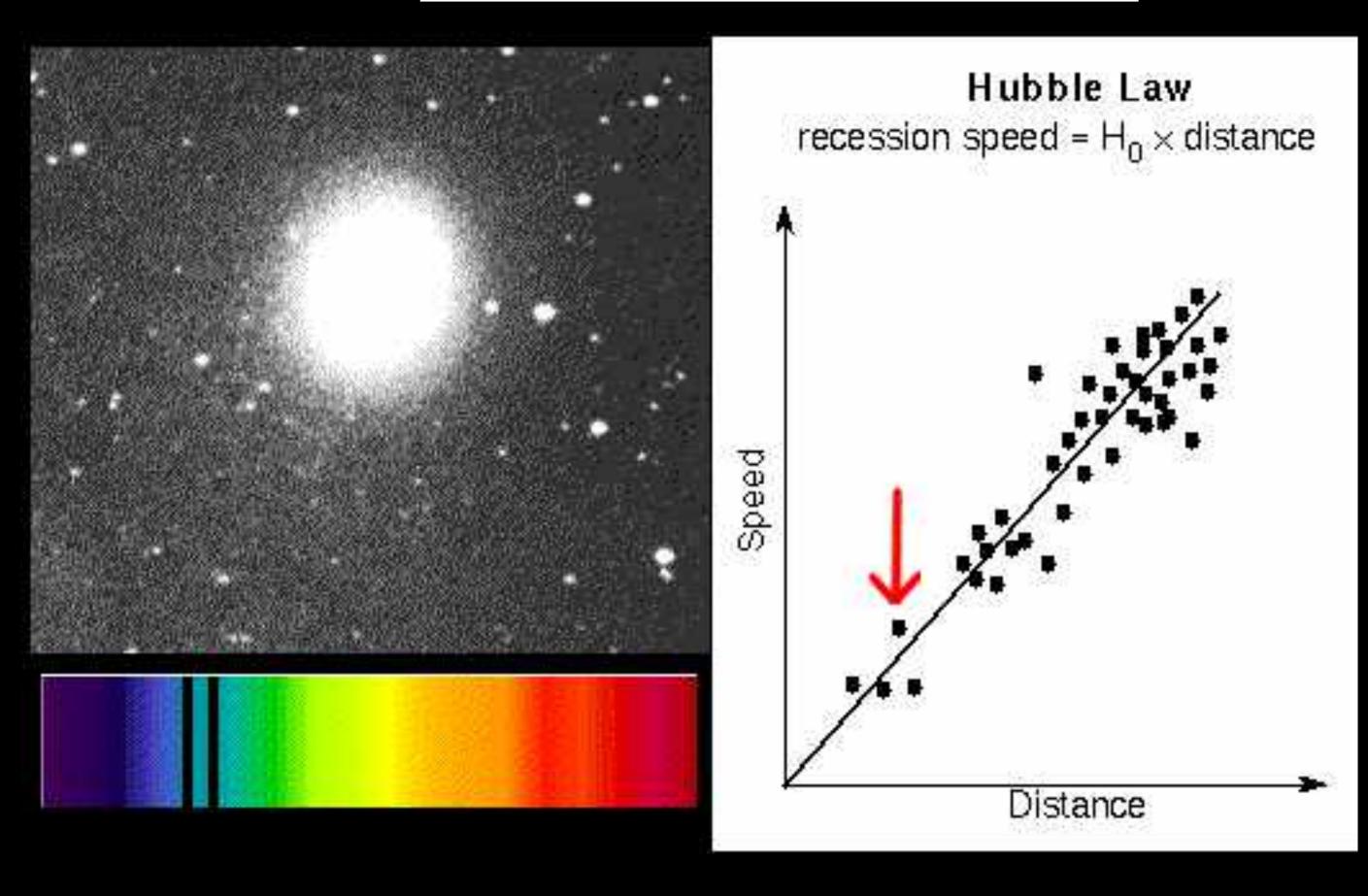
## The Expanding Universe

All galaxies move away from each other.

All galaxies have mass (stars, planets, etc.) which means they gravitationally pull on each other.

All this pulling **should** slow down the Hubble Expansion, forcing distant galaxies to slow down.

### Visualization of Hubble's Law (1929)



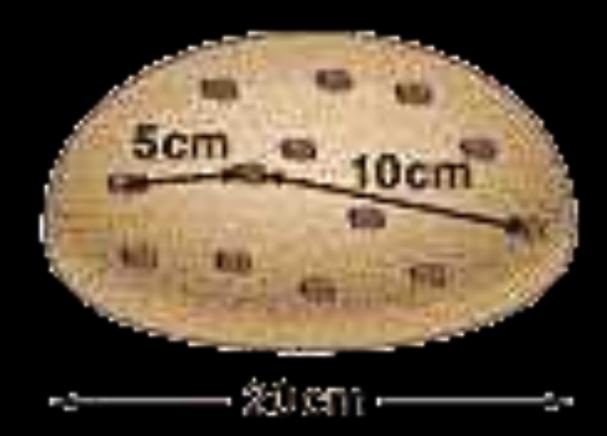


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### Way better visualization of Hubble's Law (1929) that includes bread





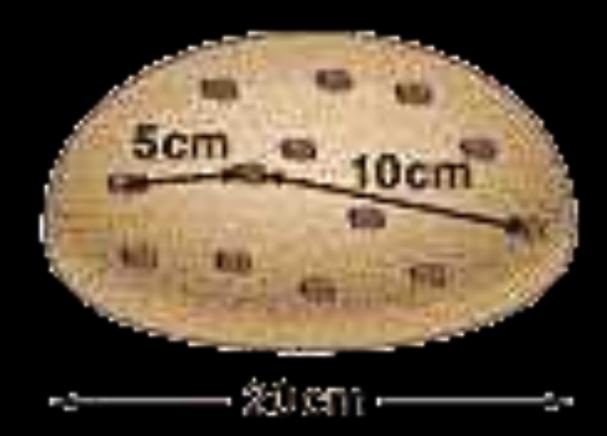
All galaxies move away from each other.

