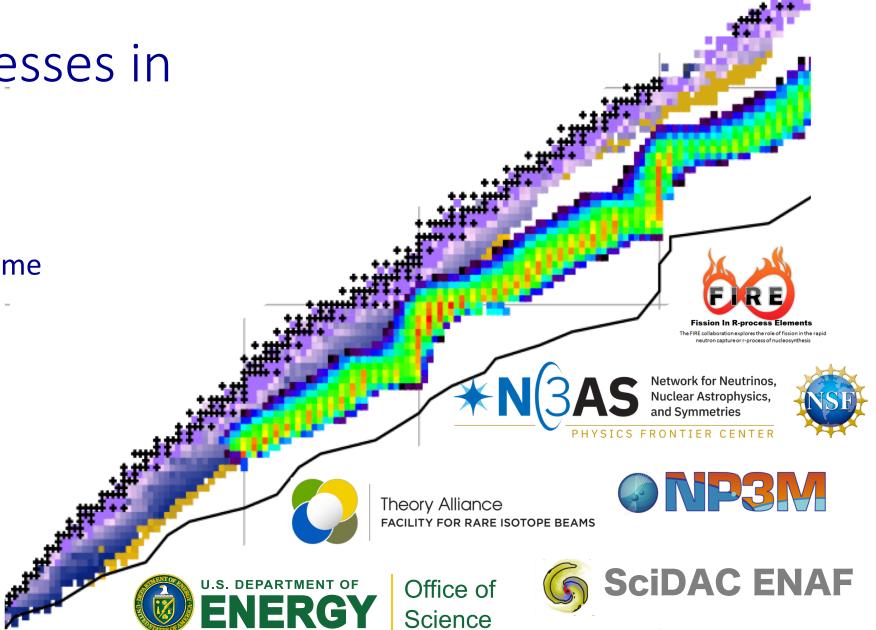
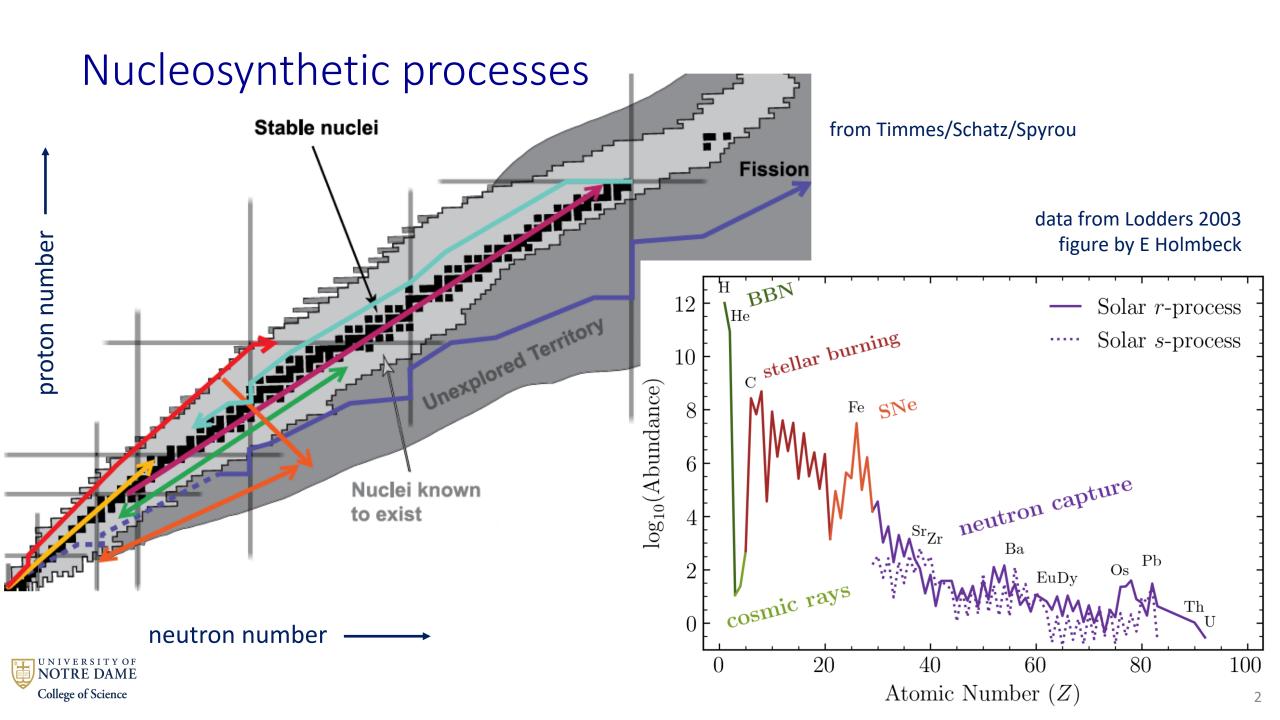
Fission processes in astrophysics

