

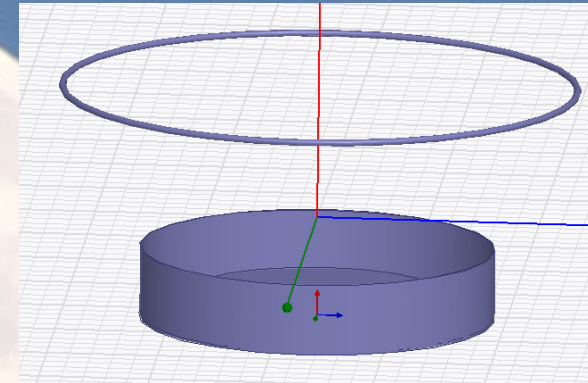
Circularly Polarized Patch Feed for 1296 MHz

Swedish EME-meeting May 2013

SM6FHZ and SM6PGP

Outline

- Prerequisite
- Features
- Design considerations
 - Size
 - Choke
 - BFR
- Performance
- Mechanical dimensions
- Realization
- On the air experience
- Conclusions



Why?

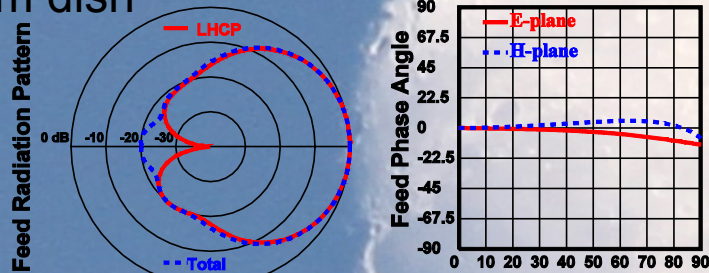
- SM6PGP wanted to use his 1.8 m dish on 23 cm EME
- He felt that a W/G feed was far too bulky in this small dish
- He figured that a patch feed with quadrature feeding might be the answer
- Together we set the goals and limitations for dimensions and performance
- Using the 70 cm feeds as a starting point the simulations soon showed a feasible design with good performance

Features

- Compact circularly polarized patch feed suitable for small dishes
- Top notch performance
- Focus on easy manufacturing and low tolerance sensitivity
- Suitable for f/D from 0.28 to 0.45 covered in two variants
- Comprising a choke (baffle) structure as well as a Beam Forming Ring

InDish Performance with BFR

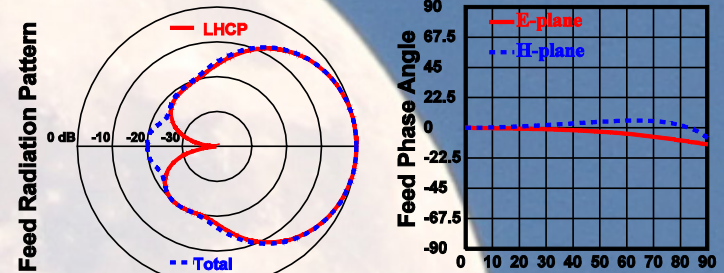
Performance in 23 cm patch feed with BFR
a 1.8 m dish



Dish diameter = 7.9λ Feed diameter = 0.7λ

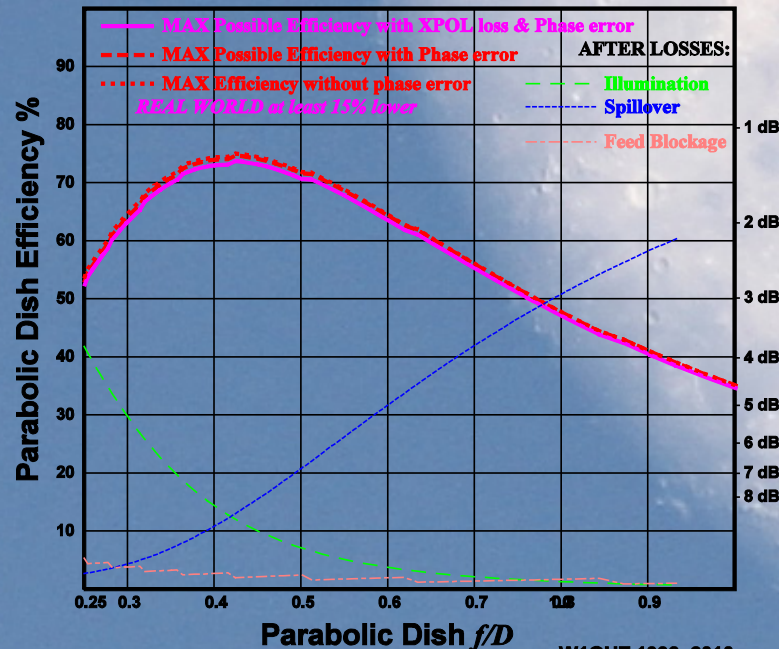
Rotation Angle around specified
Phase Center = 0.025λ beyond aperture

Performance in 23 cm patch feed with BFR
a 2.3 m dish

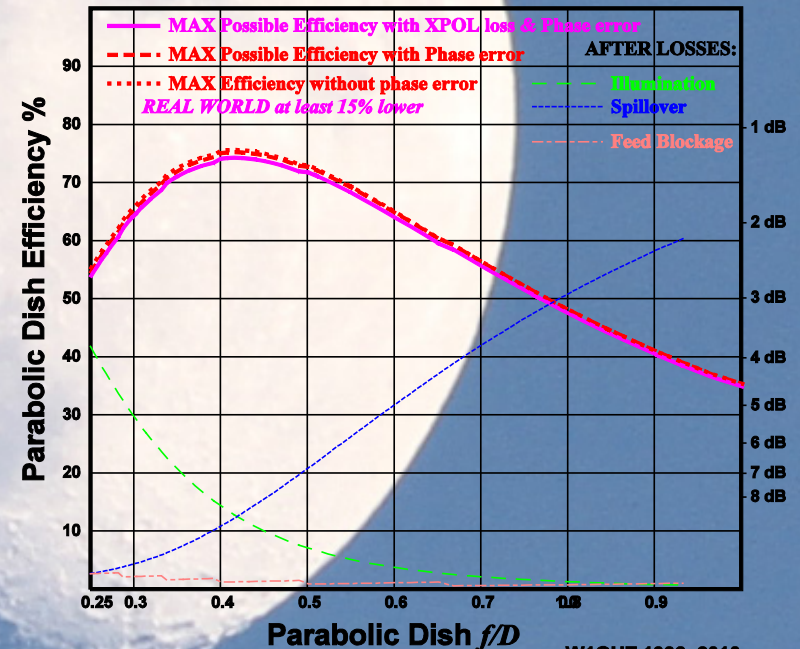


Dish diameter = 10λ Feed diameter = 0.7λ

Rotation Angle around specified
Phase Center = 0.025λ beyond aperture



W1GHZ 1998, 2010

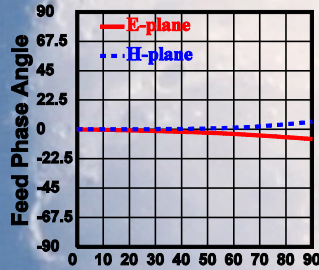
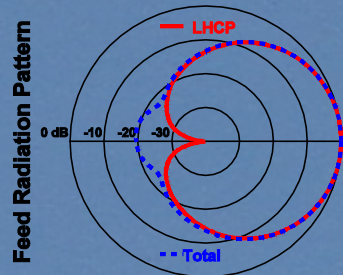


W1GHZ 1998, 2010

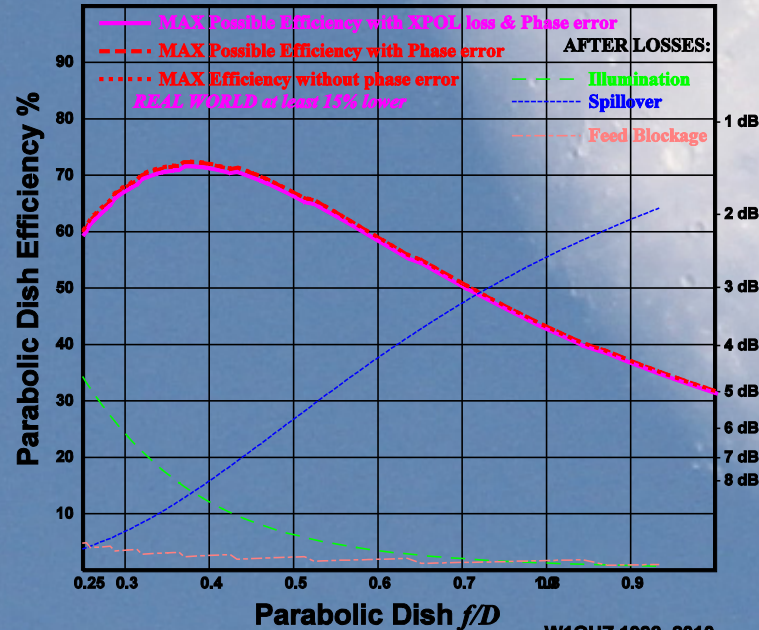
InDish Performance without BFR

Performance in
a 1.8 m dish

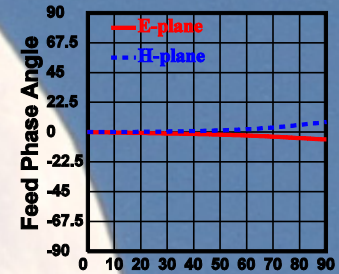
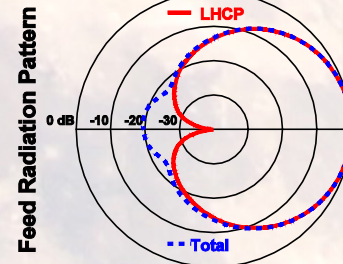
23 cm Patch feed



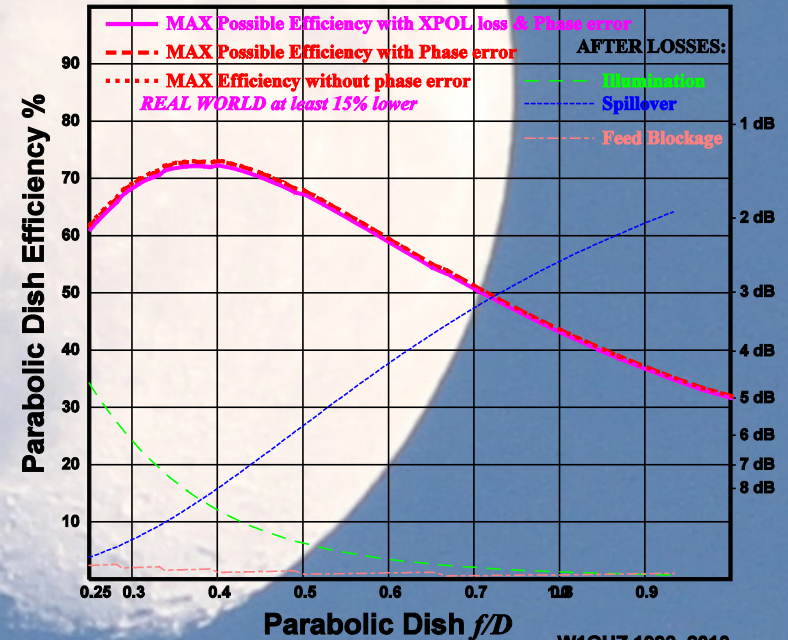
Dish diameter = 7.8λ Feed diameter = 0.7λ
Rotation Angle around specified
Phase Center = 0.03λ inside aperture



Performance in
a 2.3 m dish



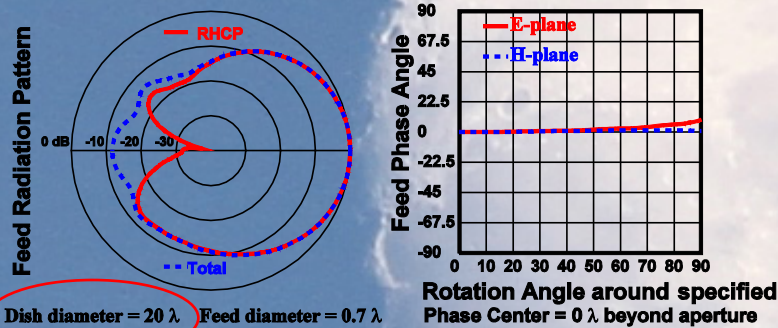
Dish diameter = 10λ Feed diameter = 0.7λ
Rotation Angle around specified
Phase Center = 0.04λ inside aperture