All galaxies have mass (stars, planets, etc.) which means they gravitationally pull on each other.

All this pulling **should** slow down the Hubble Expansion, forcing distant galaxies to slow down.

But that's not what we see!!!!!

### Way better visualization of Hubble's Law (1929) that includes bread



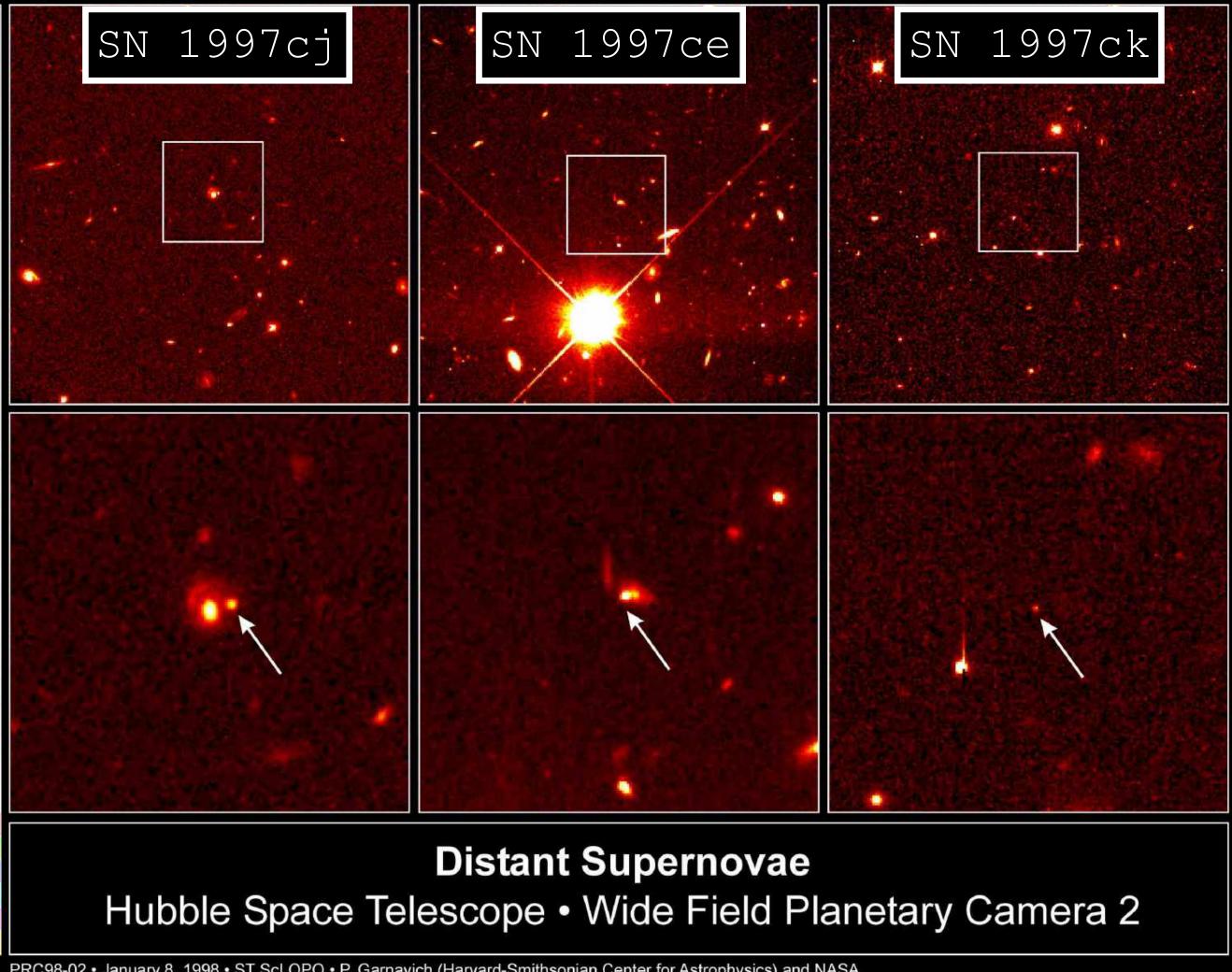


Instead, in 1997, astronomers found something very weird.

Distant supernovae appeared fainter than expected.

Thus, they must be farther away than we thought.