Rebecca Surman University of Notre Dame

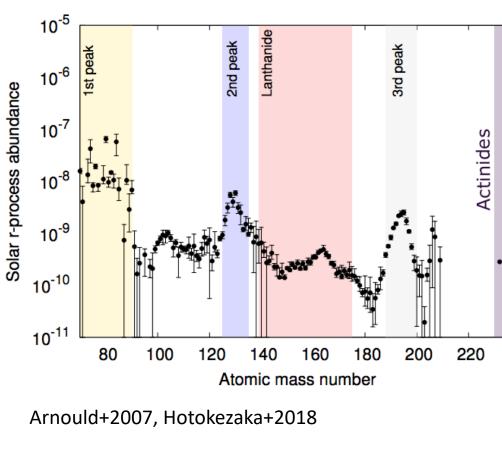
FIESTA 2024 22 Nov 2024



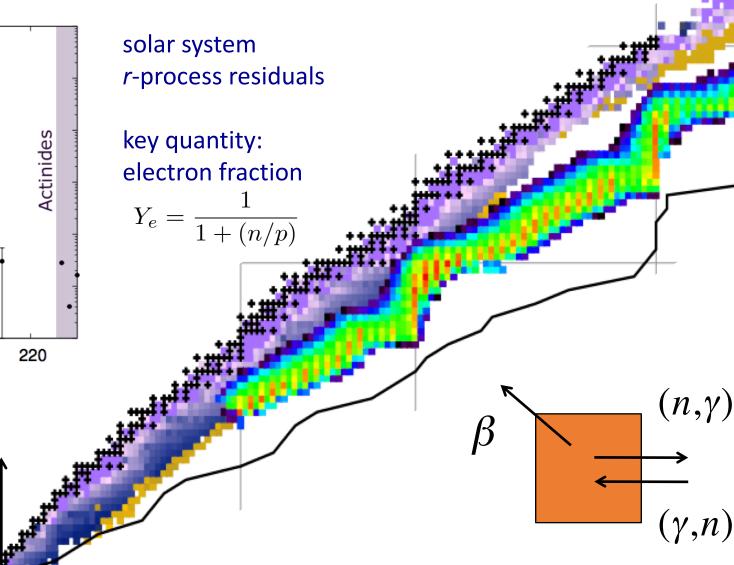




r-process nucleosynthesis



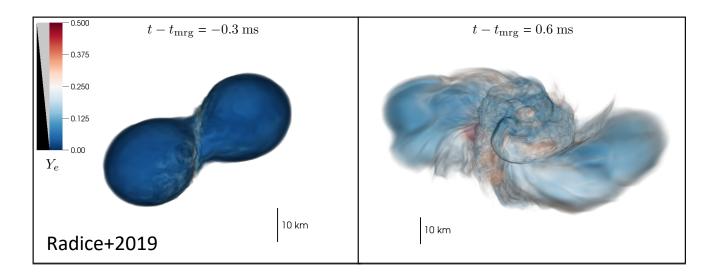
proton number Z



neutron number N

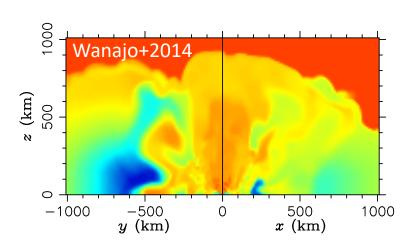


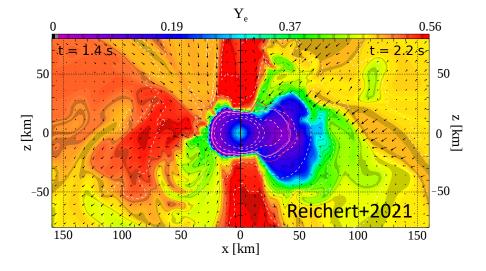
r-process astrophysical sites



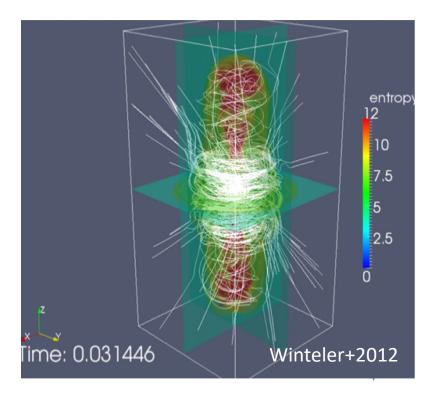
Binary neutron star mergers

(also neutron starblack hole mergers?)