InDish Performance for two alternatives

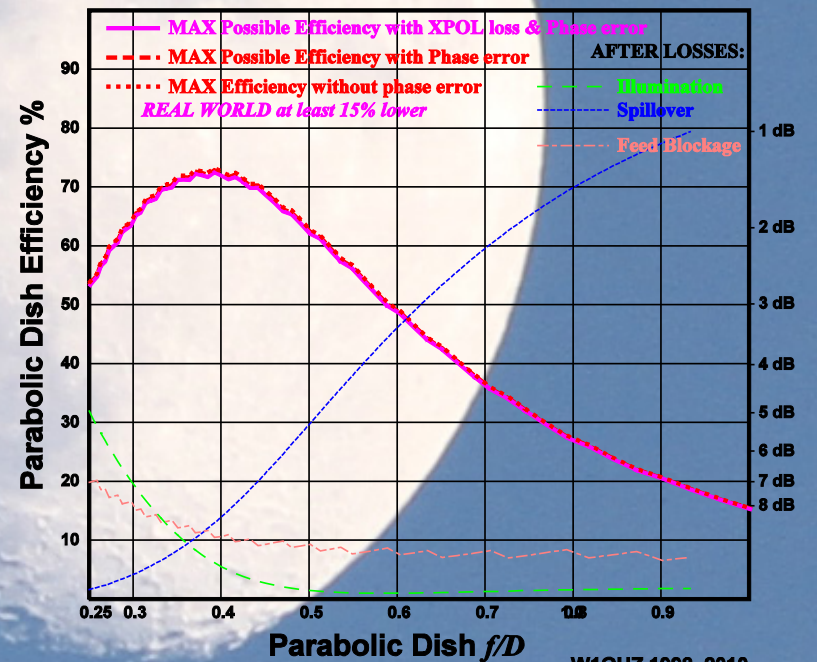
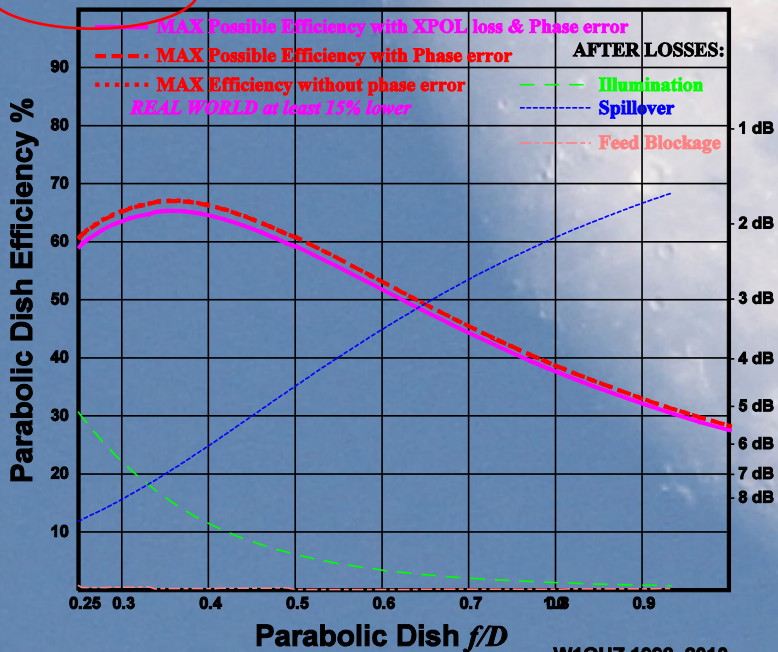
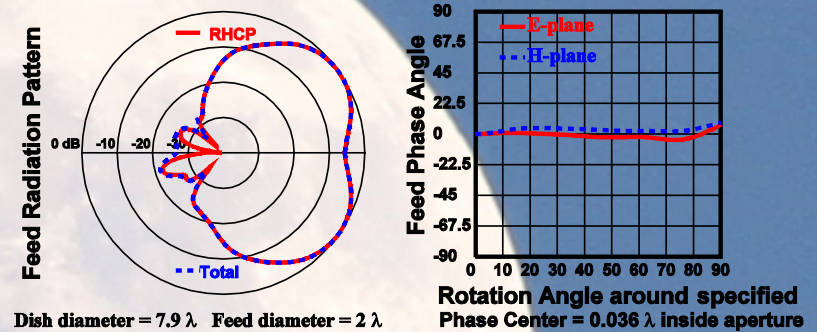
Performance with a DFC Septum feed

OK1DFC Septum Feed - nominal dimensions

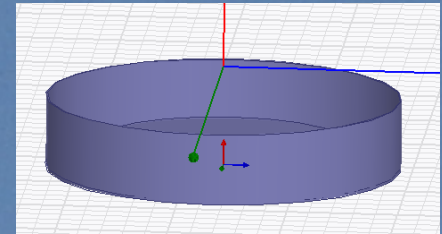


Performance with a FHZ Kumar feed

FHZ Kumar $0.14L$ in $7.9L$ dish



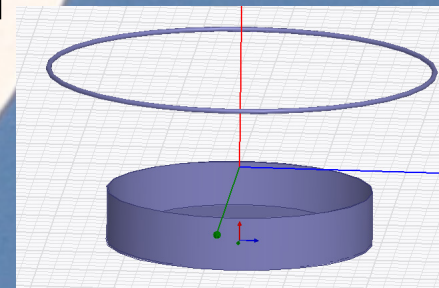
Why a choke



- The simulations on 70 cm feeds in 2010 showed that a patch feed did not have equal Beam Width in the two cardinal cuts. This need to be taken care of in order to get good efficiency together with decent noise performance.
- It was demonstrated on 70 cm that a choke (or baffle if you like) made the BW almost equal in all cuts.
- The choke also improved the Front to Back Ratio and phase error.

BFR - choke interaction

- The BFR sharpens up the main lobe
 - This leads to higher efficiency for f/D 's around 0.4
 - It also enhances the side radiation for the feed improving the antenna temperature with less ground noise pick-up
- The BFR does not affect the good properties from the choke

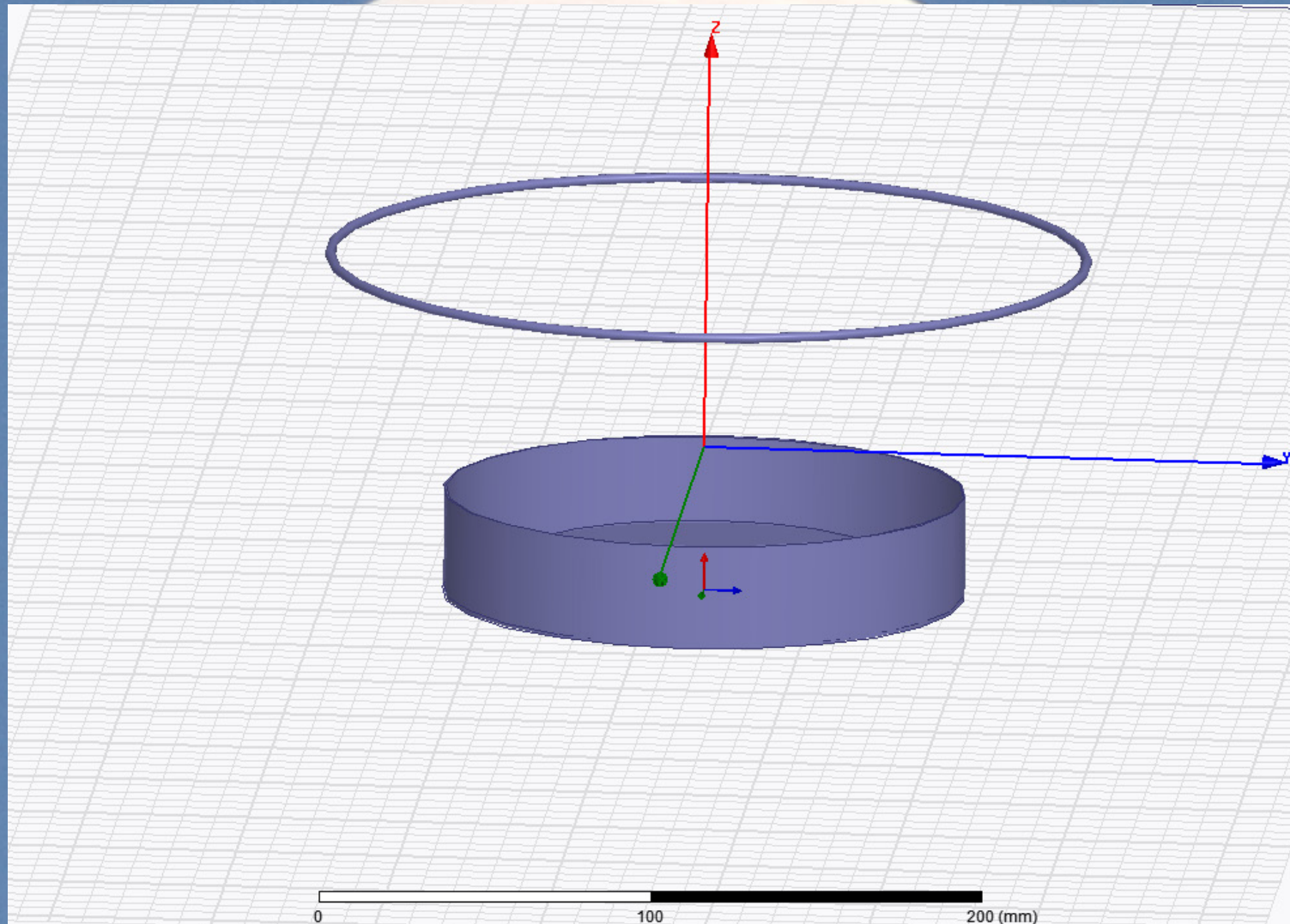


A large, bright, cratered moon is shown in a clear blue sky. The moon is the central focus, with its surface covered in numerous craters of various sizes. The text "BFR variant" is overlaid on the moon's surface.