Peter Garnavich, Harvard-Smithsonian Center for Astrophysics, the High-z Supernova Search Team, and NASA/ESA



PRC98-02 • January 8, 1998 • ST Scl OPO • P. Garnavich (Harvard-Smithsonian Center for Astrophysics) and NASA



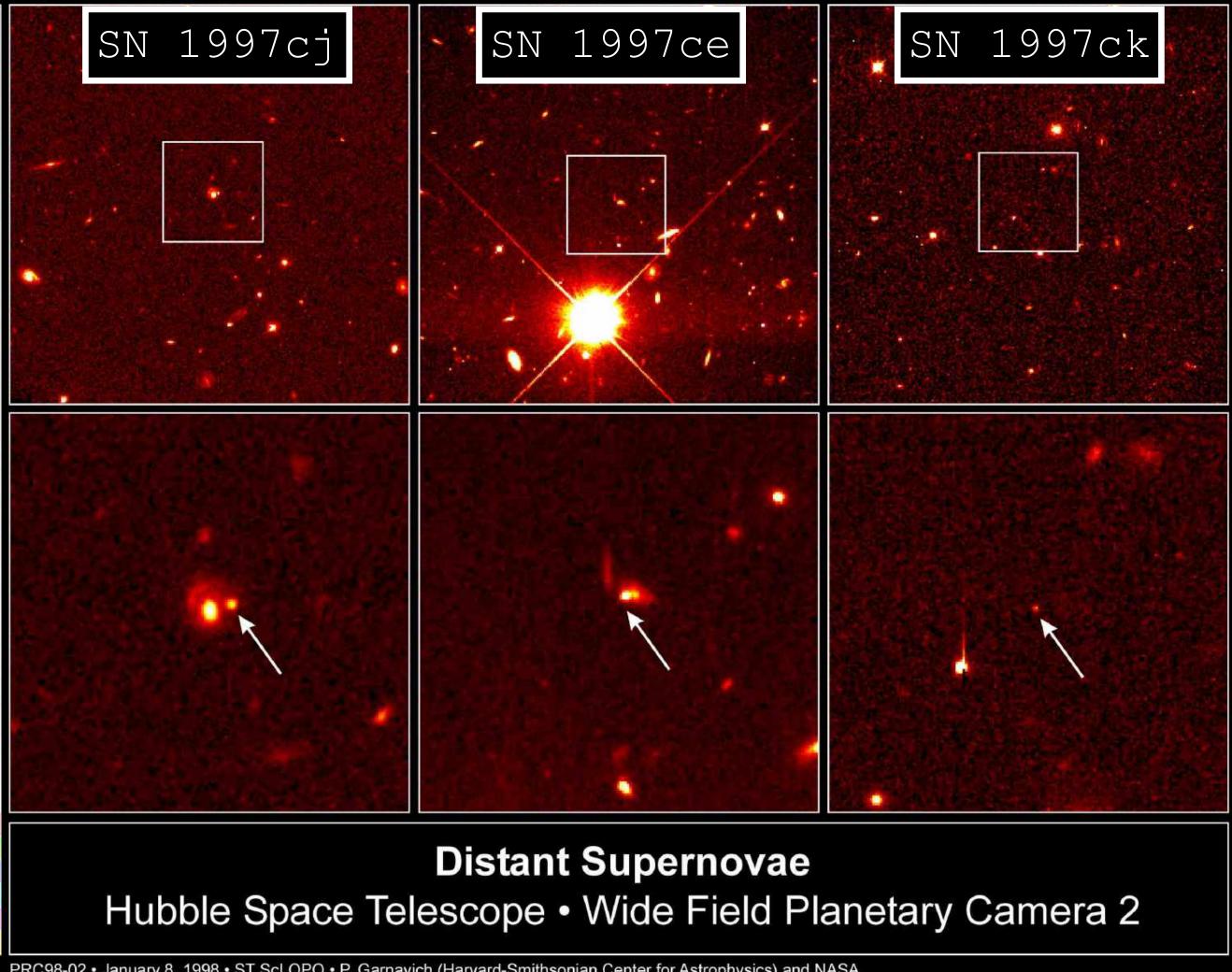
Instead, in 1997, astronomers found something very weird.

Distant supernovae appeared fainter than expected.

Thus, they must be farther away than we thought.

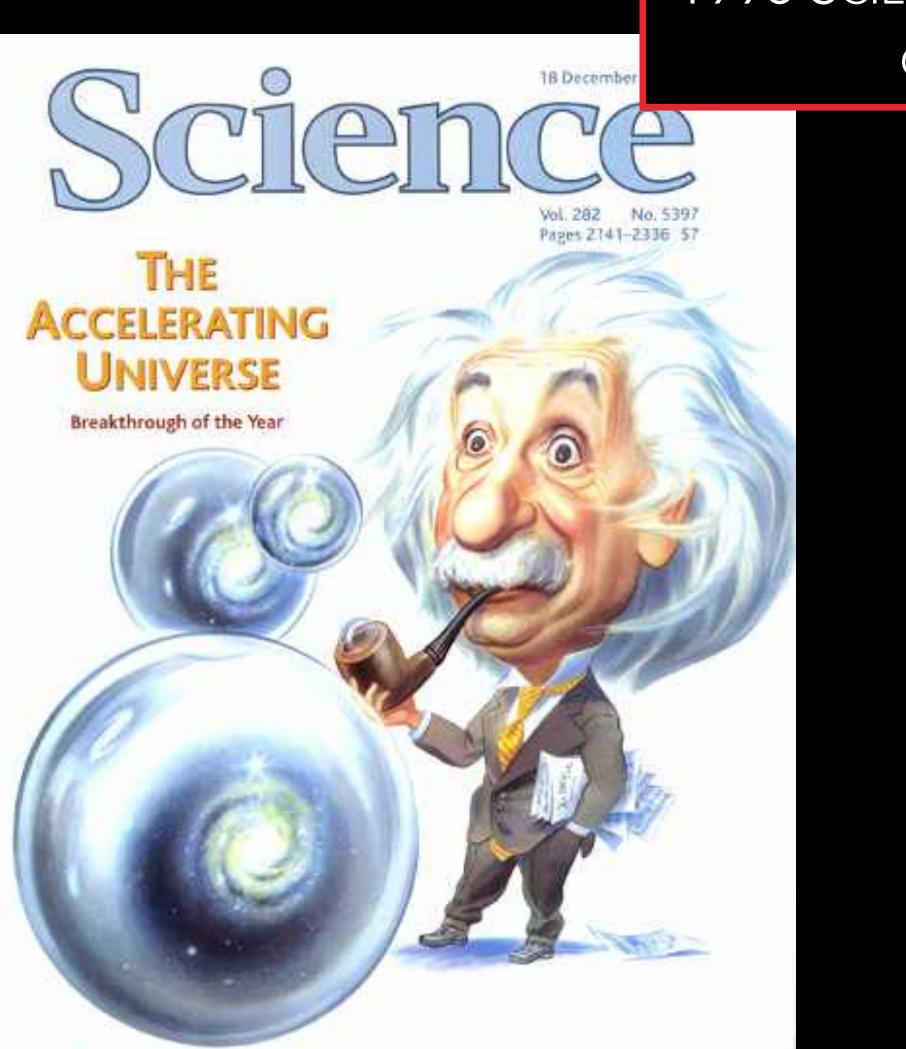
Something must be forcing the Universe to expand faster and faster to explain the dim supernovae, not slowing down as expected. But what?