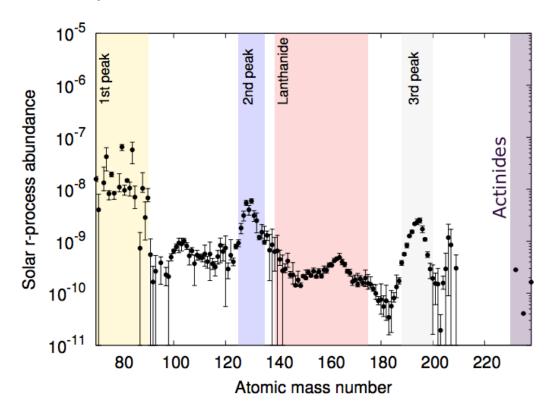




Exotic supernovae (?)



r-process observables: abundance patterns

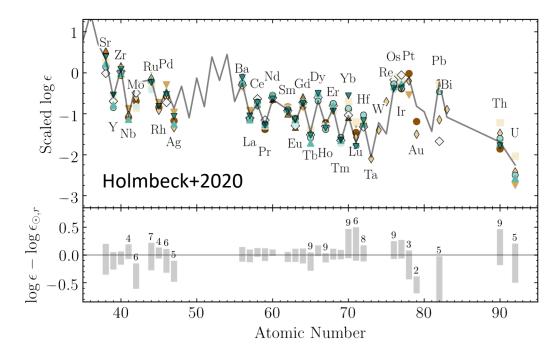


Arnould+2007, Hotokezaka+2018

r-process elements in metal-poor stars

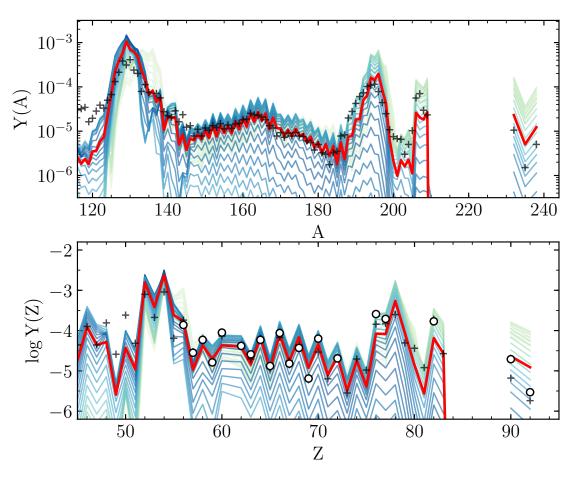
solar system r-process residuals



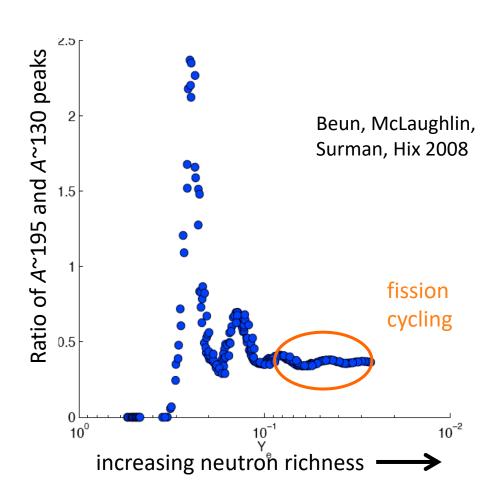




consistent r-process pattern: evidence for fission cycling?



Holmbeck, Sprouse, Mumpower, Vassh, Surman, Beers, Kawano 2019



connecting fission properties to r-process observables

 Can we exploit knowledge of fission and fissioning isotopes to interpret r-process observables?

• What can we learn about the fission properties of neutron-rich actinides from *r*-process observables?

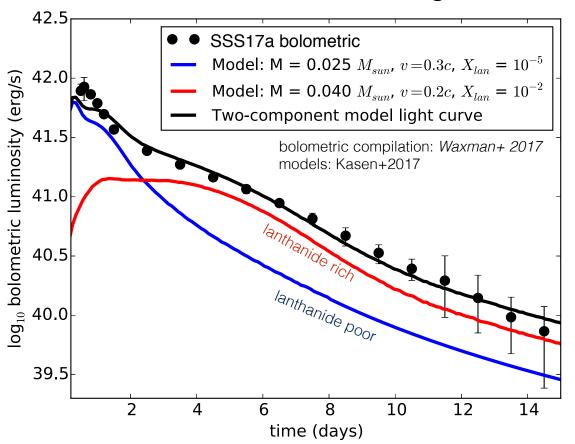
connecting fission properties to *r*-process observables

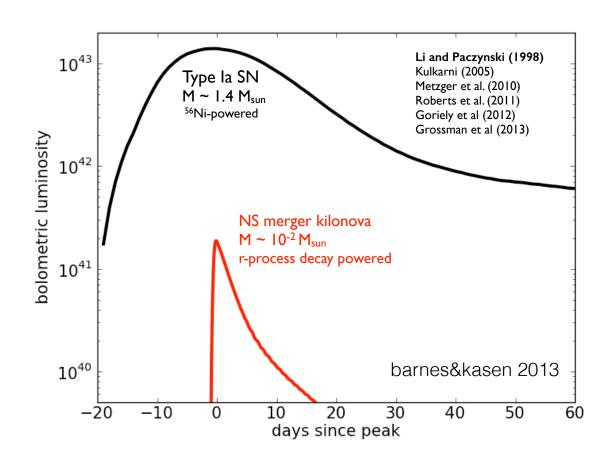
 Can we exploit knowledge of fission and fissioning isotopes to interpret r-process observables?