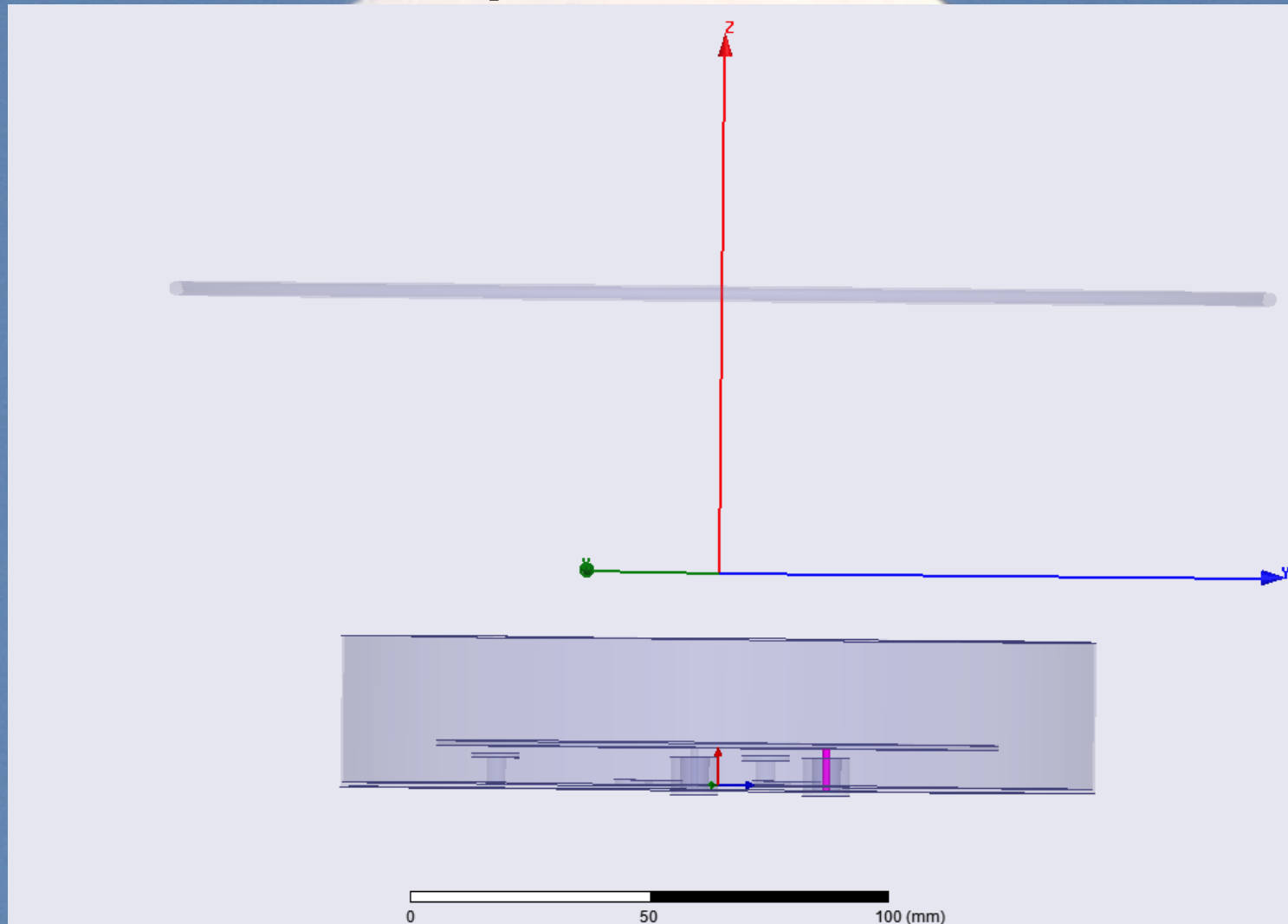
BFR variant

For dishes with f/D 's in the 0.34 to 0.43 range

Solid model from the simulation



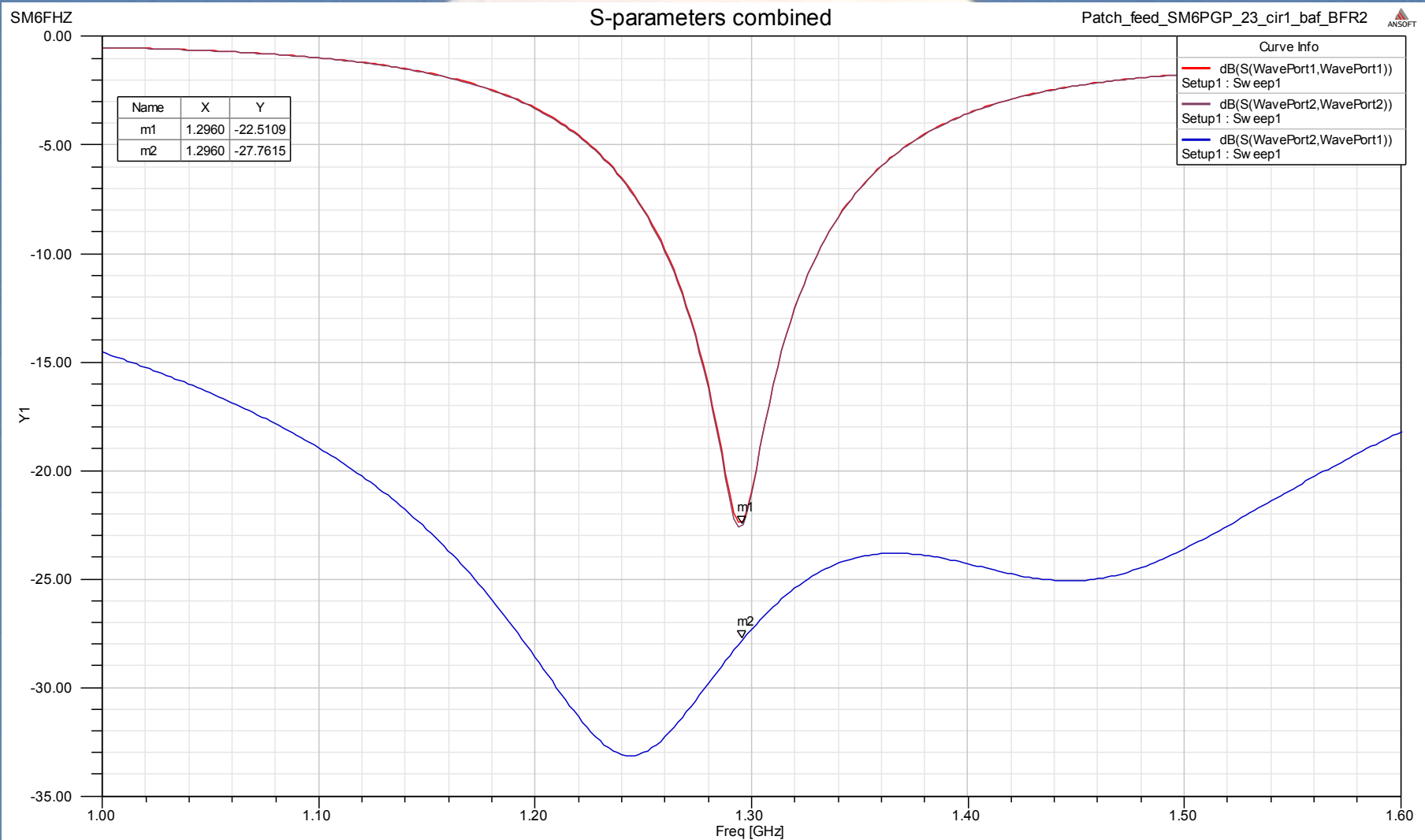
Transparent model



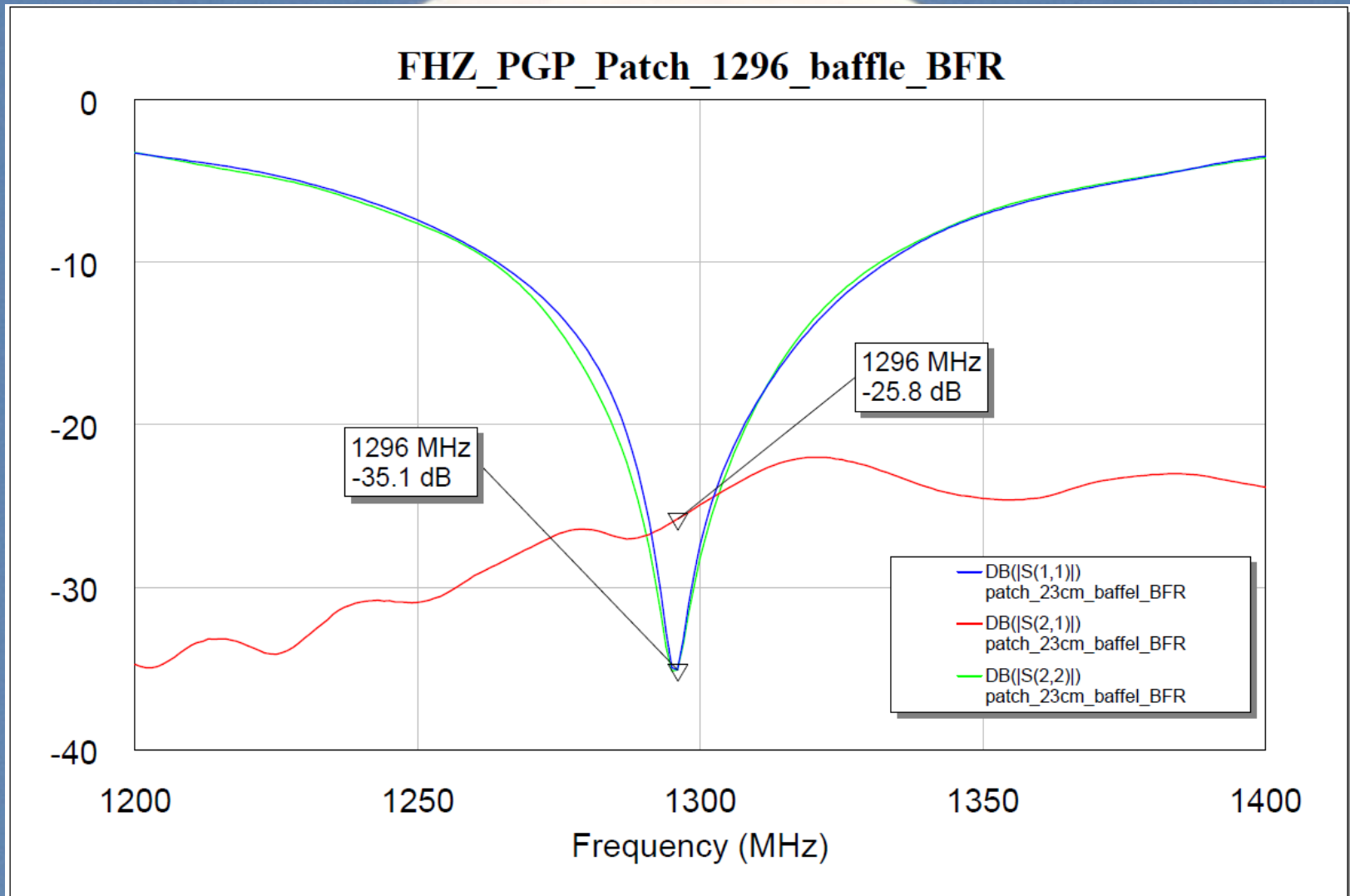
Dimensions with BFR

- Reflector 160 mm outer diam, 1 mm thick Cu or brass
- Baffle inner height 31 mm
- Baffle thickness 0.5 mm
- Patch 119 mm diam, 1 mm thick, 8 mm up from reflector (surface to surface)
- Probe 23.5 mm from center, $\frac{1}{4}$ " coax w 1.6 mm inner conductor. Outer conductor stops 2 mm from the patch
- Phase is center 51 mm above the reflector
- BFR 232 mm diam (centre of body) , 104 mm above reflector (centre of body) , 3 mm Cu or Al tube

