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## Automated data analysis

- High resolution MF (MDI) EGSO project (2002-2005)
  - sunspot magnetic field SMF
- Automated detection -
- **Solar Feature Catalogues:**
- Zharkova et al., 2005, Sol Phys, 228, 365



 Low resolution MF (WSO) – synoptic maps of SBMF

## **Solar Feature Catalogues - EGSO**

Zharkova et al., 2005, Sol Phys, 228, 365 http://solar.inf.brad.ac.uk

- Sunspot Catalogue (from 1996-05-19 19:08:35 to 2010-05-31 19:51:32)
  - Automated feature detection and data extraction
  - About 40000 observation processed
  - ~370,000 sunspots and 100 000 ARs stored and processed

#### • Sunspot Catalogue (SOHO MDI) (Zharkov et al., 2005)

- Space Observations, Accuracy & Image Quality
- Synoptic Continuum images every 6 hour
- LOS Magnetogram Data

#### • AR Catalogue (Meudon+MDI) (Benkhalil et al., 2006)

- Meudon Ca II K3 images
- Meudon H-alpha images
- MDI LOS Magnetograms
- Filaments and prominences (Meudon) (Fuller et al., 2005)
- Meudon H-alpha images



UT (1996-01-01 To 2004-12-31)

#### **SBMF (top) and sunspot MF (bottom) phase** between them is π~11y (Zharkov et al, 2008, Stix 1976)

#### Cycle 23 -Solar Background MF



MF flux Sun Spots



#### 2. 1y-4y residuals for BMF (top) and excess SMF (bottom) reveal additional phase of π/4 ~ 2.5 years – Zharkov et al, 2008



MF flux Sun Spots 1y-4y



Time, in CR



## PCA acts as a prism for magnetic waves Philosophy of PCA

- Introduced by Pearson (1901) and Hotelling (1933) to describe the variation in a set of multivariate data in terms of a set of uncorrelated variables
- We typically have a data matrix of *n* observations on *p* correlated variables

 $x_1, x_2, \dots x_p$ 

PCA looks for a transformation of the x<sub>i</sub> into p new variables y<sub>i</sub> that are uncorrelated

## To reduce dimension – weighted average

- Weighted average based on some criterion. Which one? -→
- Looking for a transformation of the data matrix X (nxp) such that

$$Y = \boldsymbol{\delta}^T \boldsymbol{X} = \delta_1 X_1 + \delta_2 X_2 + \ldots + \delta_p X_p$$

• Where  $\delta = (\delta_1, \delta_2, ..., \delta_p)^T$  is a column vector of weights with

$$\delta_1^2 + \delta_2^2 + ... + \delta_p^2 = 1$$

## **One good criterion**

- Maximize the variance of the projection of the observations on the Y variables
- Find  $\delta$  so that

 $Var(\delta^T \mathbf{X}) = \delta^T Var(\mathbf{X}) \delta$  is maximal

 The matrix **C=Var(X)** is the covariance matrix of the X<sub>i</sub> variables

# Calculating eignevalues and eigenvectors

• The eigenvalues  $\lambda_i$  are found by solving the equation

 $det(C-\lambda I)=0$ 

• Eigenvectors are columns of the matrix A such that



## So PCA gives

- New variables  $Y_i$  that are linear combination of the original variables  $(x_i)$ :
- $Y_i = a_{i1}x_1 + a_{i2}x_2 + \dots + a_{ip}x_p$ ; i=1..p
- The new variables Y<sub>i</sub> are derived in decreasing order of importance;
- they are called 'principal components'

### **Interpretation of PCA**

- The new variables (PCs) have a variance equal to their corresponding eigenvalue  $Var(Y_i) = \lambda_i$  for all i=1...p
- Small  $\lambda_i \Leftrightarrow$  small variance  $\Leftrightarrow$  data change little in the direction of component  $Y_i$
- The relative variance explained by each
   PC is given by λ<sub>i</sub> /Σλ<sub>i</sub>
- PCs can be assigned to separate physical processes

## **SBMF results: Scree plot-Eigenvalues vs variances**

 2 main eigenvalues covering 40% of variance – dipole source



#### Two magnetic waves of the opposite polarities extracted in SMF with PCA (Zharkova et al., 2012, MNRAS)





#### **Derivatives of 2 main EOFs**



From top left figure above

- component 1
- component 2

resultant

These latitudinal ICs were modeled with the updated 2layers Parker's dynamo fitting amplitudes, phase shifts and number of equator crossings Popova et al, 2013

#### **Cross-correlation of 8 EOFs**



## Classification of and predicted principal components

(Zharkova et al, 2012, Shepherd et al, 2014)



## **Summary curve**



Space Climate 6, 4 April I2016

- Mathematical laws from PCs: Symbolic regression -Hamiltonian approach Schmidt & Lipson, 2009, Science) Mathematical law for the first principal component:
- $F_{1}(t) = \sum_{k=1,..,5} A_{k} \cos(\omega_{k,1}t + \phi_{k,1}) \cos(B_{k,1} \cos(\omega_{k,1}t + \phi_{k,1}))$ 
  - Mathematical law for the second principal component:
- $F_2(t) = \sum_{k=1,..,5} A_k \cos(\omega_{k,2}t + \phi_{k,2})\cos(B_{k,2} \cos(\omega_{k,2}t + \phi_{k,2}))$