Peter Garnavich, Harvard-Smithsonian Center for Astrophysics, the High-z Supernova Search Team, and NASA/ESA



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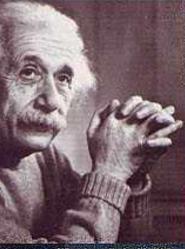
### **1998 SCIENTIFIC BREAKTHROUGH** OF THE YEAR

### SPACE EINSTEIN'S REPULSIVE IDEA He invented antigravity in desperation and

abandoned it first chance he got-but it may be the most powerful force in the universe

### By MICHAEL D. LEMONICK

LBERT EINSTEIN NEVER DID LIKE THE idea of antigravity. It wasn't that he had a problem with farfetched noal relativity theories made the astonishing assertion that time, space and matter could be squeezed and stretched like so much India rubber. The trouble was that some sort of antigravity force-Einstein called it the "cosmological term"-was required to make the predictions of general relativity



### NOT SO DUMB What he called a blunder may have been a Nobel-callber discover

match what astronomers believed the actual universe looked like. And that extra term marred the mathematical elegance of pushing it his beloved equations. The great physicist What that something might be is, at

greatest blunder."

### **O** GRAVITY WHAT IT IS: An

attractive force that pulls matter together like a rubber band

HOW IT OPERATES: Gravity weakens over distance; when the distance between two galaxies doubles, the force between them is one-fourth as strong

WHAT THAT MEANS: As the universe expands, gravity is less and loss effective at slowing the

and study a distant supernova-an exploding star-astronomers from two rival

research teams have jointly gathered the strongest evidence yet that the expansion of the universe is actually speeding up, like a rocket with its throttle wide open. And that means something is

was hugely relieved when the discovery of the expanding universe in the 1920s let full of theorists," says Adam Riess of the him cross out what he declared was "my greatest blunder." Space Telescope Science Institute in Balti-more, Md., leader of the collaboration, "and But he might have been a bit too hasty. 20 ideas will fall out." For now, the un- years ago, when two Last week scientists made a powerful case | known force is simply being called "dark en- independent teams of astronthat Einstein's blunder may actually have been another Nobel-worthy prediction. Using the Hubble Space Telescope to find

TIME, APRIL 16, 2001

Super nova 11 billion years ago, observed by Hubble telescope over past few years

ness of such a supernova with how bright it that dark energy, functioning as a kind of the stars will simply wink out-and the uniappears in the sky gives a good measure of antigravity, was their best guess. how far away it is and thus how long ago Critics argued that there might be a meekest of whimpers.

EXPANSION SLOWING DOWN

llig Bang 14 billion

Vears ago

EXPANSION SPEEDING UP

WHAT IT IS: A property of empty space that exerts an outward force like a compressed spring at every point in space

HOW IT OPERATES: A given volume of space always has the same amount of dark energy, so when the distance between two galaxies doubles, the force pushing them away from each other is twice as strong

WHAT THAT MEANS: As the universe expands the volume of space increases, which means more dark energy. By now, 14 billion years after the Big Bang, antigravity has overwhelmed gravity aster and faster

in cosmic history its light to such a project-and indeed, Perlmutter was emitted. Then, by and others are working on that idea, almeasuring how fast each su- though it will take years to get off the pernova is moving away from groun Earth in the overall ballooning If space really does see he with dark of the universe, it can be deter- energy, the fate of the universe, a matter of mined what the expansion rate was | longstanding debate, will be clear. With at different times in the past. more dark energy today than yesterday, and To everyone's astonishment, both more of the stuff tomorrow than today, the

groups found that instead of the gradual, cosmos should fly apart faster and faster as gravity-driven slowdown they expected, the rate was getting faster. Says Saul Perl-as some have predicted, with billions of mutter of Lawrence Berkeley National Laboratory in California, who heads one apocalypse. Tens of billions of years from of the groups: "We spent at least a year | now, our Milky Way galaxy will find itself is highly consistent. struggling to understand what we were alone in empty space, with its nearest neighbors too far away to see. In the end,

TIME: APRIL 16: 2001

tergalactie dust, which could contaminate the brightness measurements. But the new observations seem to have closed that loopbole. The newly identified supernova went off about 11 billion years ago-about 50% record holder. "If the dust were there," says Lawrence Berkeley astrophysicist Pe-ter Nugent, a member of Perlmutter's team much dimmer than it was."