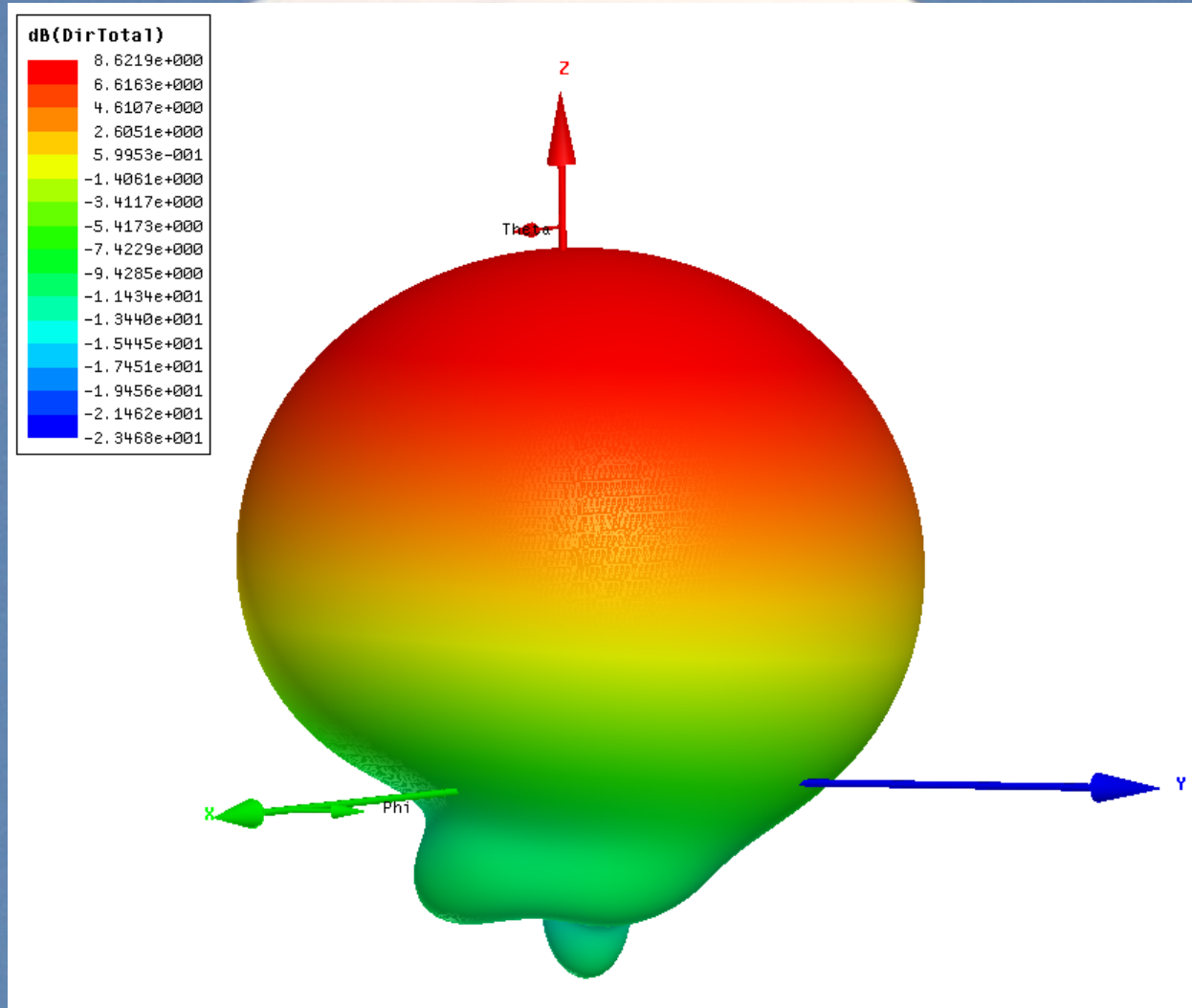
S11, S22, S21 combined



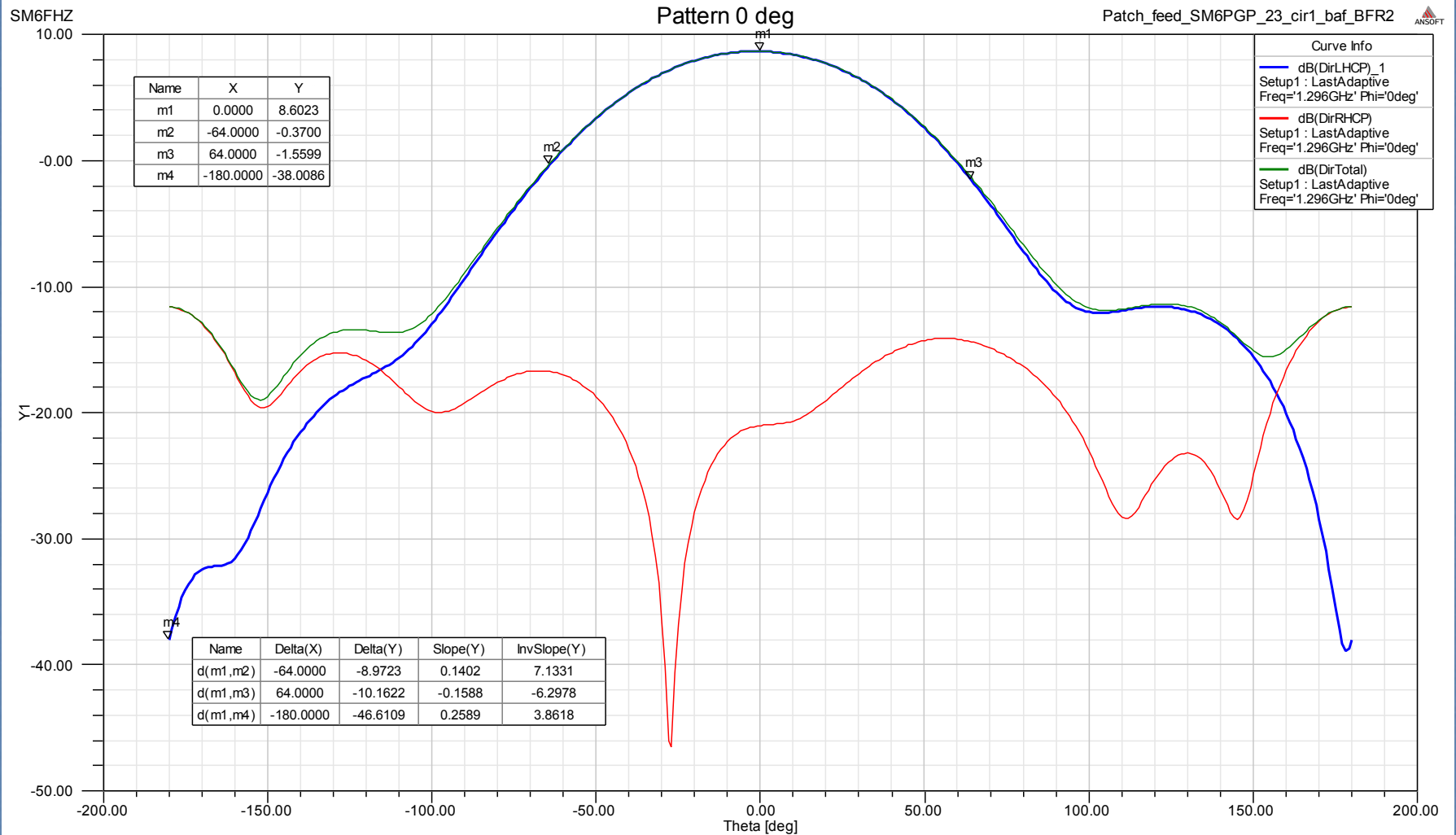
SM6PGP measurements on live model



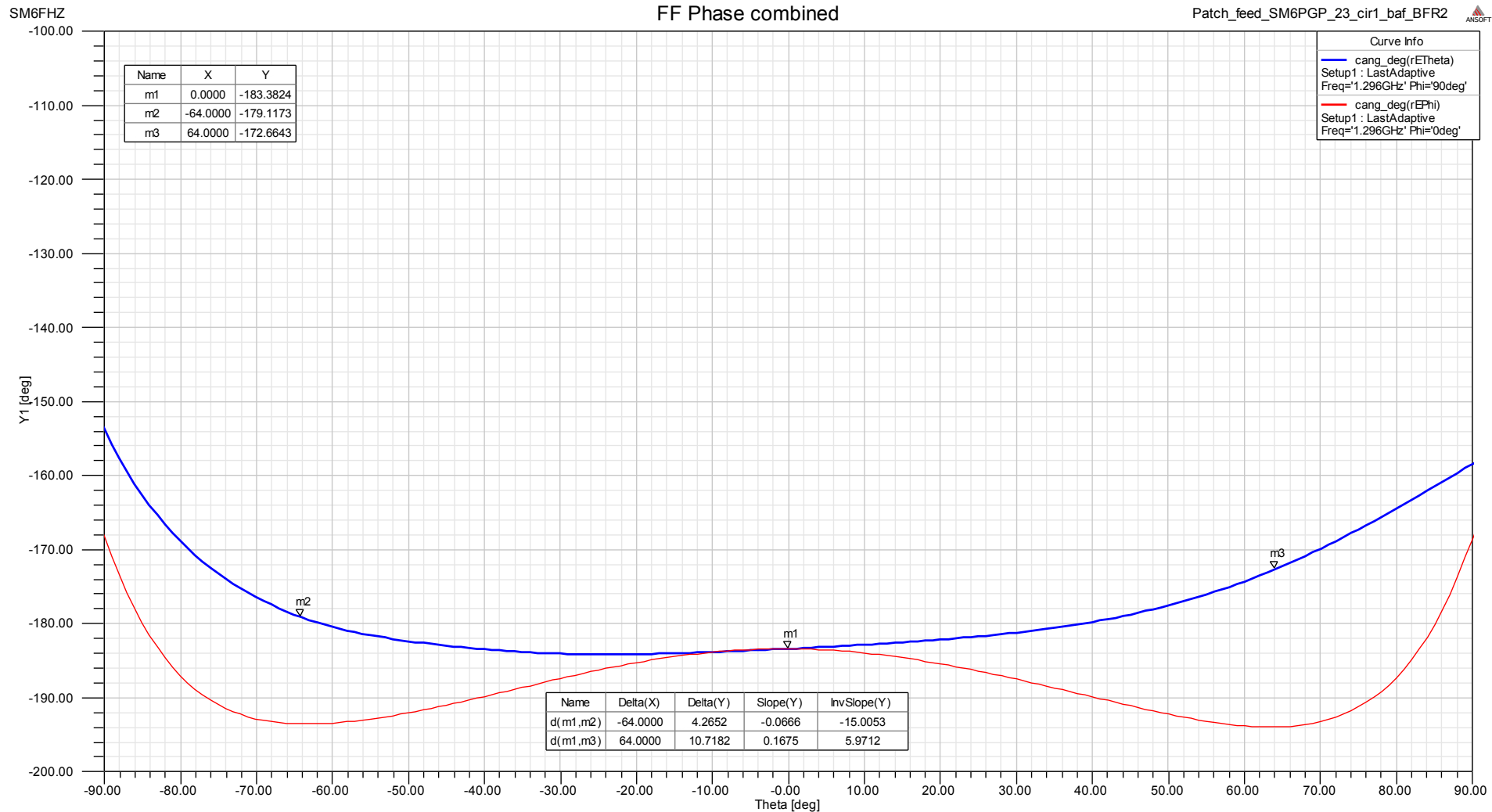
3D Total Power Far Field pattern



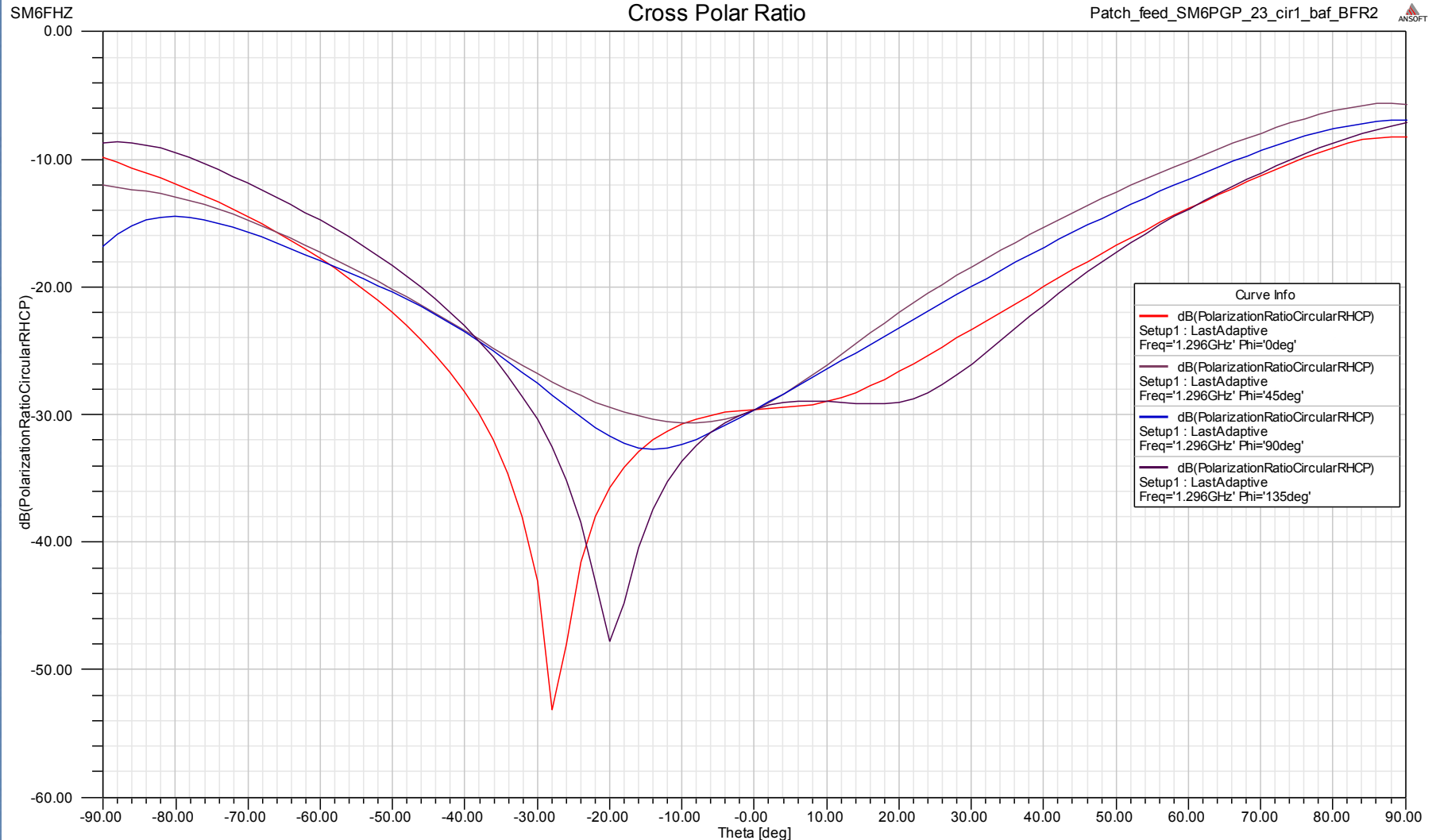
Far Field Pattern 0 deg



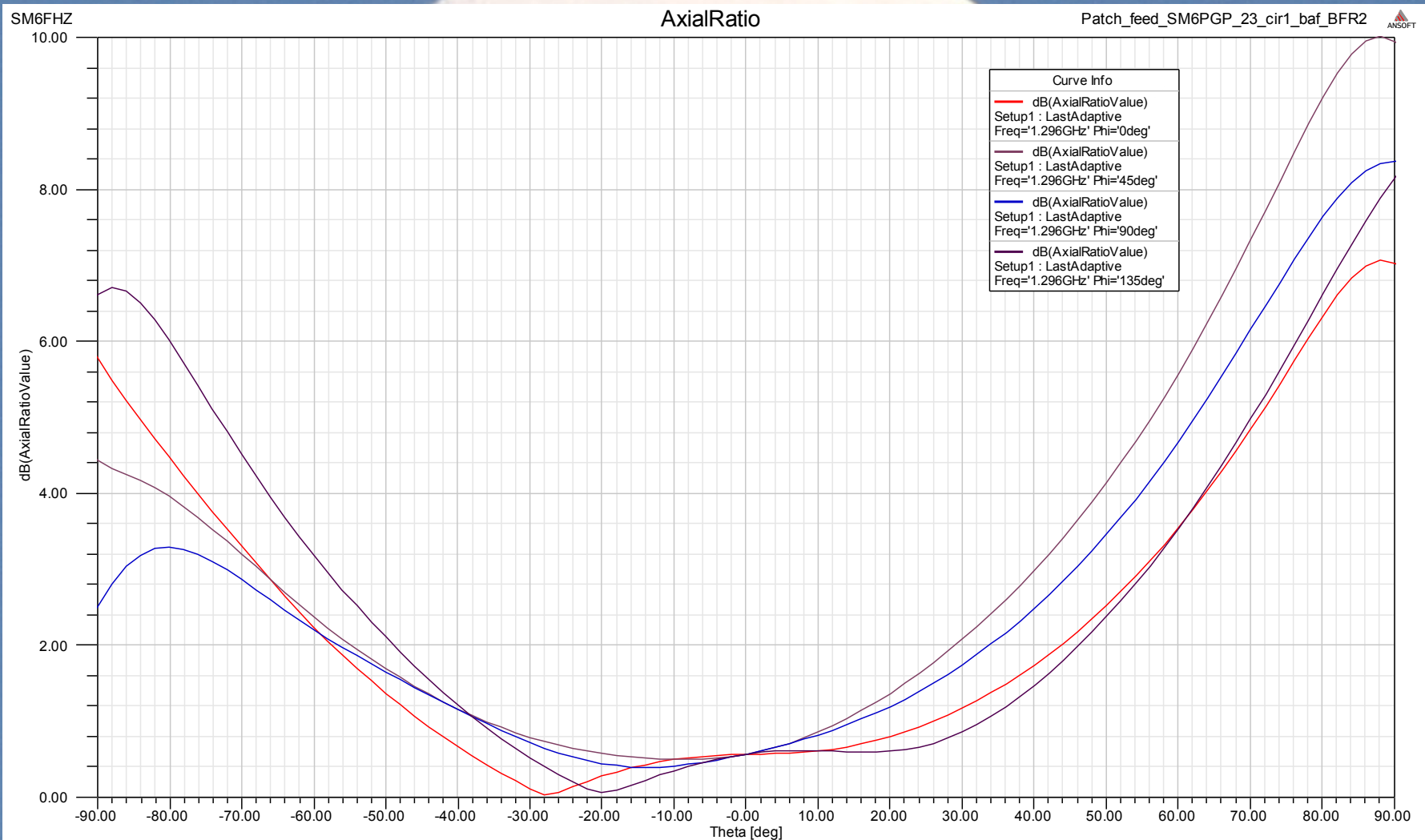
Far Field Phase



Cross Polar Ratio