• What can we learn about the fission properties of neutron-rich actinides from *r*-process observables?

r-process observables: electromagnetic signatures

kilonova SSS17a bolometric light curve

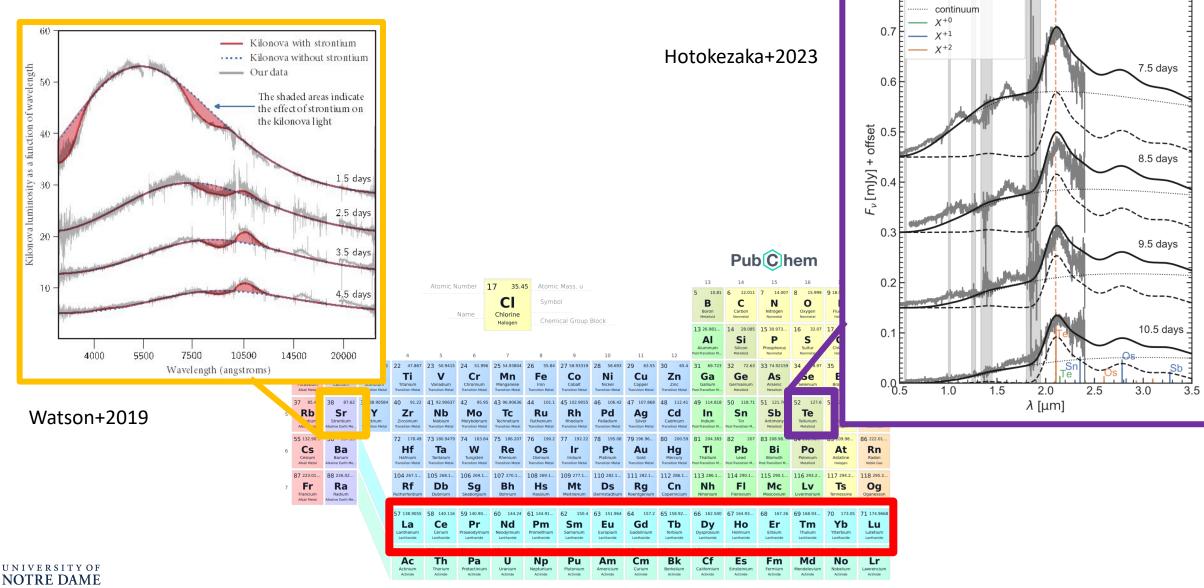






NSM evidence for specific elements

College of Science



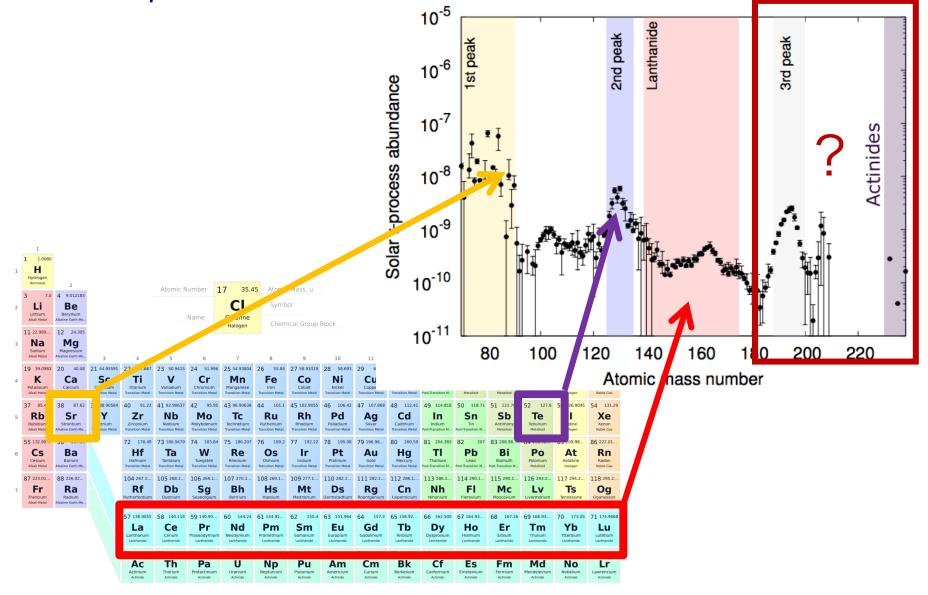
AT2017gfo

M1 line + continuum