more conventional explanation, such as infurther back in time than the previous and Riess's collaborator on the new re-search, "the supernova would have been The new supernova's remoteness was even more important for another reason. "If dark energy is really the explanation for what we see," says Riess, a member of the rival team, "then its effect should have been weaker in the early universe." That's because while the force of gravity between galaxies falls as they move farther apart, dark energy is a property of space and gets stronger as the universe expands. Shortly after the Big Bang, when the universe took up relatively little space, there wasn't much dark energy. Now much bigger, the modern universe has more space and thus more energy to shove galaxies apart. Sure enough, this distant supernova shows that the expansion was slower long ago. While the new observations go a long way toward confirming that dark energy is real, astronomers would love to see a few more distant supernovas, just to be sure. Unfortunately, that won't be happening soon. The Hubble pictures that Riess and Nugent analyzed were all taken purely by chance, while the telescope was looking for other things. Aiming at distant galaxies in hopes a supernova will go off is an ineffi-cient use of the telescope's valuable time. The best bet would be a satellite devoted verse will end not with a bang but with the

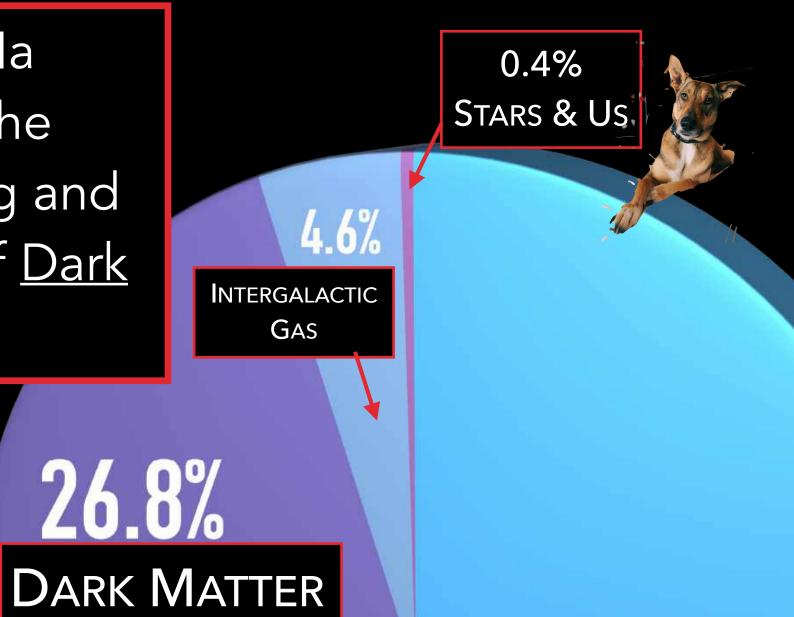


Observations of distant Type Ia supernovae have shown that the Universe is "currently accelerating and that over 2/3 of it is in the form of <u>Dark</u> <u>Energy.</u>"

### Nobel Prize Ceremony, 2011

My advisor, Peter Nugent

Saul Perlmutter, Nobel Laureate in Physics



### DARK ENERGY

68.2%



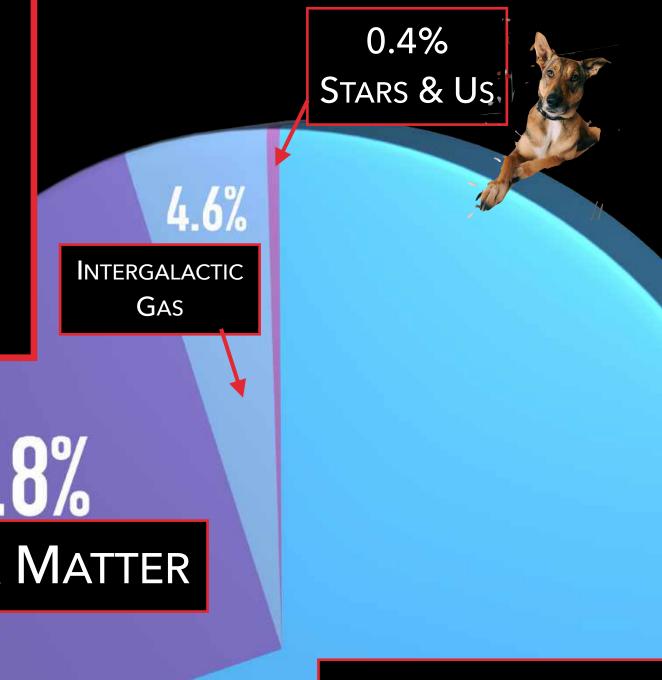


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### Nobel Prize Ceremony, 2011