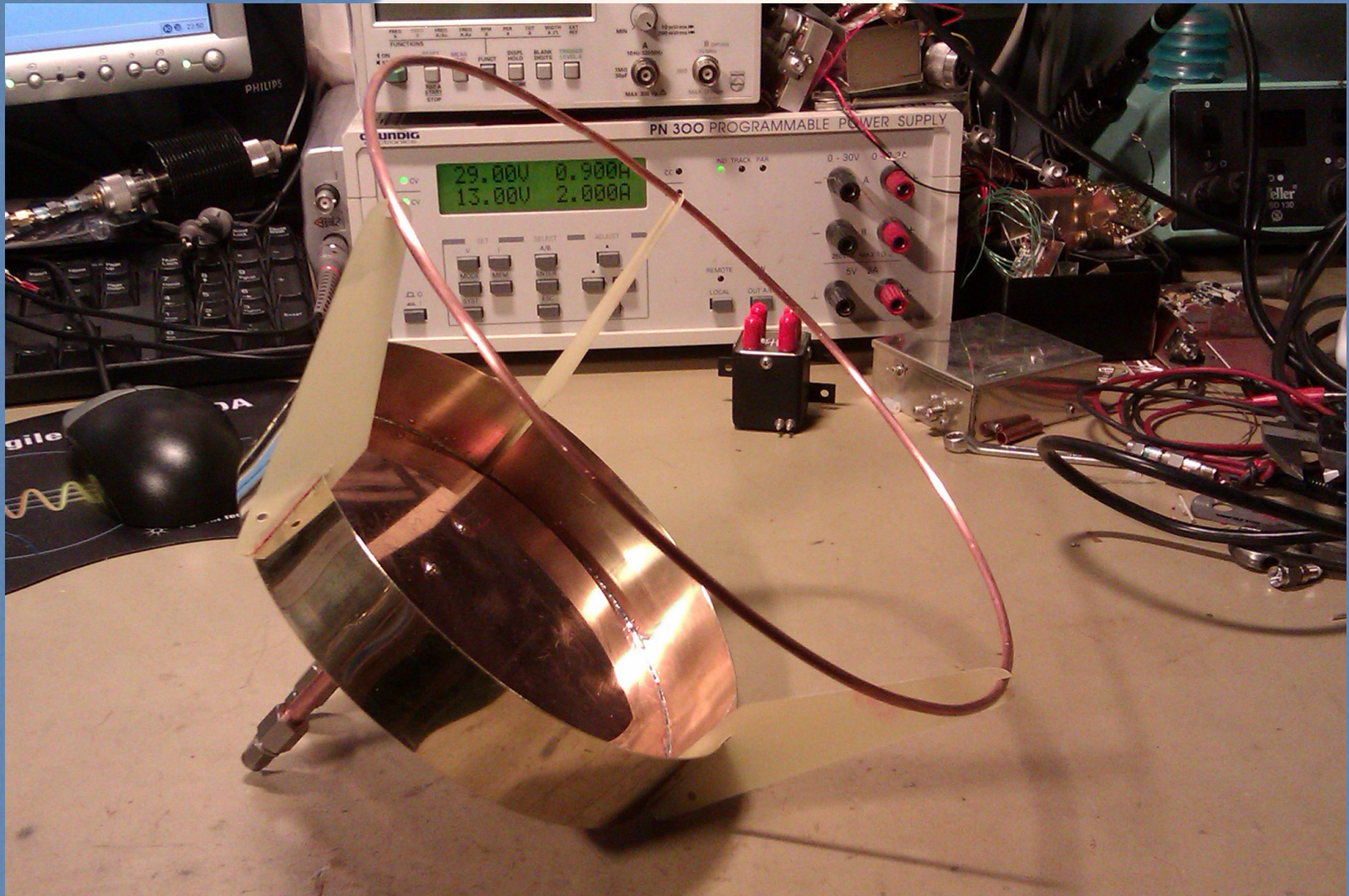
Axial Ratio



Parts cut to fit



Realization

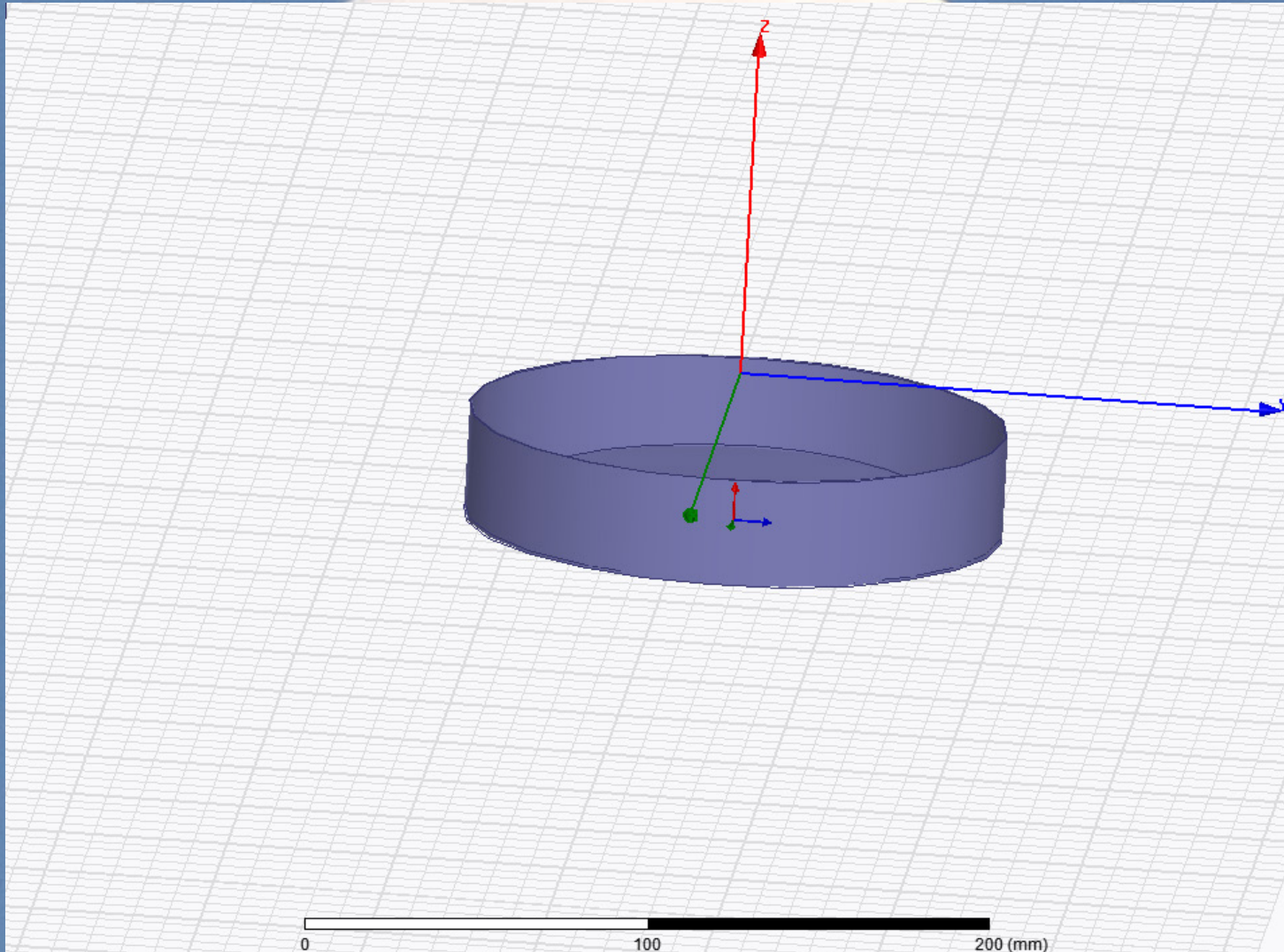


A large, bright, yellowish-white moon is centered in the upper half of the frame against a clear, deep blue sky. The moon's surface is covered in numerous craters of various sizes, and its overall appearance is slightly hazy or soft. The text is overlaid on the moon's surface.