[Te III] 2.10μm

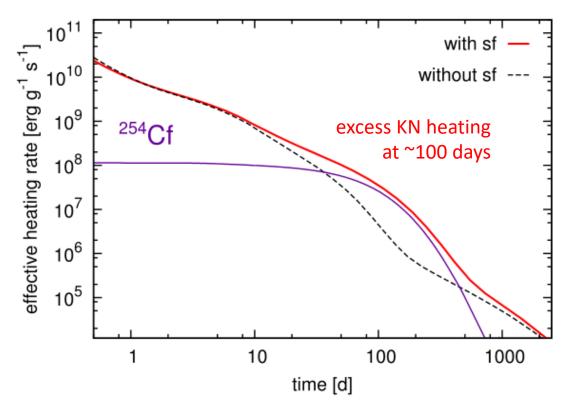
[Te I] 2.10µm

NSM evidence for specific elements





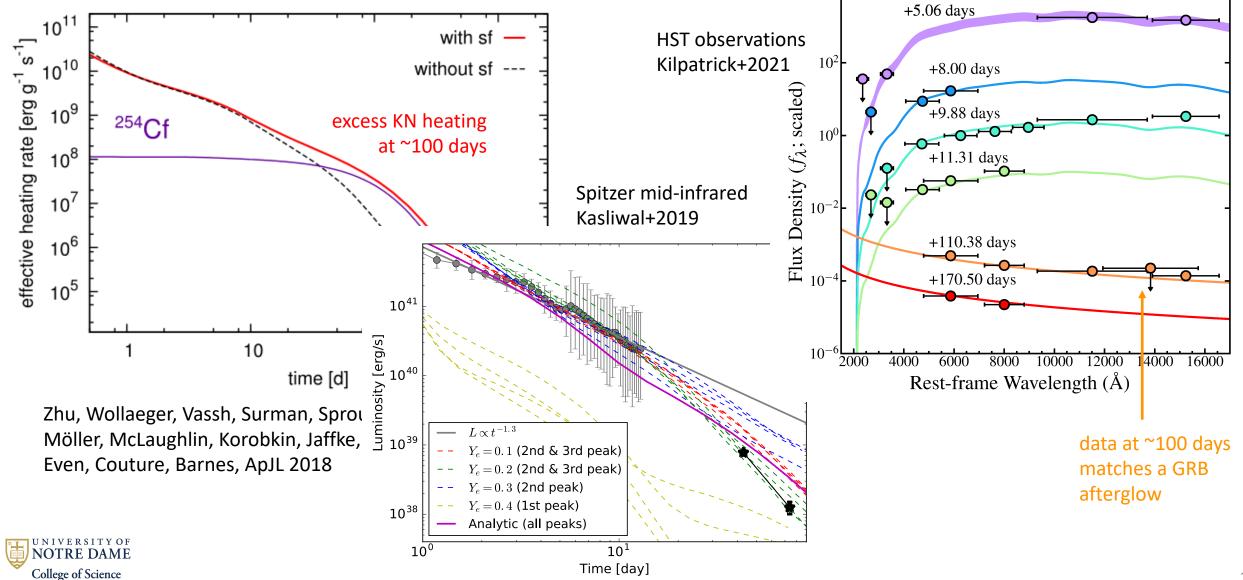
Did the GW170817 merger produce actinides?



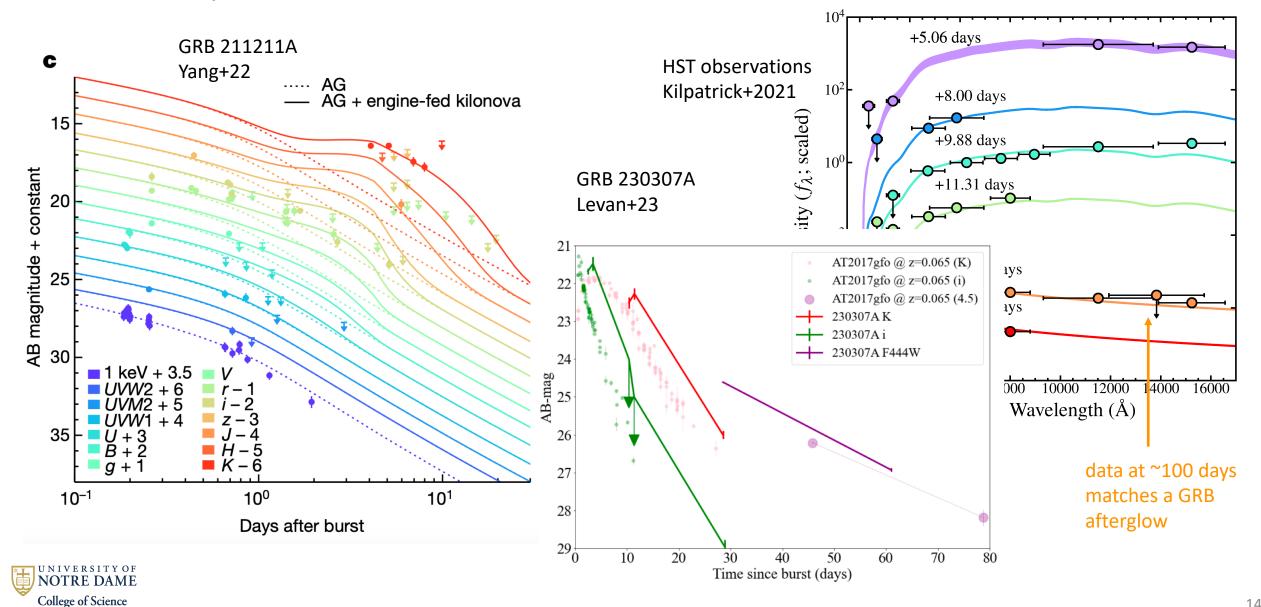
Zhu, Wollaeger, Vassh, Surman, Sprouse, Mumpower, Möller, McLaughlin, Korobkin, Jaffke, Holmbeck, Fryer, Even, Couture, Barnes, ApJL 2018



