Formation of the Solar System

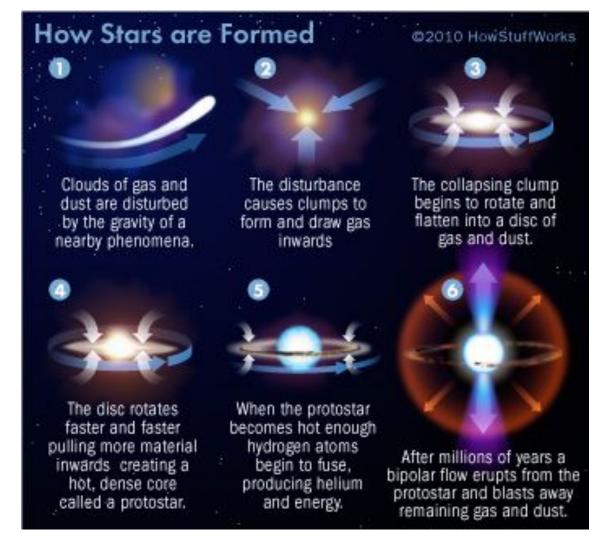
Ms. Larsh

Primeval Nebulas & The Formation of the Solar System



A previously undisturbed **nebula**, a cloud of gas and dust, slowly expands outward in a beautiful display of colors

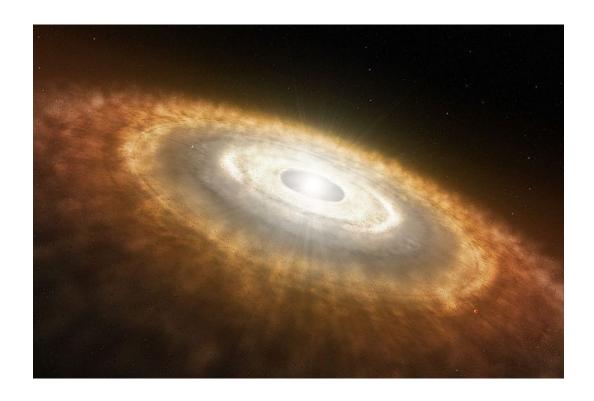
All solar systems and thus, stars originate from these nebulas



Formation of a Star

- A nebula collapses due to gravity
- As material collapses inward the nebula begins to spin
- 3. As the material spins faster the nebula flattens out into a **disk**
- 4. The disk continues to spin faster, with more material clumping in the center to create a **protostar**
- 5. When the pressure & temperature become intense enough, nuclear fusion occurs and a main sequence star is formed

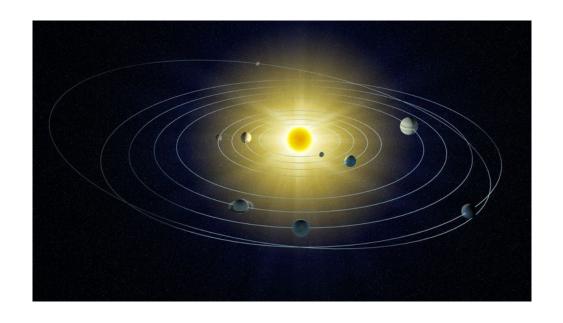
Main Sequence Stars



A star is formed because the pressure and temperature became intense enough for Hydrogen (H) atoms to fuse into Helium (He) atoms

All main sequence stars involve nuclear fusion of $\mathbf{H} \rightarrow \mathbf{He}$

Gravitational Sorting of Materials

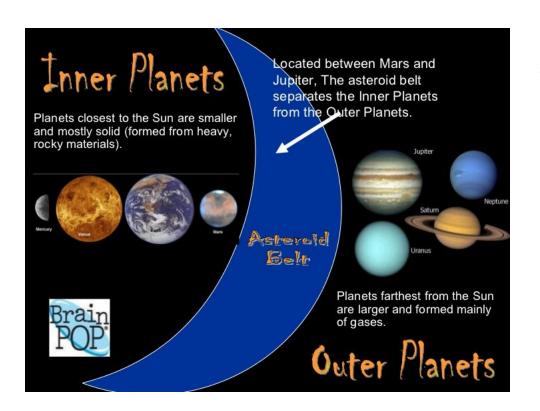


Once nuclear fusion occurs, the **inward collapse** due to gravity is **slowed** by the **outward force** of nuclear fusion

*For every action there is an equal and opposite reaction

Over time, the disk continues to cool and the **gravitational sorting of materials** causes the denser materials to stay closer to the Sun, while the lighter materials are pushed further from the Sun by solar winds

Planet Formation / Location



Denser material near the Sun formed the Terrestrial or Inner Planets: Mercury, Venus, Earth, & Mars

Lighter material away from the Sun formed the Gas Giants or Outer Planets: Jupiter, Saturn, Uranus, & Neptune

Sun: Low-Medium Mass Main Sequence Star

All main sequence stars involve nuclear fusion of hydrogen into helium

 $H \rightarrow He$

The sun is the ultimate source of energy for life on Earth