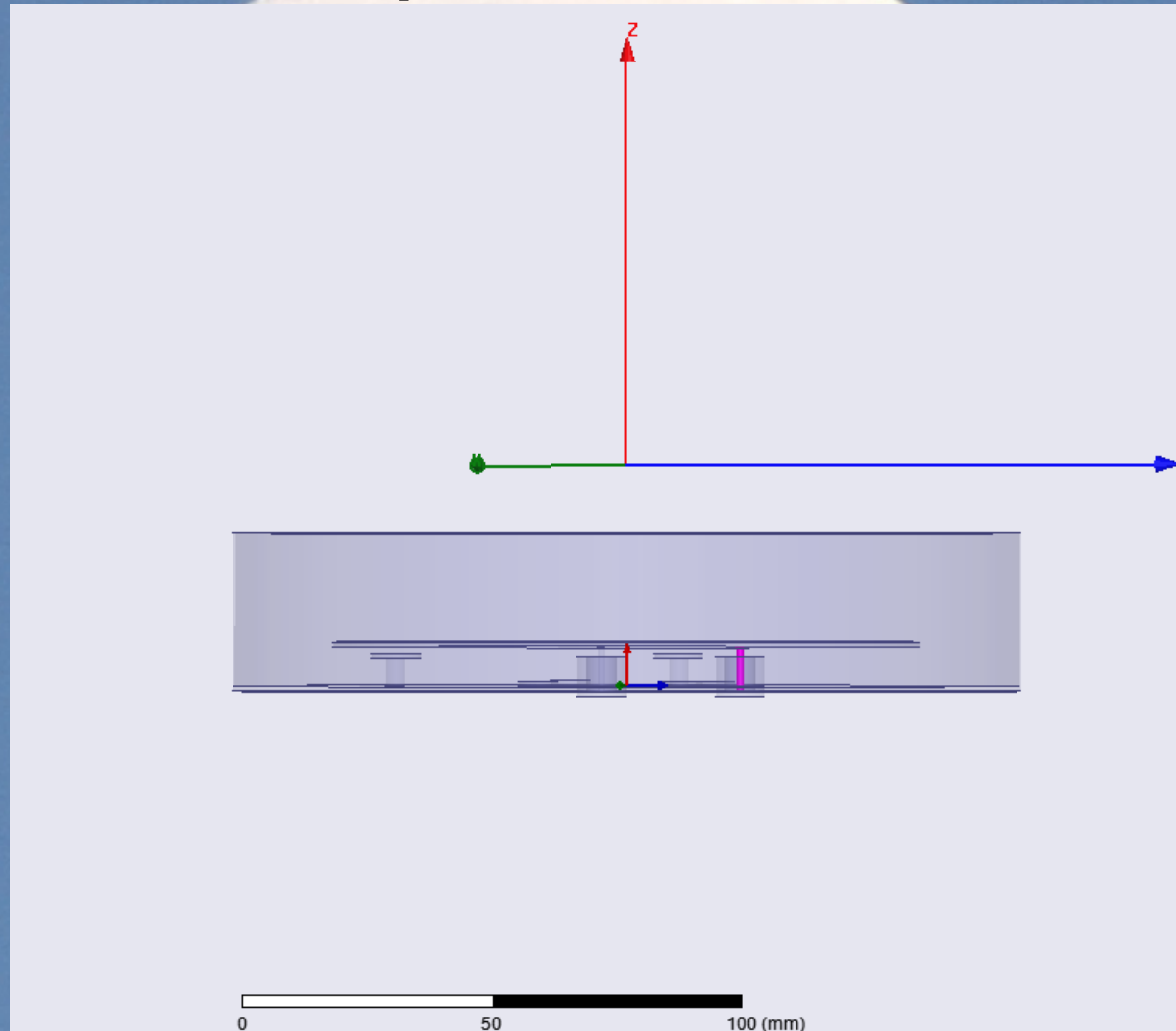
Low f/D variant

For dishes with f/D 's in the 0.30 to 0.38 range

Solid model from the simulation



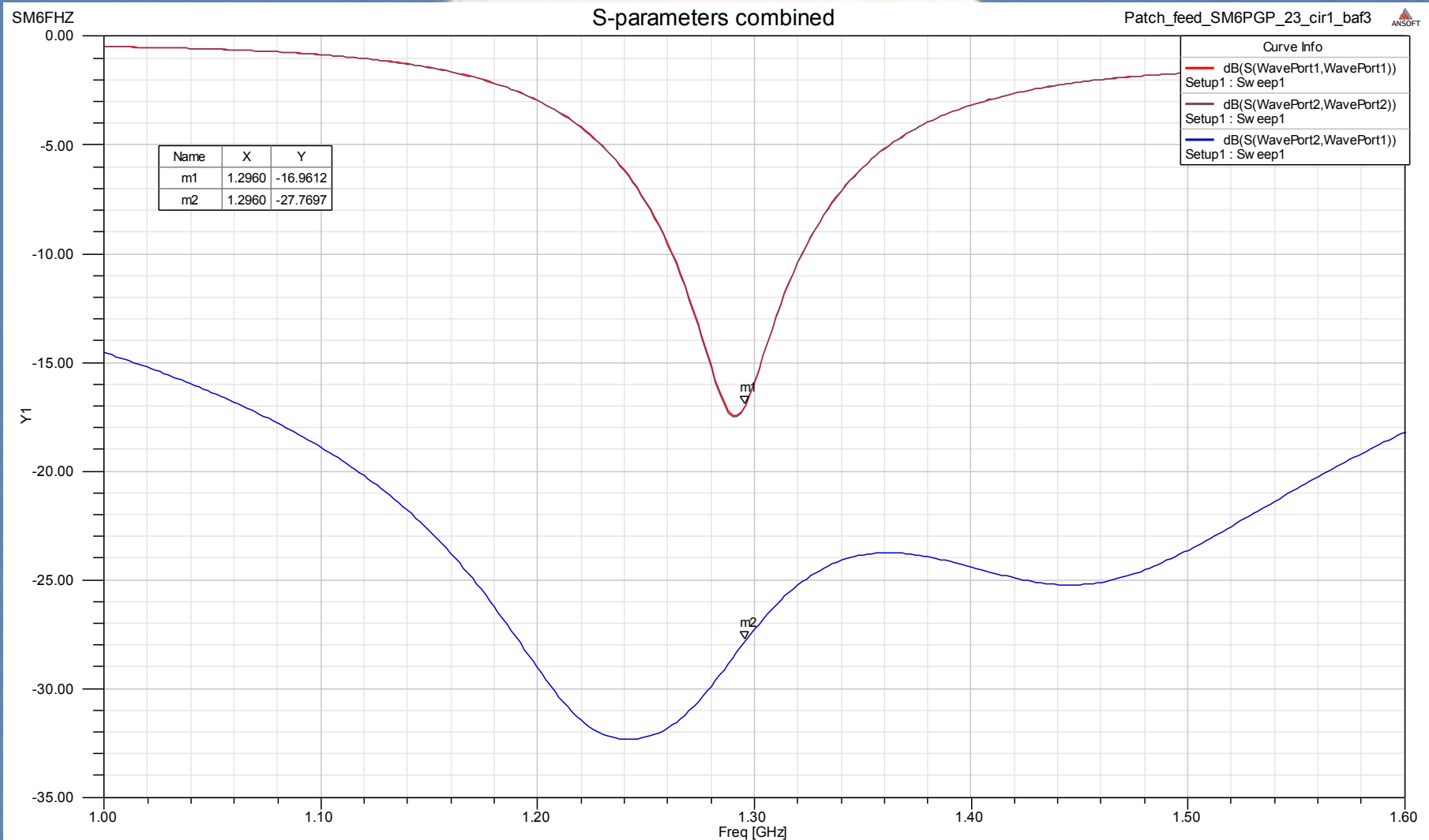
Transparent model



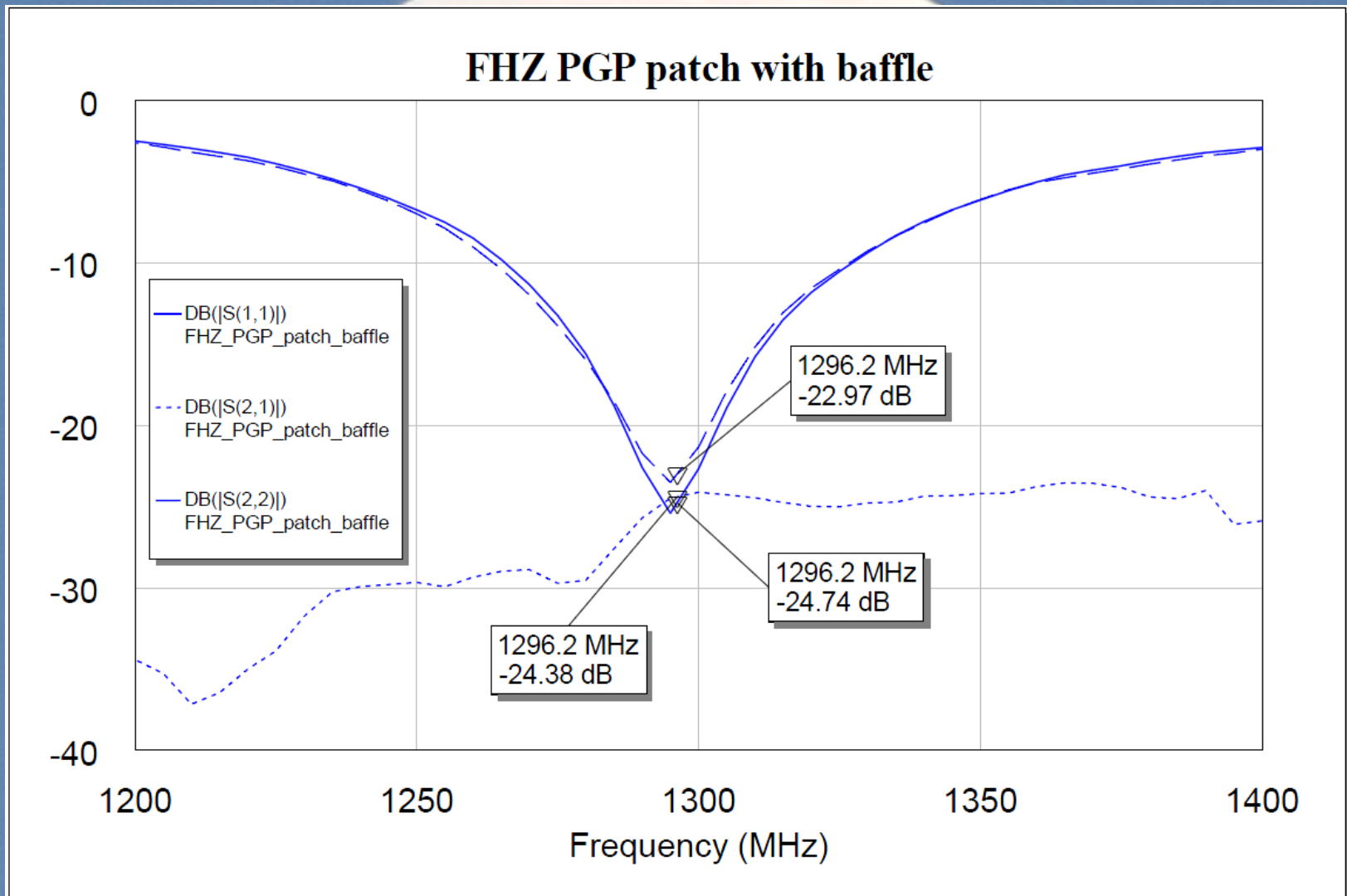
Dimensions without BFR

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- Baffle thickness 0.5 mm
- Patch 119 mm diam, 1 mm thick, 8 mm up from reflector (surface to surface)
- Probe 23.5 mm from center, $\frac{1}{4}$ " coax w 1,6 mm inner conductor. Outer conductor stops 2 mm from patch
- Phase center 36 mm above the reflector