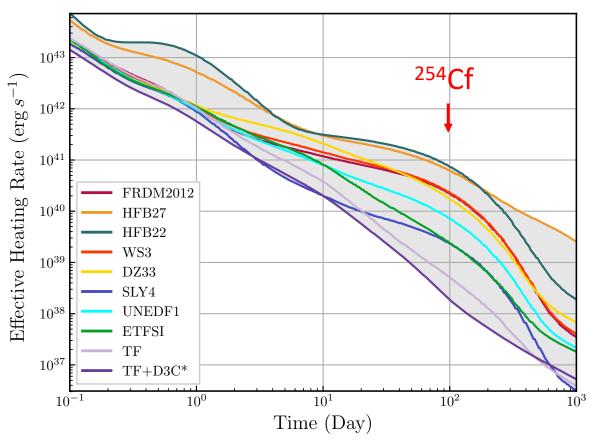
Did the GW170817 merger produce actinides?



Subsequent KNe show similar late time behavior

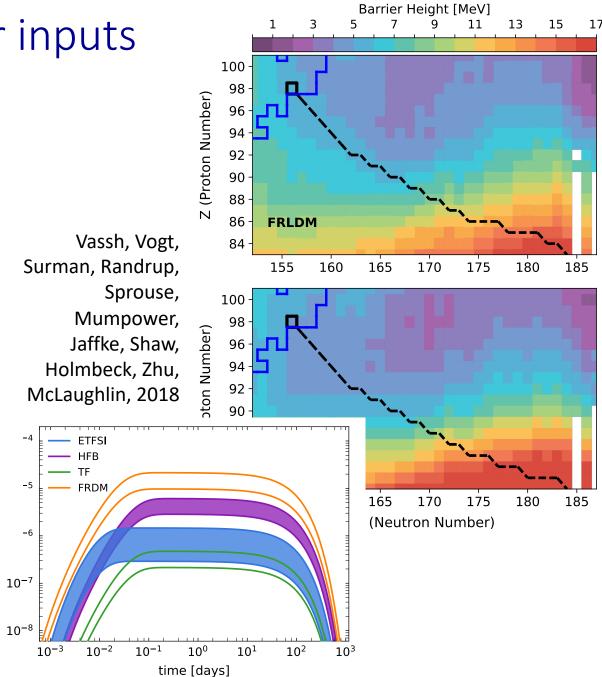


²⁵⁴Cf: dependence on nuclear inputs

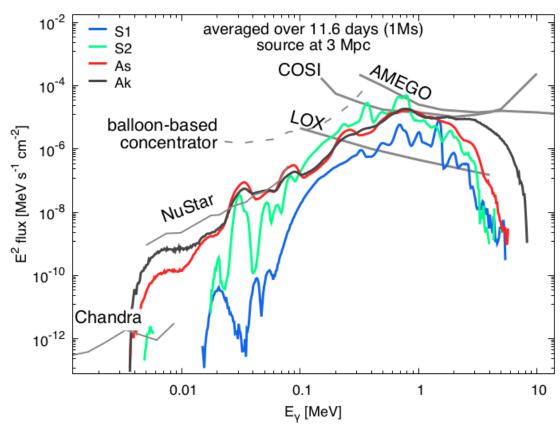


Zhu, Lund, Barnes, Sprouse, Vassh, McLaughlin, Mumpower, Surman 2021





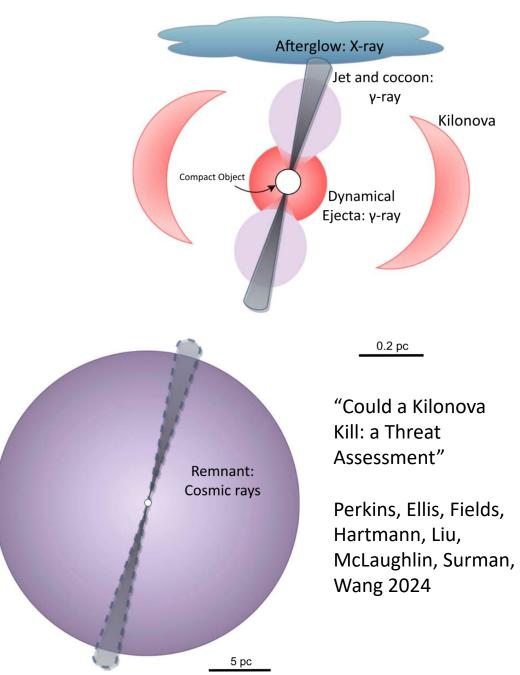
Gamma rays from a nearby event



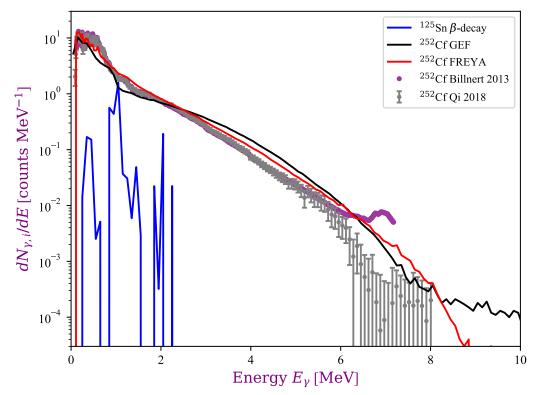
Korobkin, Hungerford, Fryer, Mumpower, Misch, Sprouse, Lippuner, Surman, Couture, Bloser, Shirazi, Evan, Vestrand, Miller 2020

