# Classical Physics – Fall 2019

## Section A: Mechanics

- **3)** A particle of mass *m* moves in a central force with potential energy  $U(r) = Kr^4$ .
- **a.** For the force to be attractive, what is the sign of *K*? Don't guess. Prove it.
- **b.** What is the effective potential energy, V(r)?
- **c.** For a given value of angular momentum, *l*, find *r* for a circular orbit.
- **d.** What is the total energy of the circular orbit? Express your answer in terms of K and  $r_0$  is the radius of the circular orbit.
- **e.** What is the period of the circular orbit? Express your answer in terms of K, m, and  $r_0$ .

#### Solution)

## a)

The central force is defined as  $F = -\nabla U$ 

$$\mathbf{F} = -\frac{\partial U}{\partial r}\hat{r} = -4Kr^3\hat{r}$$

This corresponds with a cubic restoring force. For positive K the minus sign ensures that the force is toward the center of force, for negative K the force will point away from the center of force. Thus:

#### K > 0: attractive

## b)

The effective potential is defined as:

$$V(r) = U(r) + \frac{l^2}{2mr^2}$$
$$V(r) = Kr^4 + \frac{l^2}{2mr^2}$$

## c)

The kinetic energy in polar coordinates is:

$$T=\frac{m}{2}\left(\dot{r}^2+r^2\dot{\theta}^2\right)$$

The Lagrangian has the form L = T - U:

$$L = \frac{m}{2} \left( \dot{r}^2 + r^2 \dot{\theta}^2 \right) - K r^4$$

The Lagrange equations of motion are:

$$\frac{d}{dt}\left(\frac{\partial L}{\partial \dot{q}}\right) - \frac{\partial L}{\partial q} = 0$$

For the generalized coordinate *r*:

$$\frac{d}{dt}\left(\frac{\partial L}{\partial \dot{r}}\right) - \frac{\partial L}{\partial r} = 0$$

$$\frac{\partial L}{\partial r} = mr\dot{\theta}^2 - 4Kr^3$$
$$\frac{\partial L}{\partial \dot{r}} = m\dot{r} \quad \rightarrow \quad \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{r}}\right) = m\ddot{r}$$

For the generalized coordinate  $\theta$ :

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{\theta}} \right) - \frac{\partial L}{\partial \theta} = 0$$

$$\frac{\partial L}{\partial \theta} = 0$$
$$\frac{\partial L}{\partial \dot{\theta}} = mr^2 \dot{\theta} \quad \rightarrow \quad \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{\theta}} \right) = mr^2 \ddot{\theta} + 2mr\dot{r}\dot{\theta}$$

The equations of motion are:

$$m\ddot{r} - mr\dot{\theta}^2 + 4Kr^3 = 0 \qquad (1)$$
$$\frac{d}{dt}(mr^2\dot{\theta}) = 0 \qquad (2)$$

For circular motion  $r=r_0=const$  ( $\dot{r}=0,\ddot{r}=0$ ) in (1):

$$-mr_0\dot{\theta}^2 + 4Kr_0^3 = 0 \qquad (3)$$

From (2):

$$mr^2\dot{\theta} = const = l$$

Plugin

$$\dot{\theta}^2 = \frac{l^2}{m^2 r_0^4}$$

into (3):

$$-mr_0 \frac{l^2}{m^2 r_0^4} + 4Kr_0^3 = 0$$
$$4Kr_0^3 = \frac{l^2}{mr_0^3}$$
$$r_0^6 = \frac{l^2}{4mK}$$

d)

The total energy is:

$$E = T + U$$
$$E = \frac{m}{2} \left( \dot{r}^2 + r^2 \dot{\theta}^2 \right) + Kr^4$$

At  $r = r_0$ :

$$E = \frac{m}{2}r_0^2\dot{\theta}^2 + Kr_0^4$$

where

$$\dot{\theta}^2 = \frac{l^2}{m^2 r_0^4} = \frac{4mKr_0^6}{m^2 r_0^4} = \frac{4Kr_0^2}{m}$$
$$E = \frac{m}{2}r_0^2 \frac{4Kr_0^2}{m} + Kr_0^4$$
$$E = 2Kr_0^4 + Kr_0^4$$
$$E = 3Kr_0^4$$

e)

Using the period of the orbits T:

$$\dot{\theta} = \frac{2\pi}{T}$$

where

$$\dot{\theta} = 2r_0 \sqrt{\frac{K}{m}}$$

We get:

$$T = \frac{2\pi}{\dot{\theta}}$$
$$T = \frac{\pi}{r_0} \sqrt{\frac{m}{K}}$$