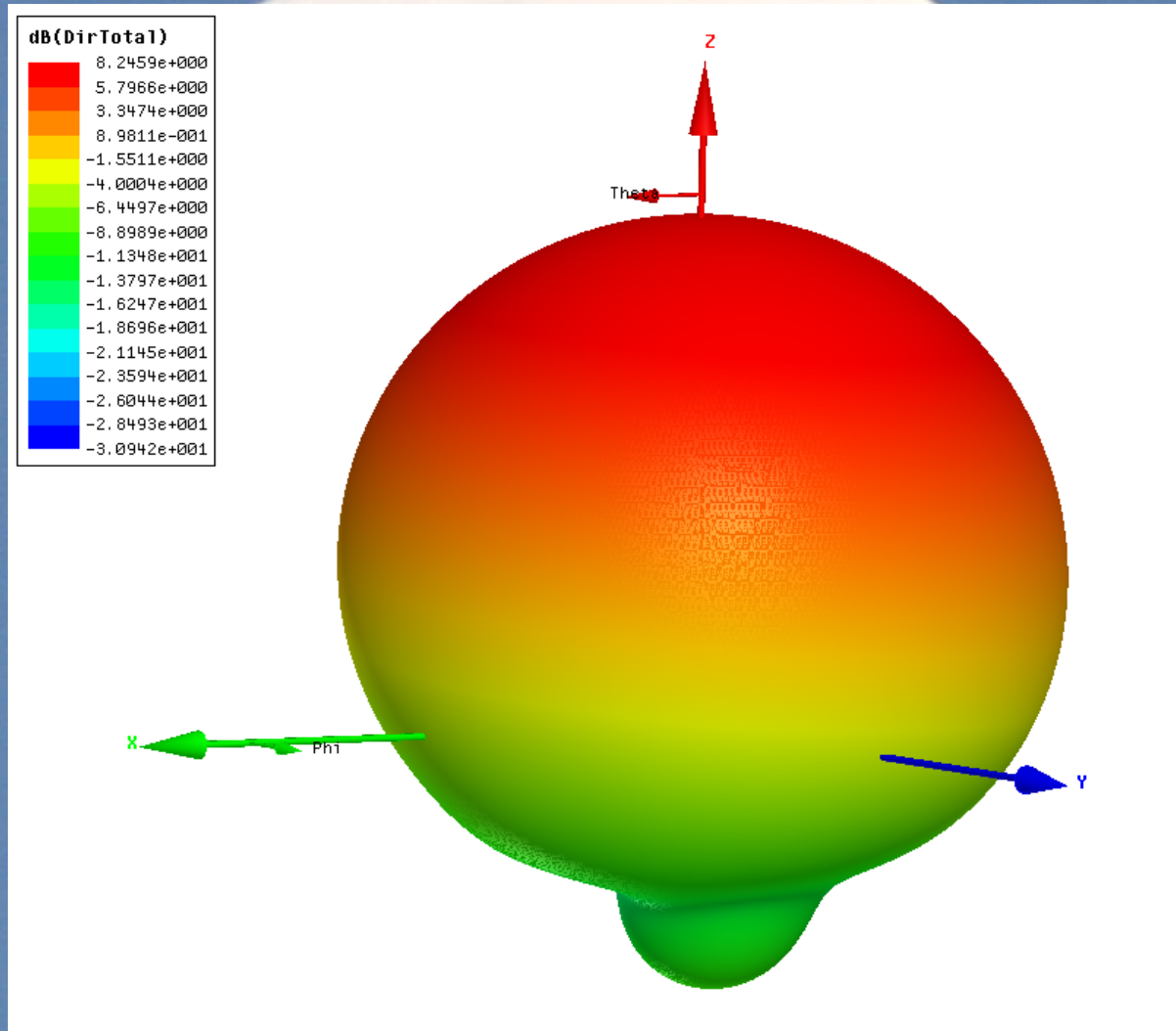
S11, S22, S21 combined



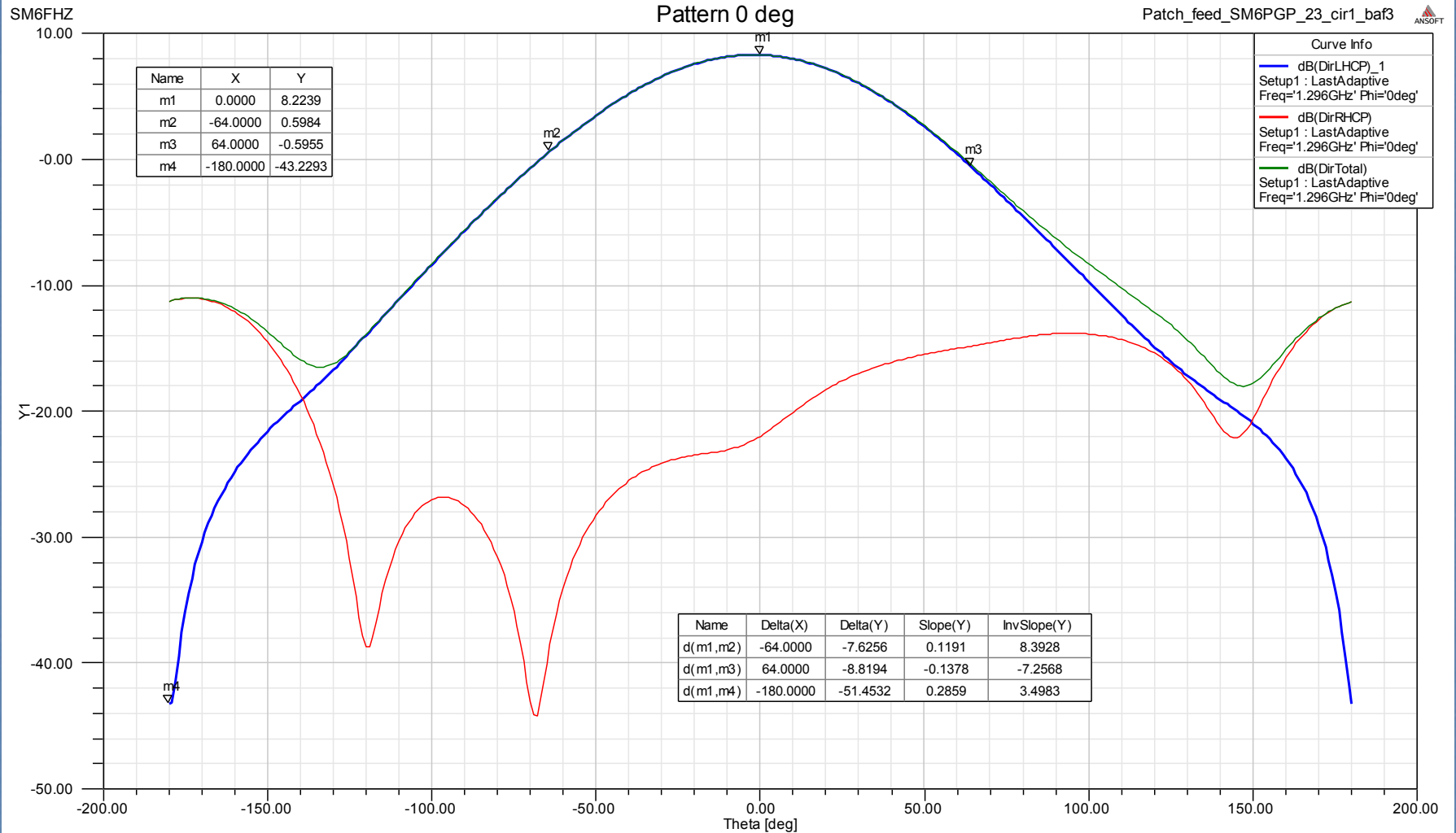
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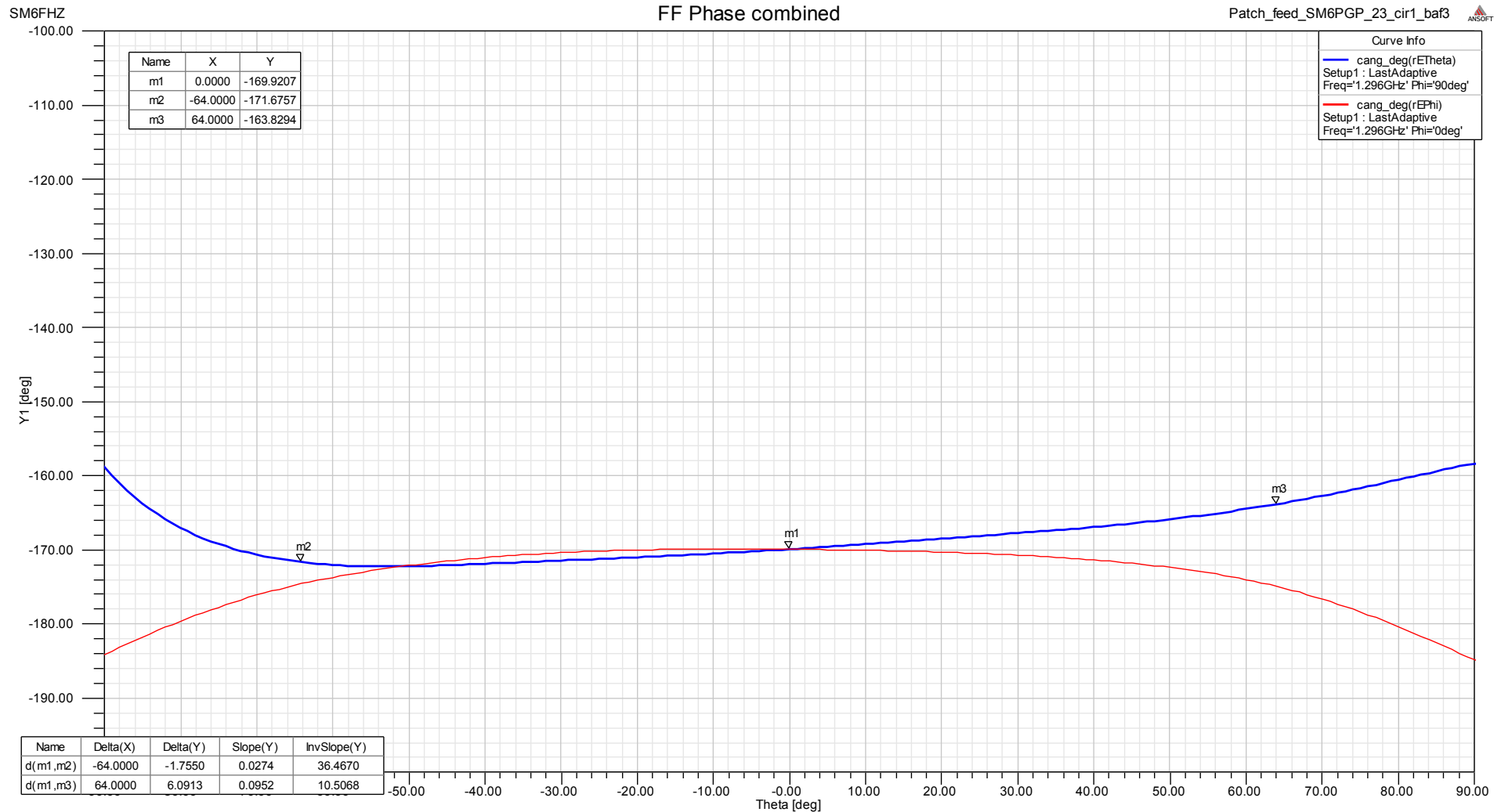
3D Total Power Far Field pattern