advisor Peter Nugen

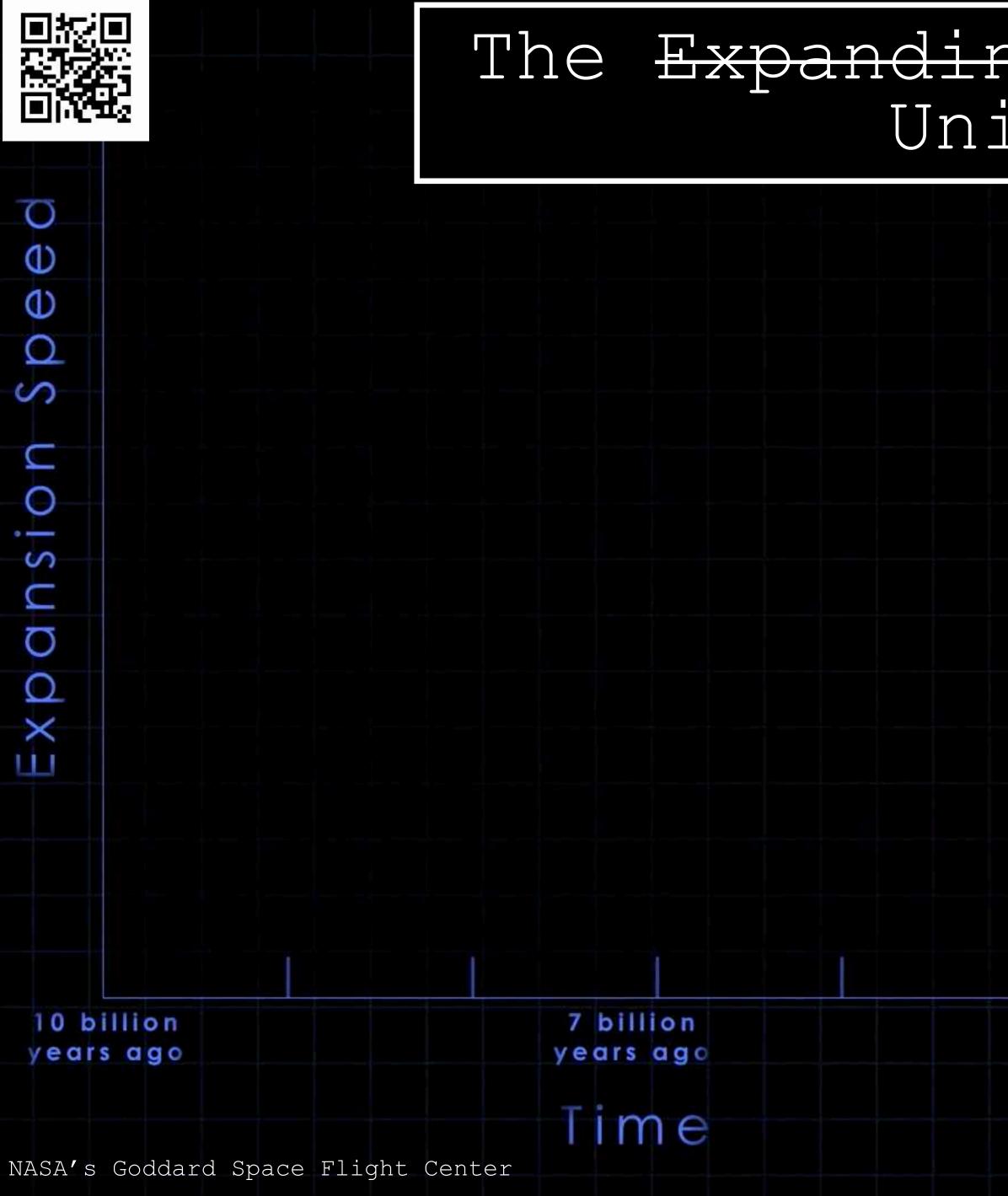
**IMPORTANT:** Dark Energy IS NOT THE SAME AS Dark Matter