also Hotokezaka+2016; Li 2019; Wu+2019; Ruiz-Lapuente, Korobkin 2020

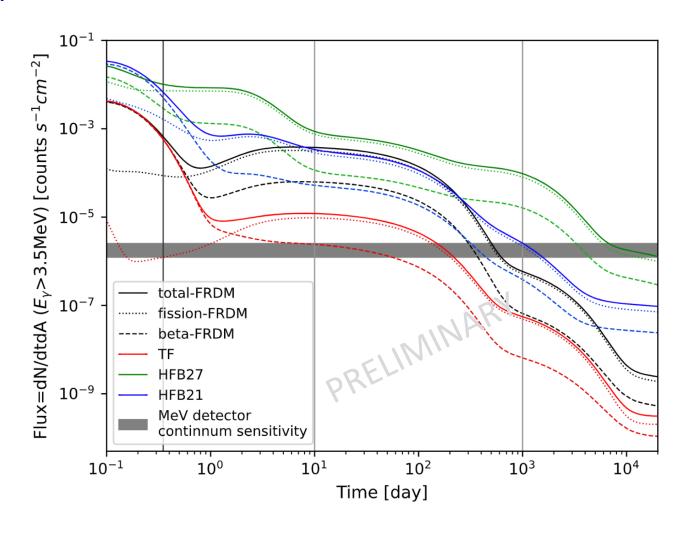




Gamma rays from fission



Wang, Vassh, Sprouse, Mumpower, Vogt, Randrup, Surman, ApJL 2020



Wang, Vassh+ in preparation 2024

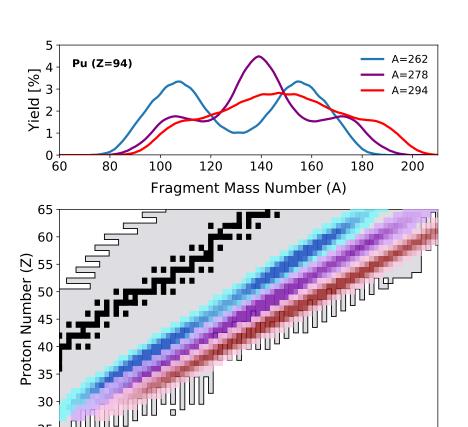


connecting fission properties to *r*-process observables

• Can we exploit knowledge of fission and fissioning isotopes to interpret *r*-process observables?

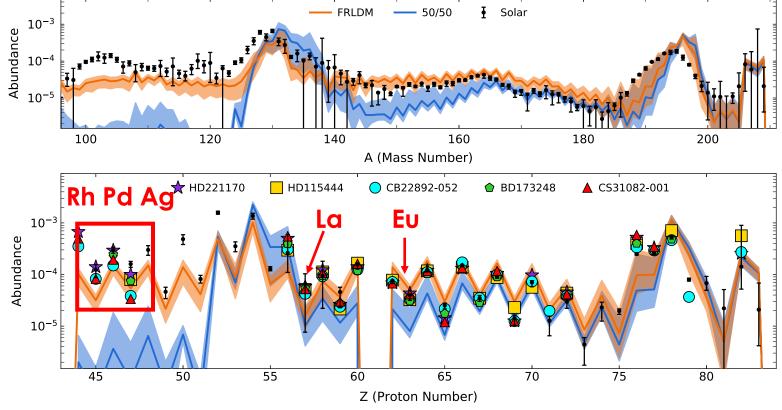
 What can we learn about the fission properties of neutron-rich actinides from r-process observables?