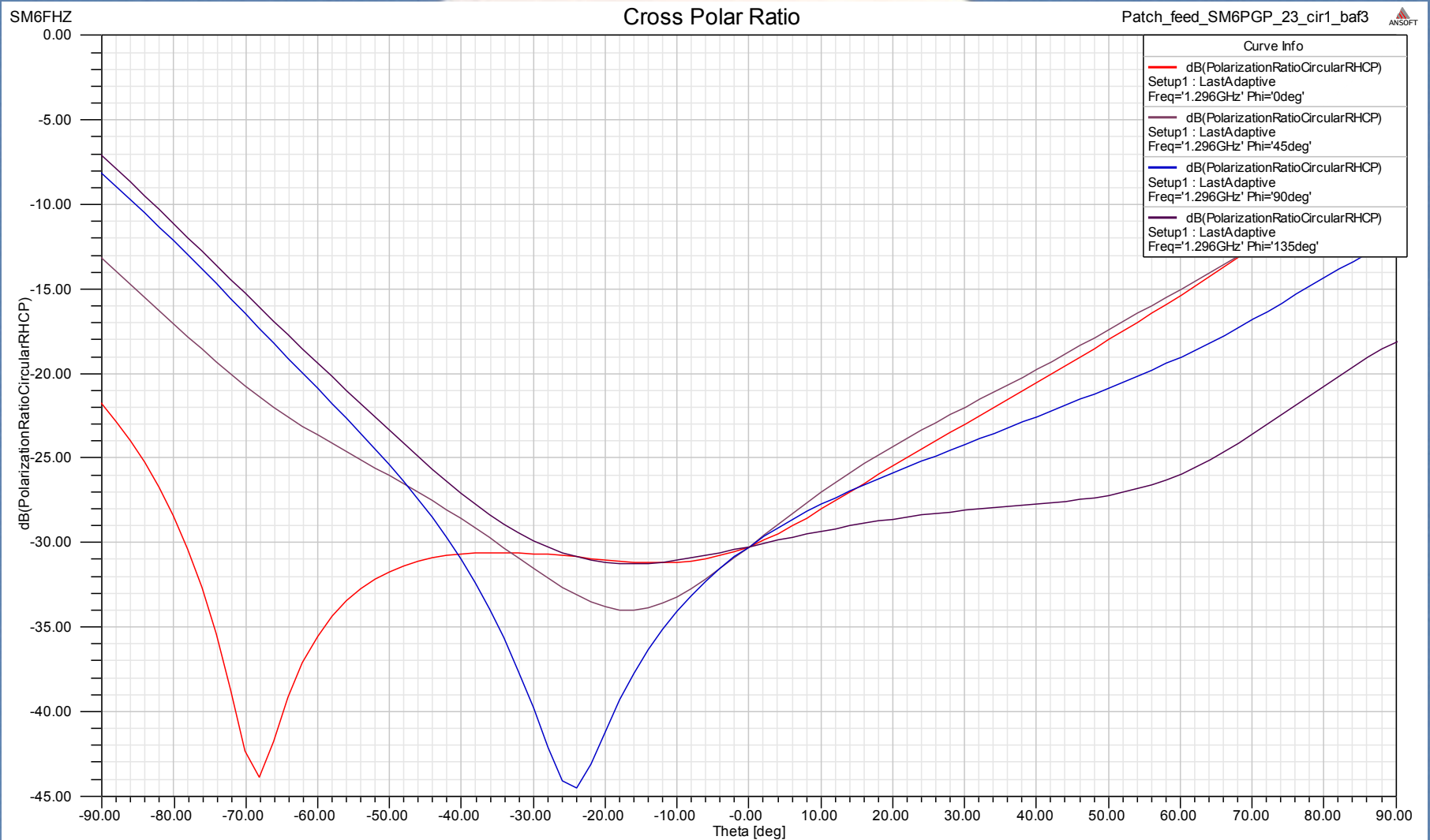
Far Field Pattern 0 deg



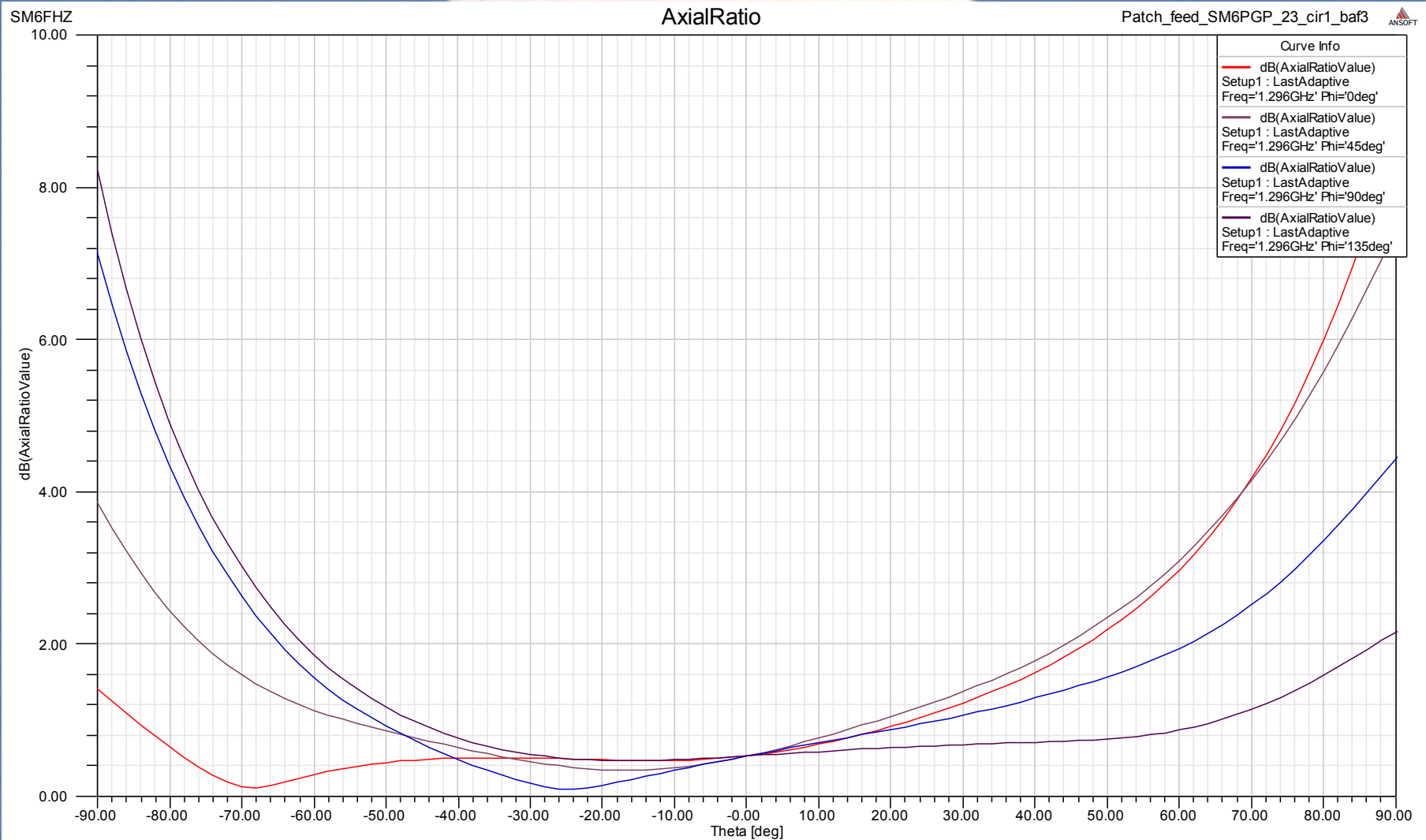
Far Field Phase



Cross Polar Ratio



Axial Ratio



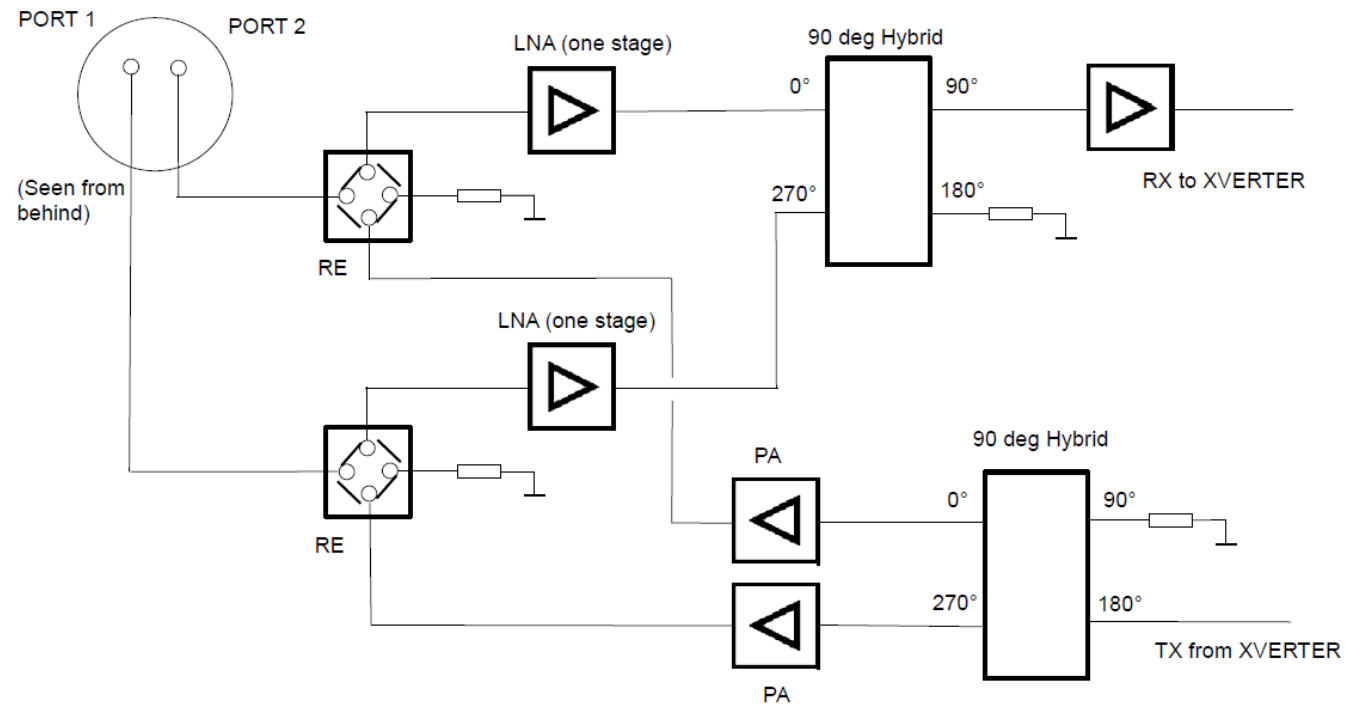
Feeding of the patch

- In order to get circular polarization you need to feed the two probes in quadrature i.e. 90 deg phase difference between the two probes with coherent signals
- This can be accomplished by a branch line coupler (3 dB hybrid) directly to the probes and gives RHCP on one port and LHCP on the other port of the hybrid
- It can also be done with the hybrids on the low power side and phase and amplitude matched PA's and Pre Amps and thus make use of spatial power combination of the two PA's. See SM6PGP solution
- All phase and amplitude errors introduced in the complete feeding network will be a part of the axial ratio of the radiated polarization i.e. keep track of all phase and amplitude variations (e.g. cable lengths etc)

One good way to set it up

SM6PGP 23 CM SETUP FOR CIRCULAR POLARIZATION WITH PATCH FEED

PATCH ANTENNA WITH TWO PORTS 90 DEG OUT OF PHASE



PHASE MATCHED LNA:s & PA:s NEEDED
FINAL HIGH POWER COMBINER **NOT** NEEDED

20130227 - SM6PGP

SM6PGP live model in dish



SM6PGP results

- 1.8 m solid dish (0.4 f/D)
- 7.5 dB solar noise with 0.4 dB LNA
- Heard OK2DL and LX1DB on SSB via the moon.
- Hears the ON0EME beacon weak on CW. Hard to get full call.
- Four JT65 and two CW initials worked with a Tx power of 300 W at the feed so far. More than 10 different CW stations heard during the Dubus 23 cm event this year

Conclusions

- This feed may be an enabler for new stations with smaller dishes to get on the moon on 23 cm
- Easy manufacturing, low cost and minimal dimensions together with good performance makes it an attractive alternative for the smaller dishes
- An excellent choice for feeding your small dish for the upcoming SK6OSO EME event (July 13/14, 2013)
 - Feed it in quadrature for circular polarization. Also an excellent choice for linear polarization in deep dishes, use each port separately for linear polarization (i.e. V&H) for terrestrial communication.
- Again, is not easy to feed a small, flat dish in an efficient way. For the smaller dishes, go for a f/D in the 0.3 to 0.4 range
- For the larger dishes the Kumar type W/G feed is still the king of feeds for the lower f/D 's

Thank you for your attention



Hope to see you via EME on 23 cm from SK6OSO