### DARK ENERGY

**68.2%** 





## The <del>Expanding</del> Accelerating Universe







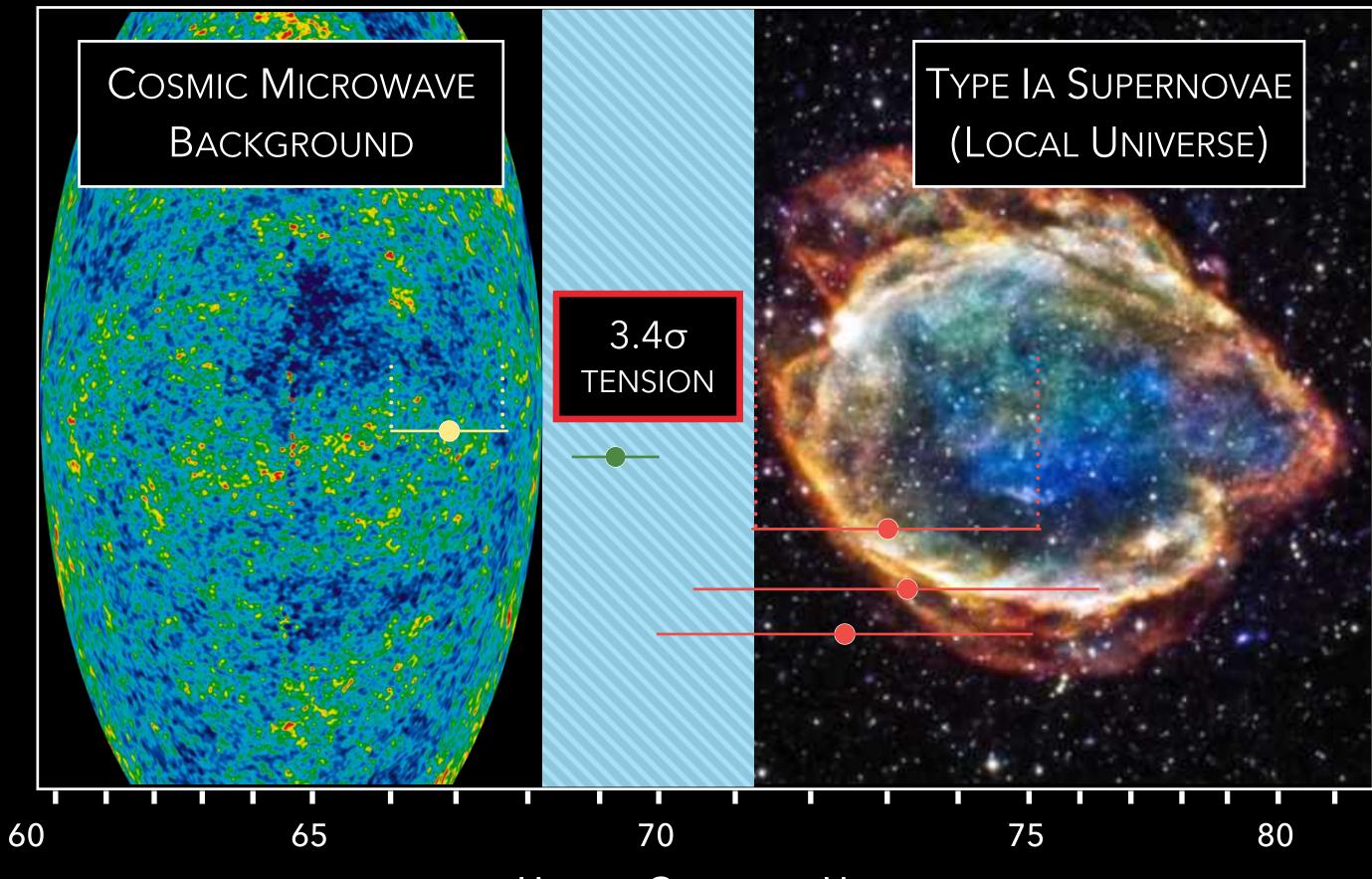


NASA/STScI/G. Bacon



### But this expansion rate is actually not well-constrained

### EARLY UNIVERSE observations give a 'low' value



HUBBLE CONSTANT H<sub>0</sub> (km s<sup>-1</sup> Mpc<sup>-1</sup>)

LOCAL UNIVERSE (present-day) observations give a 'high' value





## But this expansion rate is actually not well-constrained

EARLY UNIVERSE observations give a 'low' value

60

- Need an independent method to VAE resolve this tension.
  - Traditional local measurements of the Hubble Constant are Type Ia supernovae.
  - I want to push the envelope by using a totally independent measure that hasn't been used before: Type IIP
    - supernovae.
    - HUBBLE CONSTANT H<sub>0</sub> (km s<sup>-1</sup> Mpc<sup>-1</sup>)

LOCAL UNIVERSE (present-day) observations give a 'high' value





### Remember Type II Supernovae?



Pressure from *neutrinos* that drives the expansion & subsequent explosion







(aka big stars that explode without the help of a companion star)

> I use these babies (specifically Type IIP) to try to measure the rate of the expansion of the Universe!





# The expansion rate tells us something about the Fate of the Universe





NASA/STScI/G. Bacon

# The expansion rate tells us something about the Fate of the Universe





NASA/STScI/G. Bacon

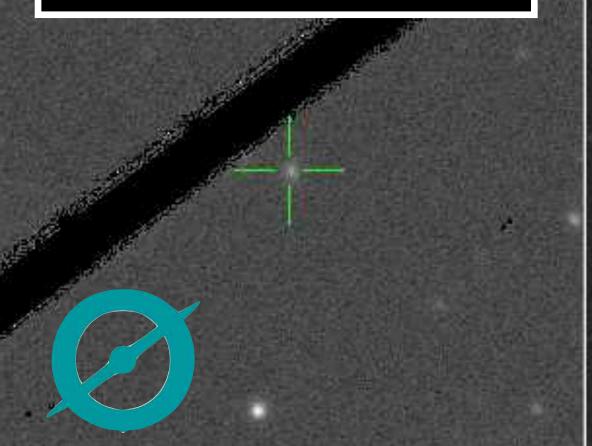


### We need your help finding supernovae!

Help discover supernovae from Pan-STARRS1.

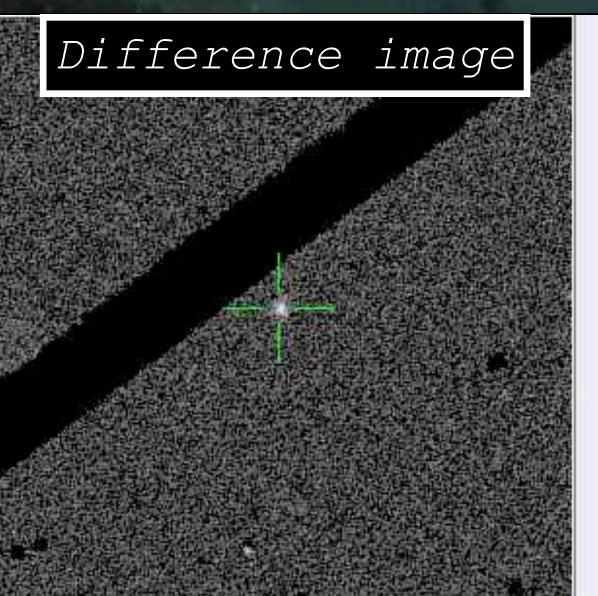


### Target image (SN "detection")



Reference image

http://supernova.galaxyzoo.org



TASK

TUTORIAL

Does the source centred in the green crosshairs in the rightmost image look like a detection of a real supernova?