## Fitting the PCs to measured in cycle 24 and prediction to 25-26



Space Climate 6, 4 April I2016

#### Summary PC and modulus summary PC vs sunspot data (Shepherd et al, 2014, ApJ)



Space Climate 6, 4 April I2016

## **Predicted summary curve on the millennium timescale**



Periods - grand:350-400 years and super-grand:1900-2000 years

## Updated curve for 3000 years (blue) versus a curve by Usoskin et al. (red)



## 2 layer dynamo model explaining some PCA features

- Dynamo model was not even considered yet while we did PCA in 2010-12 and SEA 2014
- Started discussing possible mechanisms since the end of 2010
- In 2011 we considered 2 layer Parker's model (1993) with meridional circulation
- 2013 first model paper appeared in Annals in Geophysics (Popova et al, 2013)

#### **Dynamo in Two-Layer Medium**



We included the meridional flows in each layer:

$$\frac{\partial B}{\partial t} + \frac{\partial (VB)}{\partial \theta} = \beta \Delta B, \qquad \frac{\partial A}{\partial t} + V \frac{\partial A}{\partial \theta} = \alpha B + \beta \Delta A, \tag{2.3}$$

$$\frac{\partial b}{\partial t} + \frac{\partial (vb)}{\partial \theta} = D\cos\theta \frac{\partial a}{\partial \theta} + \Delta b, \qquad \frac{\partial a}{\partial t} + v \frac{\partial a}{\partial \theta} = \Delta a, \qquad (2.4)$$

here  $V(\theta), v(\theta)$  are the meridional flows in the respective layers.

Following Parker we prescribe r = 0 for the radial boundary between two layers and use boundary conditions:

$$b = B, \qquad a = A, \qquad \frac{\partial b}{\partial r} = \beta \frac{\partial B}{\partial r}, \qquad \frac{\partial a}{\partial r} = \frac{\partial A}{\partial r}.$$
 (2.5)

In view of the symmetry conditions  $\alpha(-\theta) = -\alpha(\theta)$ ,  $V(-\theta) = -V(\theta)$  the above system of equations can be considered in only one (e.g., the northern) hemisphere using antisymmetry (dipolar symmetry) or symmetry (quadrupolar symmetry) conditions at the equator.

We obtained Hamilton-Jacobi equation for eqs. (2.3) and (2.4) by a method similar to the method described in Popova et al. (2010).



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#### **Quadrupole components of poloidal magnetic field (BMF)**



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Undamped Wave Equation: Solution to Initial Value Problem (2 of 3)

Solution of a wave equation with forced oscillations

$$y(t) = \frac{F_0}{m(\omega_0^2 - \omega^2)} \left(\cos \omega t - \cos \omega_0 t\right)$$

• To simplify the solution even further, let  $A = (\omega_0 + \omega)/2$  and  $B = (\omega_0 - \omega)/2$ . Then  $A + B = \omega_0 t$  and  $A - B = \omega t$ . Using the trigonometric identity  $\cos(A \pm B) = \cos A \cos B \mp \sin A \sin B$ ,

it follows that  $\cos \omega t = \cos A \cos B + \sin A \sin B$  $\cos \omega_0 t = \cos A \cos B - \sin A \sin B$ 

and hence  $\cos \omega t - \cos \omega_0 t = 2 \sin A \sin B$ 

#### Undamped Equation: Beats (3 of 3)

Using the results of the previous slide, it follows that

$$y(t) = \left[\frac{2F_0}{m(\omega_0^2 - \omega^2)}\sin\frac{(\omega_0 - \omega)t}{2}\right]\sin\frac{(\omega_0 + \omega)t}{2}$$

- When  $|\omega_0 \omega| \stackrel{\scriptscriptstyle \sim}{=} 0$ ,  $\omega_0 + \omega$  is much larger than  $\omega_0 \omega$ , and  $sin[(\omega_0 + \omega)t/2]$  oscillates more rapidly than  $sin[(\omega_0 \omega)t/2]$ .
- Thus motion is a rapid oscillation with frequency  $(\omega_0 + \omega)/2$ , but with slowly varying sinusoidal amplitude given by

$$\frac{2F_0}{m|\omega_0^2-\omega^2|}\left|\sin\frac{(\omega_0-\omega)t}{2}\right|$$

This phenomena is called a **beat**.
Beats occur with two tuning forks of nearly equal frequency.







## **Conclusions:**

## **EOFs components: cycles 21-23**

- Principal components of SBMF are paired
- The strongest PCs cover more than 40% of variance
- These PCs are shown to reflect 2 dynamo waves travelling with increasing phase shift from one hemisphere to another
- The waves intercept with the increased turbulence one year prior and after the cycle maximum
- Cross-correlation shows a presence of quadruple sources in all the cycles and possible sextuple ones in cycle 23
- Mathematical laws derived with Hamiltonian approach (Euriqa) was used for prediction of the reduction of the solar activity in cycles 25-27 – next Maunder Minimum

Prediction for 3000 years backwards fits the main warming and cooling periods – Sun gave us the clues!