Fission yield signatures



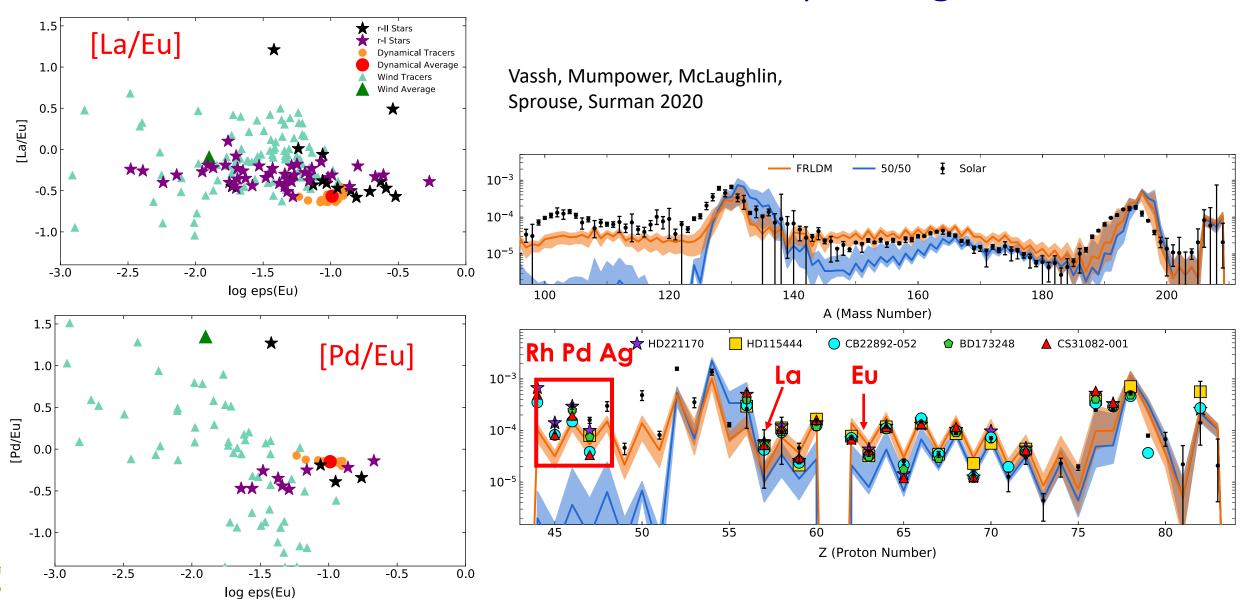
Neutron Number (N)

Vassh, Mumpower, McLaughlin, Sprouse, Surman 2020

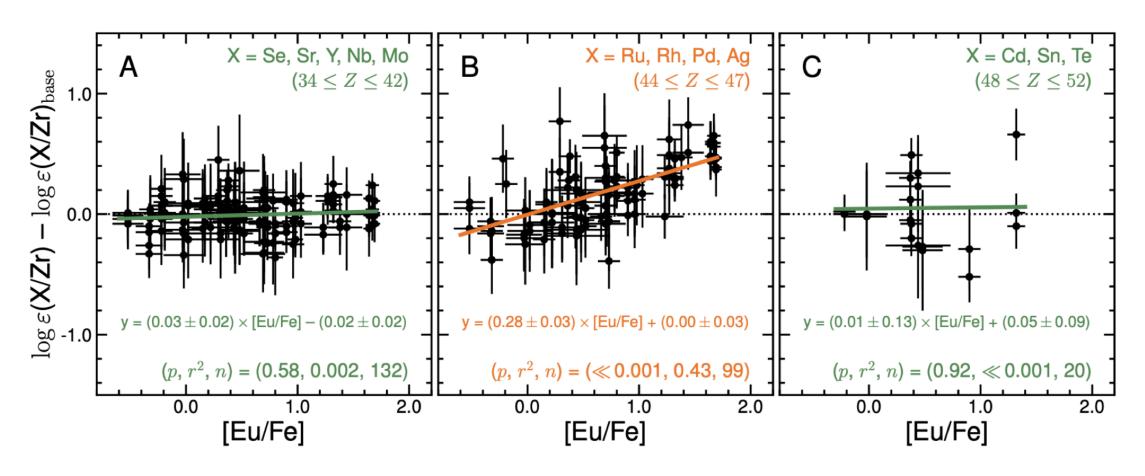




Fission yield signatures



Fission yield signatures



Roederer, Vassh, Holmbeck, Mumpower, Surman, Cowan, Beers, Ezzeddine, Frebel, Hansen, Placco, Sakari, *Science* 2023

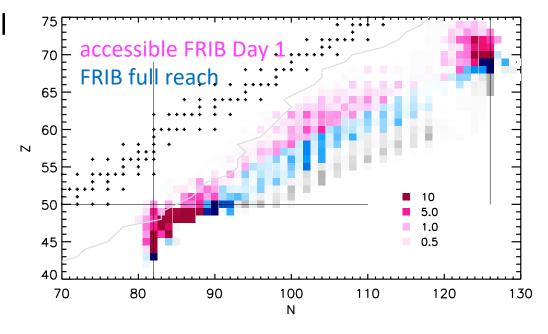


summary

The origin of the heaviest elements in the *r*-process of nucleosynthesis has been one of the greatest mysteries in nuclear astrophysics for decades.

Despite considerable progress in the past several years, including the first direct detection of an *r*-process event, the *r*-process site(s) has not been definitively determined.

An understanding of fission is crucial for the interpretation of *r*-process observables such as abundance patterns and light curves.



Additionally, as other nuclear physics uncertainties are reduced, details of fission properties of neutron-rich actinides may be extractable from *r*-process data.

Mumpower, Surman, McLaughlin, Aprahamian, JPPNP 2016



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Rodney Orford

Guy Savard

students postdocs

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Raph Hix

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