Yes

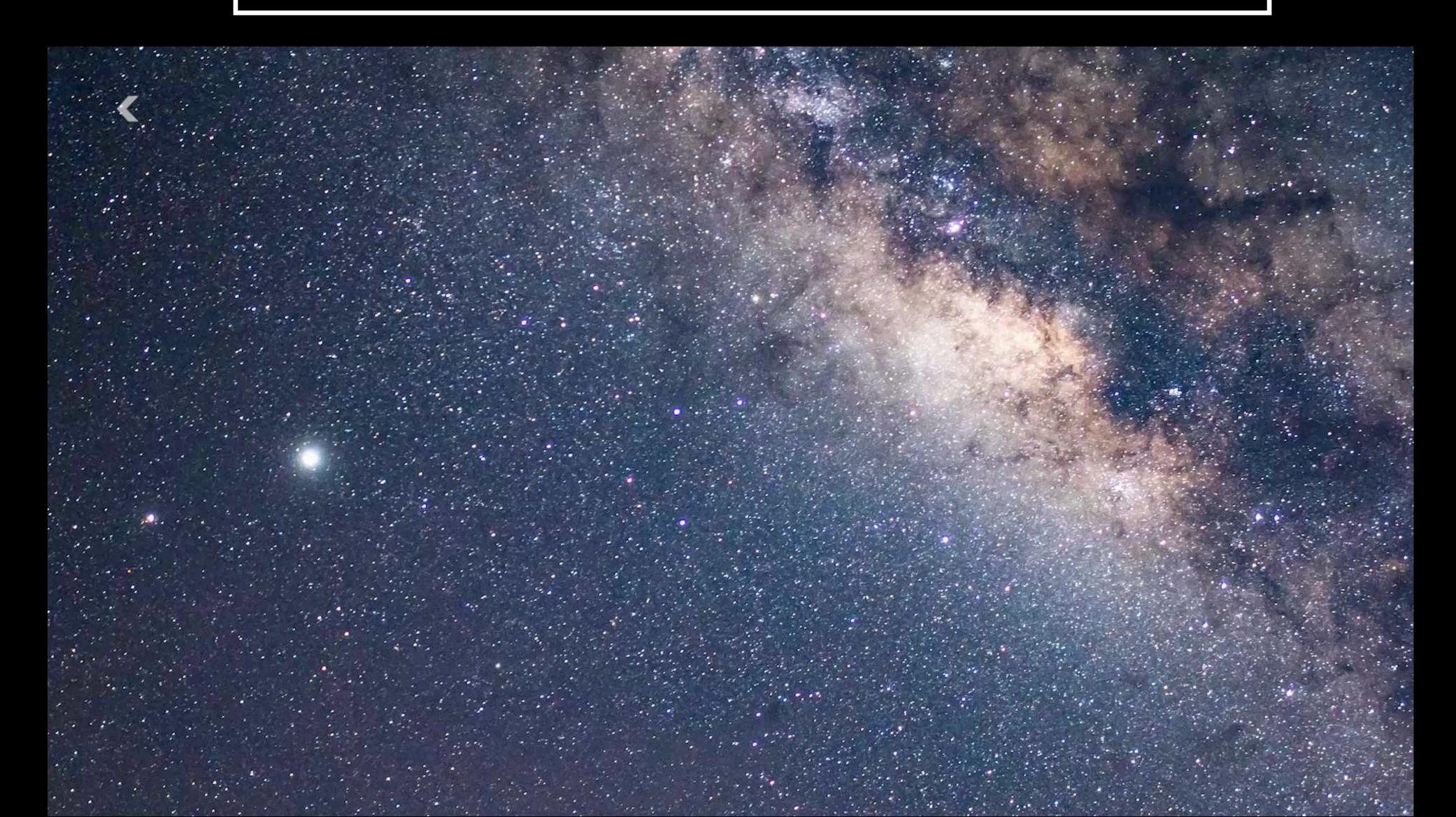
No

NEED SOME HELP WITH THIS TASK?





# Explore the cosmos with me in my new show CONSTELLATIONS





SUPERNOVAE ARE EXPLODING STARS. THEY CAN BE 2-STAR OR 1-STAR SYSTEMS.

BETELGEUSE IS (PROBABLY NOT) GOING TO EXPLODE ANYTIME SOON.

The Universe is expanding. That expansion is accelerating thanks to DARK ENERGY. WE FOUND THAT OUT BY STUDYING SUPERNOVAE.





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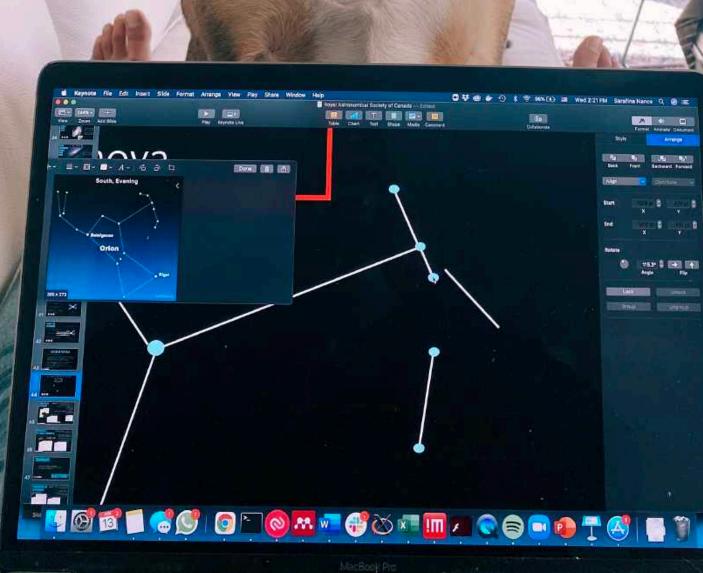
## IF YO FROM 7

SUPERNOVAE ARE EXF

BETELGEUSE IS (PRC)

THE UNIVERSE IS E> DARK ENERGY. WE

YOU CAN HELP US



## HING ER THIS:

-STAR SYSTEMS.

YTIME SOON.

LERATING THANKS TO

RNOVAE.

